ASTR7741 (Fall, 2007)

1. Spherical Free-Fall Problem

Determine exactly how long it takes a (massless) particle to "free-fall" from rest onto a point mass of mass M if the particle is initially at a distance r_0 from the point mass. To answer this question, you need to integrate (twice) the following equation of motion (eq. 3.11, p. 120 in Vol. II of Padmanabhan):

$$\frac{d^2r}{dt^2} = -\frac{GM}{r^2}\,.\tag{1}$$

More specifically, you should set $v_r = 0$ and $r = r_0$ at time t = 0, then determine a precise, analytical expression for the time τ_{ff} at which $r \to 0$.

- Hint #1: Multiply through by 2v (on the left-hand-side) and by 2dr/dt (on the right-hand-side), and integrate once to determine v^2 as a function of r, M, and r_0 .
- Hint #2: To set up the second integral, replace v by dr/dt, and remember that the velocity must be negative.

1 SPHERICAL FREE-FALL PROBLEM

ANSWER: Padmanabhan actually provides the answer to this problem, beginning with Eq. (3.11), p. 120 of Vol. II. First note that Eq. (1), above, can be rewritten in the form,

$$\frac{dv_r}{dt} = -\frac{GM}{r^2}.$$
(2)

Therefore,

$$2v_r \frac{dv_r}{dt} = -\frac{2GM}{r^2} \frac{dr}{dt}$$
(3)

$$\Rightarrow d(v_r^2) = 2GM \ d(r^{-1}) \tag{4}$$

$$\Rightarrow \text{At any } r < r_0, \qquad v_r^2 = 2GM\left(\frac{1}{r} - \frac{1}{r_0}\right) \tag{5}$$

Since $v_r = dr/dt$, this also means,

$$\frac{dr}{dt} = -\left[2GM\left(\frac{1}{r} - \frac{1}{r_0}\right)\right]^{1/2} \tag{6}$$

$$\Rightarrow \left[2GM\left(\frac{1}{r} - \frac{1}{r_0}\right)\right]^{-1/2} dr = -dt \tag{7}$$

To integrate this, it is convenient to adopt the substitution,

$$\cos^2 \zeta \equiv \frac{r}{r_0}; \text{ hence } dr = -r_0(2\cos\zeta\sin\zeta) \, d\zeta, \qquad (8)$$

in which case $\zeta = 0$ when t = 0 $(r = r_0)$ and when $r \to 0$, $\zeta \to \pi/2$. With this substitution, the equation to be integrated takes the form,

$$(2GM)^{1/2} dt = \left[\frac{1}{r_0 \cos^2 \zeta} - \frac{1}{r_0}\right]^{-1/2} r_0(2\cos\zeta\sin\zeta) d\zeta \tag{9}$$

$$\Rightarrow \left[\frac{GM}{2r_0^3}\right]^{1/2} dt = \left[\frac{1}{\cos^2 \zeta} - 1\right]^{-1/2} (\cos \zeta \sin \zeta) d\zeta \tag{10}$$

$$= \cos^2 \zeta \, d\zeta \tag{11}$$

$$\Rightarrow \left[\frac{GM}{2r_0^3}\right]^{1/2} t = \int_0^\zeta \cos^2 \zeta \, d\zeta \tag{12}$$

$$= \left[\frac{1}{2}\zeta + \frac{1}{4}\sin(2\zeta)\right]_{0}^{\zeta} \tag{13}$$

$$= \frac{1}{2}\zeta + \frac{1}{4}\sin(2\zeta) \tag{14}$$

Hence, when $\zeta = \pi/2$,

$$t = \tau_{\rm ff} = \frac{\pi}{4} \left[\frac{2r_0^3}{GM} \right]^{1/2} = \left[\frac{\pi^2 r_0^3}{8GM} \right]^{1/2}$$
(15)

But also note that $\bar{\rho}_0 \equiv 3M/(4\pi r_0^3)$, hence, $(r_0^3/M) = 3/(4\pi \bar{\rho}_0)$, which means,

$$\tau_{\rm ff} = \left[\frac{3\pi}{32 \ G\bar{\rho}_0}\right]^{1/2}.$$
 (16)