

Algebra-based Physics II Review for Chap. 24-27

Cha. 24. Electromagnetic waves

Basic characters:

$$\begin{array}{llll}
 E(t) = E_0 \sin \omega t & E_{rms} = \frac{E_0}{\sqrt{2}} & E = cB & c = f \cdot \lambda \\
 B(t) = B_0 \sin \omega t & B_{rms} = \frac{B_0}{\sqrt{2}} & \omega = 2\pi f & c = \frac{1}{\sqrt{\epsilon_0 \mu_0}}
 \end{array}$$

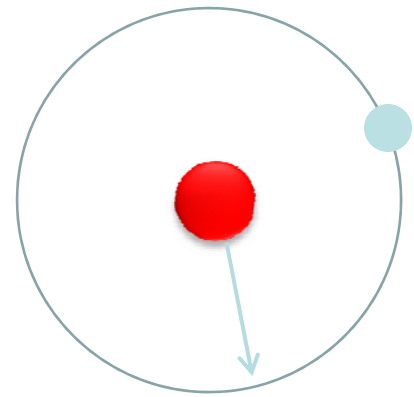
Energy density:

$$\begin{array}{l}
 u = \frac{\epsilon_0}{2} E^2 + \frac{1}{2\mu_0} B^2 = \epsilon_0 E^2 = \frac{1}{\mu_0} B^2 \\
 \bar{u} = \frac{\epsilon_0}{2} E_{rms}^2 + \frac{1}{2\mu_0} B_{rms}^2 = \epsilon_0 E_{rms}^2 = \frac{1}{\mu_0} B_{rms}^2
 \end{array}$$

Intensity:

$$\begin{array}{l}
 S = \frac{P}{A} = \frac{E}{t \cdot A} = cu \quad \bar{S} = c\bar{u} \\
 S = cu = \frac{c\epsilon_0}{2} E^2 + \frac{c}{2\mu_0} B^2 = c\epsilon_0 E^2 = \frac{c}{\mu_0} B^2
 \end{array}$$

Example: What fraction of the power radiated by the sun is intercepted by the planet Mercury? The radius of Mercury is $2.44 \times 10^6 \text{m}$, and its mean distance from the sun is $5.79 \times 10^{10} \text{m}$. Assume that the sun radiates uniformly in all direction.



Doppler Effect:

$f_o = f_s \left(1 \pm \frac{v_{rel}}{c} \right)$	Doppler shift + if v_{rel} is towards - if v_{rel} is away
$\frac{\lambda_s}{\lambda_o} = 1 \pm \frac{v_{rel}}{c}$	

Polarization:

Malus' Law: $S = S_0 \cos^2 \theta$	Calculate the intensity change
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Cha. 25 & 26: Geometric Optics

Index of refraction:

$$n = \frac{c}{v} = \frac{\lambda_{vac}}{\lambda}$$

Snell's Law: $n_1 \sin \theta_1 = n_2 \sin \theta_2$

Apparent depth: $d' = d \left(\frac{n_2}{n_1} \right)$ $\begin{cases} n_1 \text{ contains incident ray} \\ n_2 \text{ contains refracted ray} \end{cases}$

Total Internal reflection (critical angle): $\sin \theta_c = \frac{n_2}{n_1}$

Brewster's Angle: $\tan \theta_B = \frac{n_2}{n_1}$ $\begin{cases} n_1 \text{ contains incident ray} \\ n_2 \text{ contains refracted ray} \end{cases}$

Mirrors and Lenses

Law of reflection: $\theta_r = \theta_i$

Mirror/lens equation: $\frac{1}{d_o} + \frac{1}{d_i} = \frac{1}{f}$

Magnification: $m = \frac{h_i}{h_o} = -\frac{d_i}{d_o}$

focal length f $\begin{cases} 0.5 * R \text{ for spherical mirror} \\ - \text{ for convex mirror, diverging lens} \\ + \text{ for concave mirror, converging lens} \end{cases}$

Cha. 27: Interference & Diffraction

Linear superposition principle $\left\{ \begin{array}{l} \text{In phase: Constructive} \\ \text{Out of the phase: Destructive} \end{array} \right.$

Double slit $m = 0, 1, 2, 3, \dots$

$$d \sin \theta = \begin{cases} m\lambda & \text{bright fringes (constructive)} \\ (m + \frac{1}{2})\lambda & \text{dark fringes (destructive)} \end{cases}$$

Single slit $m = 1, 2, 3, \dots$

$$\sin \theta = \frac{m\lambda}{W} \quad \text{dark fringe of order } m$$

Thin film interference

$$2t + \text{shift} ? = m\lambda \quad \text{Constructive interference}$$

$$2t + \text{shift} ? = m\lambda + \frac{1}{2}\lambda \quad \text{Destructive interference}$$

Cha. 27: Interference & Diffraction

Circular aperture - Resolving power - Rayleigh Criterion

$$\sin \theta \approx \theta > 1.22 \frac{\lambda}{D}$$

Diffraction grating Principal maxima (bright fringes)

$$\sin \theta = m \frac{\lambda}{d}$$
