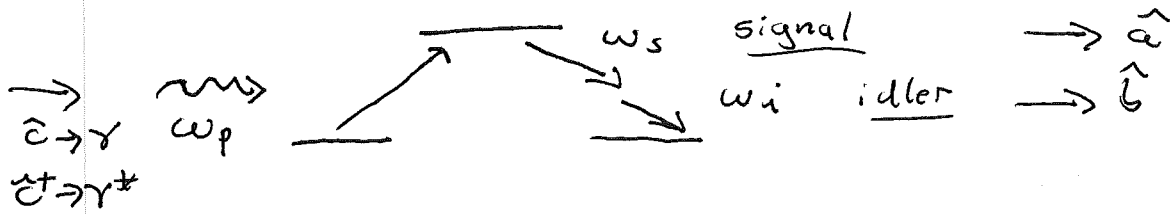


7.7 Two Mode Squeezed Vacuum

In χ_2 parametric downconversion: the general case



$$|1\rangle_{\omega_p} \rightarrow |1\rangle_{\omega_s} |1\rangle_{\omega_i}$$

We do not have to have $\omega_s = \omega_i$
polarization and directions can be different

$$|1\rangle_{\omega_p, \vec{k}_p, s_p} \rightarrow |1\rangle_{\omega_s, \vec{k}_s, s_s} |1\rangle_{\omega_i, \vec{k}_i, s_i}$$

\uparrow \uparrow \uparrow \uparrow \uparrow
 frequency wave number polarization

They are all connected

Freq.	$\hbar \omega_p = \hbar \omega_s + \hbar \omega_i$	cons. Eng
Wave number	$\hbar \vec{k}_p = \hbar \vec{k}_s + \hbar \vec{k}_i$	cons. mom.
polarization	$\hbar \vec{e}_p = \hbar \vec{e}_s + \hbar \vec{e}_i$	cons. z mom.

so in general signal and idler in different modes labeled by $\omega_i \neq \omega_s$; $\vec{k}_i \neq \vec{k}_s$; $\vec{e}_i \neq \vec{e}_s$

To model this we associate \hat{a} with one mode and \hat{b} with the other $\Rightarrow [\hat{a}, \hat{b}^\dagger] = 0$

The hamiltonian is

$\gamma =$ pump field

$$\hat{H}_i = \hbar K [\gamma \hat{a} \hat{b} - \gamma^* \hat{a}^\dagger \hat{b}^\dagger]$$

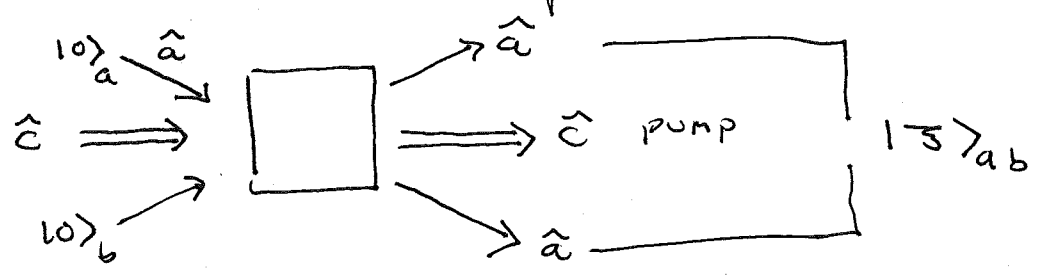


This implies a two mode evolution

$$\hat{U} = e^{-i\gamma \hat{H}_{int}}$$

$$= \exp \left[\frac{\gamma}{2} (\hat{a} \hat{b} + \hat{a}^\dagger \hat{b}^\dagger) \right] = \hat{S}_{\frac{\gamma}{2}}^{ab}$$

where $\frac{\gamma}{2} = r e^{i\theta}$ is the squeeze operator which inherits the phase of the pump.



where $|0\rangle_a |0\rangle_b \Rightarrow \hat{S}_{\frac{\gamma}{2}}^{ab} |0\rangle_a |0\rangle_b = |z\rangle_{ab}$

This is a two-mode squeezed state

The squeezing is distributed across the two modes. Note $|z\rangle_{ab} \neq |z\rangle_a |z\rangle_b$

and hence entanglement can occur; that is quantum correlations between modes. The two-mode quadratures can be defined

$$\hat{X}_1^{ab} \equiv \frac{1}{2^{3/2}} [(\hat{a} + \hat{a}^\dagger) + (\hat{b} + \hat{b}^\dagger)] \approx \left[\hat{X}_1^a + \hat{X}_1^b \right] \frac{1}{\sqrt{2}}$$

$$\hat{X}_2^{ab} \equiv \frac{1}{2^{3/2} i} [(\hat{a} - \hat{a}^\dagger) + (\hat{b} - \hat{b}^\dagger)] \approx \left[\hat{X}_2^a + \hat{X}_2^b \right] \frac{1}{\sqrt{2}}$$

Recall $[X_1^a, X_2^a] = [X_1^b, X_2^b] = \frac{i}{2}$

$$\begin{aligned} \Rightarrow [\hat{X}_1^{ab}, \hat{X}_2^{ab}] &= \frac{1}{2} [X_1^a + X_1^b, X_2^a + X_2^b] \\ &= \frac{1}{2} \{ [X_1^a, X_2^a] + [X_1^a, X_2^b] + [X_1^b, X_2^a] + [X_1^b, X_2^b] \} \\ &= \frac{1}{2} \left\{ \frac{i}{2} + \frac{i}{2} \right\} = i/2 \end{aligned}$$

So $[\hat{X}_i^{ab}]$ obey same commutator as X_i^a or X_i^b

The two-mode squeeze xfrm is

$$\begin{aligned} \hat{S}_{ab}^\dagger(\zeta) \hat{a} \hat{S}_{ab}(\zeta) &= \hat{a} \cosh r - e^{i\theta} \hat{b}^\dagger \sinh r \\ \hat{S}_{ab}^\dagger(\zeta) \hat{b} \hat{S}_{ab}(\zeta) &= \hat{b} \cosh r - e^{i\theta} \hat{a}^\dagger \sinh r \end{aligned}$$

Proof: since $[a, b] = 0$ this is fairly easy

with C.B.H.

$$e^{i\lambda \hat{X}} \hat{Y} e^{-i\lambda \hat{X}} = \sum_{n=0}^{\infty} \frac{(i\lambda)^n}{n!} [\hat{X}, \hat{Y}]^{(n)}$$

$$\text{First } \hat{X} = [\zeta^* \hat{a} \hat{b} - \zeta \hat{a}^\dagger \hat{b}^\dagger]^\dagger = -\hat{X} \quad \text{so let } \lambda = i$$

$$\hat{S}_{ab}^\dagger(\zeta) \hat{a} \hat{S}_{ab}(\zeta) = \sum_{n=0}^{\infty} \frac{(-1)^n}{n!} [\hat{X}, \hat{a}]^{(n)}$$

$$\text{but } [\hat{X}, \hat{a}]^{(1)} = \zeta^* [\hat{a}, \hat{a}] \hat{b} - \zeta [\hat{a}, \hat{a}^\dagger] \hat{b}^\dagger = -\zeta \hat{b}^\dagger$$

$$[\hat{X}, [\hat{X}, \hat{a}]] = (\zeta^* \hat{a})(-\zeta) = -\zeta \hat{a}$$

$$\text{even terms} \Rightarrow \hat{a} \sum_{\text{even}} \frac{\zeta^n}{n!} = \hat{a} \cosh r$$

$$\text{odd terms} \Rightarrow -e^{i\theta} \hat{b}^\dagger \sum_{\text{odd}} \frac{\zeta^n}{n!} = -e^{i\theta} \hat{b}^\dagger \sinh r$$

$$\Rightarrow \hat{S}^\dagger a \hat{S} = \hat{a} \cosh r - e^{i\theta} \hat{b}^\dagger \sinh r$$

Note since $[a, b] = 0$ $\hat{S}_{ab} = \hat{S}_{ba}$

Hence we may just make the substitution:

$$\hat{a} \rightarrow \hat{b} ; \hat{b}^\dagger \rightarrow \hat{a}^\dagger$$

$$\Rightarrow \hat{S}^\dagger \hat{b} \hat{S} = \hat{b} \cosh r - e^{i\theta} \hat{a}^\dagger \sinh r \quad \therefore$$

Using these transforms we can compute the quadratures $\Delta X_1^2 ; \Delta X_2^2$

Note

$$\begin{aligned} & \langle \xi | \hat{X}_1 | \xi \rangle \\ &= \langle 0 | \hat{S}^\dagger \hat{X}_1 \hat{S} | 0 \rangle \\ &= \frac{1}{2^{3/2}} \langle 0 | \hat{S}^\dagger [a + a^\dagger + b + b^\dagger] \hat{S} | 0 \rangle \\ &= \frac{1}{2^{3/2}} \langle 0 | \hat{S}^\dagger a \hat{S} + \hat{S}^\dagger a^\dagger \hat{S} + \hat{S}^\dagger b \hat{S} + \hat{S}^\dagger b^\dagger \hat{S} | 0 \rangle \\ &= \frac{1}{2^{3/2}} \langle 0 | \{a, a^\dagger, b^\dagger, b\} | 0 \rangle \quad \text{always zero} \end{aligned}$$

= 0 By inspection

same for $\langle \xi | \hat{X}_2^{ab} | \xi \rangle = 0$

Hence $\langle \hat{X}_i^{ab} \rangle^2 = 0$

$$\Rightarrow \boxed{\Delta X_i^{ab 2} = \langle \hat{X}_i^{ab 2} \rangle} - \langle \hat{X}_i \rangle^2$$

TRICK!

Let $\hat{c} \equiv \frac{\hat{a} + \hat{b}}{\sqrt{2}}$ $\hat{c}^\dagger = \frac{\hat{a}^\dagger + \hat{b}^\dagger}{\sqrt{2}}$ Hence

$$\hat{X}_1^{ab} = \frac{1}{2} [\hat{c} + \hat{c}^\dagger] = \frac{1}{\sqrt{2}} \left[\frac{\hat{c} + \hat{c}^\dagger}{2} \right] \equiv \frac{1}{\sqrt{2}} \hat{X}_1^c$$

$$\hat{X}_2^{ab} = \frac{1}{2i} [\hat{c} - \hat{c}^\dagger] = \frac{1}{\sqrt{2}} \left[\frac{\hat{c} - \hat{c}^\dagger}{2i} \right] \equiv \frac{1}{\sqrt{2}} \hat{X}_2^c$$

We may treat "c" as fictitious single mode

Hence $\Delta X_i^{ab} = \Delta X_i^c$ since $[a, b]$ commute!
 $[c, c^\dagger] = \frac{1}{2} ([a, a^\dagger] + [b, b^\dagger]) = i$

However $\hat{S}^\dagger \hat{c} \hat{S} = (a \cosh r - e^{i\theta} b^\dagger \sinh r) + (b \cosh r - e^{i\theta} a^\dagger \sinh r)$

$$\begin{aligned} \hat{S}^\dagger \hat{c} \hat{S} &= \hat{c} \cosh r - e^{i\theta} \hat{c}^\dagger \sinh r \\ \hat{S}^\dagger \hat{c}^\dagger \hat{S} &= \hat{c}^\dagger \cosh r - e^{-i\theta} \hat{c} \sinh r \end{aligned}$$

But these are single mode squeeze xfrms; Eq. 7.12

with the identification $\hat{c} \leftrightarrow \hat{a}$. Hence we may

lift Eq. 7.16 & 7.17 for single mode case

$$\Delta X_i^{ab} \equiv \Delta X_i^c = \frac{1}{4} \left\{ \cosh^2 r + \sinh^2 r \mp 2 \sinh r \cosh r \cos \theta \right\}$$

$$= \frac{1}{4} \left[\cosh(2r) \mp \sinh 2r \cos \theta \right]$$

Let $\theta = 0 \Rightarrow$

$$\begin{aligned} \Delta X_i^2 &= \frac{1}{4} [\cosh 2r \mp \sinh 2r] \\ &= \frac{1}{4} \left[\left(\frac{e^{2r} + e^{-2r}}{2} \right) \mp \left(\frac{e^{2r} - e^{-2r}}{2} \right) \right] \\ &= \frac{1}{4} \begin{cases} e^{-2r} & < 1/4 \\ e^{2r} & > 1/4 \end{cases} \end{aligned}$$

So there is squeezing in X_1

The decomposition into number states now has form

$$|\xi\rangle_{ab} = \sum_{nm} c_{nm} |n\rangle_a |m\rangle_b$$

Hence there is entanglement in number!

$$\begin{aligned} \text{Note } \hat{a} |00\rangle &= \hat{a} |0\rangle_a |0\rangle_b = 0 \\ \hat{b} |00\rangle &= |0\rangle_a \hat{b} |0\rangle_b = 0 \end{aligned}$$

IN SIMILAR PROCEDURE TO SINGLE MODE

$$\begin{aligned} \hat{S}^\dagger \hat{a} \hat{S} &\longrightarrow \hat{S} \hat{a} \hat{S}^\dagger = \hat{a} \cosh r + e^{i\theta} \hat{b}^\dagger \sinh r \\ \hat{S} \hat{b} \hat{S}^\dagger &= \hat{b} \cosh r + e^{i\theta} \hat{a}^\dagger \sinh r \end{aligned}$$

Hence

$$\begin{aligned} \hat{a} |0\rangle &= 0 \\ \Rightarrow \hat{S}^\dagger [\hat{a} \hat{S} \hat{S}^\dagger |0\rangle] &= 0 \\ \Rightarrow (\hat{S} \hat{a} \hat{S}^\dagger) |\xi\rangle &= 0 \quad ; \quad (\hat{S} \hat{b} \hat{S}^\dagger) |\xi\rangle = 0 \end{aligned}$$

Let $\mu \equiv \cosh r$ $\nu = e^{i\theta} \sinh r$

$$\Rightarrow [\mu \hat{a} + \nu \hat{b}^\dagger] |\mathbb{Z}\rangle = 0 \quad \text{and} \quad [\mu \hat{b} + \nu \hat{a}^\dagger] |\mathbb{Z}\rangle = 0$$

$$\Leftrightarrow |\mathbb{Z}\rangle_{ab} \equiv |\mathbb{Z}\rangle_{ba} \quad \forall n$$

$$\Leftrightarrow \boxed{|\mathbb{Z}\rangle = \sum_n C_n |n\rangle_a |n\rangle_b}$$

$$\begin{aligned} (\mu \hat{a} + \nu \hat{b}^\dagger) |\mathbb{Z}\rangle &= \sum_{n=0}^{\infty} C_n \mu \sqrt{n} |n-1\rangle_a |n\rangle_b \quad \leftarrow \text{resum} \\ &+ \sum_{n=0}^{\infty} C_n \nu \sqrt{n+1} |n\rangle_a |n+1\rangle_b \quad \leftarrow \text{don't} \end{aligned}$$

$$= \sum_{n=0}^{\infty} C_{n+1} \mu \sqrt{n+1} |n\rangle |n+1\rangle + \sum_{n=0}^{\infty} C_n \nu \sqrt{n+1} |n\rangle |n+1\rangle = 0$$

$$\Rightarrow \boxed{C_{n+1} = -\frac{\nu}{\mu} C_n}$$

$$\Rightarrow C_1 = -\frac{\nu}{\mu} C_0$$

$$C_2 = -\frac{\nu}{\mu} C_1 = \left(-\frac{\nu}{\mu}\right)^2 C_0$$

\vdots

$$C_n = \left(-\frac{\nu}{\mu}\right)^n C_0$$

$$= \left[\frac{-e^{i\theta} \sinh r}{\cosh r} \right]^n C_0$$

$$= \boxed{(-1)^n e^{in\theta} \tanh^n r C_0}$$

To find C_0

$$\boxed{|\mathbb{Z}\rangle_{ab} = C_0 \sum_n (-1)^n e^{in\theta} \tanh^n r |n\rangle |n\rangle}$$

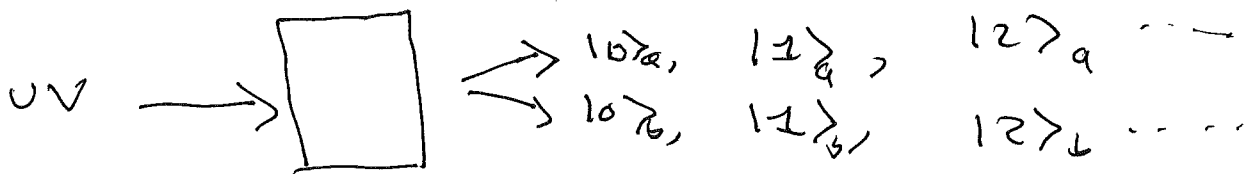
Demand

$$\begin{aligned}
 1 &= \langle \bar{3} | \bar{3} \rangle_{ab} = |c_0|^2 \sum_n (\tanh^2 r)^n \quad (0 < \tanh r < 1) \\
 &= |c_0|^2 \frac{1}{1 - \tanh^2 r} \\
 &= |c_0|^2 \frac{1}{\text{sech}^2 r} = |c_0|^2 \cosh^2 r
 \end{aligned}$$

WLOG

$$c_0 = \frac{1}{\cosh r} = \text{sech } r$$

$$|\bar{3}\rangle_{ab} = \text{sech } r \sum_n (-1)^n e^{in\theta} \tanh^n r |n\rangle_a |n\rangle_b$$



Spontaneous Parametric Downconversion

$$r \ll 1 \Rightarrow \text{sech } r \cong 1 \quad \tanh r \cong r$$

$$\begin{aligned}
 |\bar{3}\rangle_{ab}^{r \ll 1} &\approx |0\rangle_a |0\rangle_b + e^{i(\theta + \pi)} r |1\rangle_a |1\rangle_b + \cancel{\theta (r^2)} |2\rangle_a |2\rangle_b \\
 &\quad \text{MOSTLY VACUUM} \qquad \text{SMALL \# } |1\rangle|1\rangle \qquad \text{NEG. } |2\rangle|2\rangle
 \end{aligned}$$

neglect

The two modes $|n\rangle_a |n\rangle_b$ are strongly correlated in number. Let's ask: suppose I detect "m" in mode "a" what will I detect in mode b?

$$|m\rangle_a \langle m| \overset{\text{projector}}{\mathbb{I}}_{ab} = |m\rangle_a \sum_n c_n \overset{\delta_{mn}}{a} \langle m|n\rangle_a |n\rangle_b$$

$$= \underbrace{c_n |m\rangle_a}_{\text{state after collapse}} |m\rangle_b$$

renormalize \Rightarrow $\boxed{(|m\rangle_a \langle m|) \mathbb{I}_{ab} = |m\rangle_a |m\rangle_b}$

That is if I measure "5" photons in mode "a" I will then see "5" photons in mode "b". Highly correlated!

consider the difference operator

$$\hat{M}_{ab} \equiv \hat{n}_a - \hat{n}_b = a^\dagger a - b^\dagger b$$

$$\begin{aligned} \langle \mathbb{I} | \hat{M}_{ab} | \mathbb{I} \rangle_{ab} &= \sum_n c_n \left[\overset{1}{a} \langle n | \hat{n}_a | n \rangle_a \overset{1}{b} \langle n | n \rangle_b - \overset{1}{a} \langle n | n \rangle_a \overset{1}{b} \langle n | \hat{n}_b | n \rangle_b \right] \\ &= \sum_n c_n [n - n] \\ &\equiv 0 \end{aligned}$$

Again indicates highly correlated in number

$$\begin{aligned} \langle \mathbb{I} | \hat{n}_a | \mathbb{I} \rangle &= \langle 0 | \langle 0 | \hat{S} \hat{a}^\dagger \hat{S}^\dagger a | 0 \rangle | 0 \rangle \\ &= \langle 0 | \langle 0 | [a^\dagger c - e^* b^\dagger / z] [a c - e b^\dagger / z] | 0 \rangle | 0 \rangle \\ &= + |e|^2 \langle 0 | 0 \rangle_a \langle 0 | b b^\dagger | 0 \rangle_b \\ &= \boxed{\sinh^2 r = \langle \hat{n}_b \rangle} \text{ by symmetry } a \leftrightarrow b \end{aligned}$$

$$\begin{aligned}
\langle \hat{n}_a^2 \rangle &= \langle 0 | s^\dagger a^\dagger s s^\dagger a s s^\dagger a s s^\dagger a s | 0 \rangle \\
&= \langle 0 | [a^\dagger c - e^* b^\dagger r] [a c - e b^\dagger r] [a^\dagger c - e^* b^\dagger r] [a c - e b^\dagger r] | 0 \rangle \\
&= \langle 0 | [a^\dagger a c^2 - e a^\dagger b^\dagger c r - e^* b a c r + b b^\dagger r^2] \quad [a, b] = 0 \text{ etc} \\
&\quad \times [a^\dagger a c^2 - e a^\dagger b^\dagger c r - e^* b a c r + b b^\dagger r^2] | 0 \rangle \\
&= \langle 0 | [a^\dagger a c^2 - e a^\dagger b^\dagger c r - e^* a b c r + b b^\dagger r^2] \\
&\quad \times [a^\dagger a c^2 - e a^\dagger b^\dagger c r - e^* a b c r + b b^\dagger r^2] | 0 \rangle
\end{aligned}$$

$$= \langle 0 | \left[\begin{array}{cccc}
a^\dagger a a^\dagger a c^4 & 0 & 0 & e^\dagger a b b^\dagger c^2 r^2 \\
0 & + e^2 a^\dagger b^\dagger c^2 r^2 & + a^\dagger a b^\dagger b c^2 r^2 & 0 \\
0 & + b b^\dagger a a^\dagger c^2 r^2 & + e^2 a^\dagger b^\dagger c^2 r^2 & 0 \\
b b^\dagger a^\dagger a c^2 r^2 & 0 & 0 & + b b^\dagger b b^\dagger r^4
\end{array} \right] | 0 \rangle$$

$$\begin{aligned}
&= \langle 0 | \langle 0 | [(1 + b^\dagger b)(1 + a^\dagger a) c^2 r^2 + (1 + b^\dagger b)(1 + b^\dagger b) r^4] | 0 \rangle | 0 \rangle \\
&= c^2 r^2 + r^4 = \cosh^2 r \sinh^2 r + \frac{\sinh^4 r}{2}
\end{aligned}$$

$$\Rightarrow \Delta^2 n_a = \sinh^2 r \cosh^2 r = \frac{1}{4} \sinh^2(2r) = \Delta^2 n_b$$

Mandel & Test

$$Q_{ab} = \frac{\Delta^2 n - \langle \hat{n} \rangle}{\langle \hat{n} \rangle} = \frac{\frac{1}{4} \sinh^2(2r) - \sinh^2 r}{\sinh^2 r}$$

$$= \frac{\sinh^2 r \cosh^2 r - \sinh^2 r}{\sinh^2 r} = \frac{\cosh^2 r - 1}{\geq 1} \geq 0 \quad \text{Super Poissonian}$$

Variance $\equiv \Delta n_a^2 = \Delta n_b^2 = \frac{1}{4} \sinh^2(\alpha)$

describes fluctuations/correlations in each mode a, b separately.

Co Variance $\equiv \text{Cov} [\hat{n}_a, \hat{n}_b] \equiv \langle \hat{n}_a \hat{n}_b \rangle - \langle \hat{n}_a \rangle \langle \hat{n}_b \rangle$

describes fluctuations/correlations across both modes a, b.

The Covariance of any two random variables is

$\text{Cov} [X, Y] = \langle XY \rangle - \langle X \rangle \langle Y \rangle$

Classical Prob. Theory.

So covariance of any two observables \hat{X}, \hat{Y}

$\text{Cov} [\hat{X}, \hat{Y}] = \langle \hat{X} \hat{Y} \rangle - \langle \hat{X} \rangle \langle \hat{Y} \rangle$

measures degree of correlation!

simple example $\hat{X} = \hat{X}, \hat{Y} = \hat{X}$

$\text{Cov} [\hat{X}, \hat{X}] = \langle \hat{X}^2 \rangle - \langle \hat{X} \rangle^2 = \text{Variance.}$

Note: $\langle \hat{X} \hat{Y} \rangle = \langle \hat{X} \rangle \langle \hat{Y} \rangle \Rightarrow \text{Cov} [X, Y] = 0$

completely un-correlated

Consider $\hat{O}_P \hat{O}_S$

$\hat{M}_{\pm} \equiv \hat{n}_a \pm \hat{n}_b$

commute!
 $[\hat{n}_a, \hat{n}_b] = 0$

$\Delta M_{\pm}^2 = \langle [\hat{n}_a \pm \hat{n}_b][\hat{n}_a \pm \hat{n}_b] \rangle - \langle \hat{n}_a \pm \hat{n}_b \rangle^2$

$= \langle [\hat{n}_a^2 \pm 2\hat{n}_a\hat{n}_b + \hat{n}_b^2] \rangle - [\langle \hat{n}_a \rangle \pm \langle \hat{n}_b \rangle]^2$

$= \langle \hat{n}_a^2 \rangle \pm 2\langle \hat{n}_a\hat{n}_b \rangle + \langle \hat{n}_b^2 \rangle - [\langle \hat{n}_a \rangle \pm 2\langle \hat{n}_a \rangle \langle \hat{n}_b \rangle + \langle \hat{n}_b \rangle^2]$

$= \Delta^2 n_a + \Delta^2 n_b \pm 2 \text{Cov} [\hat{n}_a, \hat{n}_b]$

Recall $\langle \hat{M}_- \rangle \equiv 0$ Eq 7.170

but $\hat{M}_- |3\rangle = 0 \Rightarrow \hat{M}_-^2 |3\rangle = 0$

$$\Rightarrow \langle 3 | \hat{M}_-^2 | 3 \rangle = 0$$

$$\Rightarrow \Delta M_-^2 = 0 = \Delta n_a^2 + \Delta n_b^2 - 2 \text{cov}[\hat{n}_a, \hat{n}_b]$$

$$\Rightarrow \text{cov}[\hat{n}_a, \hat{n}_b] = \frac{1}{2} [\Delta n_a^2 + \Delta n_b^2] = \frac{1}{4} \sinh^2(2r)$$

$$\begin{array}{l} r \rightarrow 0 \swarrow \\ 0 \\ r \rightarrow \infty \searrow \\ \infty \end{array}$$

so as $r \rightarrow \infty$ the correlations grow as $\frac{1}{4} \cdot \frac{1}{4} e^{4r}$
or very fast!

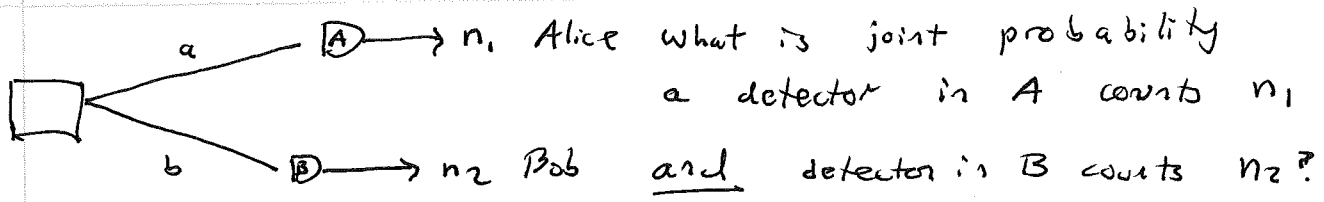
Again from classical prob. theory gives two random variables we define the linear correlation func.

$$J(x, y) \equiv \frac{\text{cov}[x, y]}{\Delta x \cdot \Delta y} \in [0, 1]$$

which is just the normalized covariance

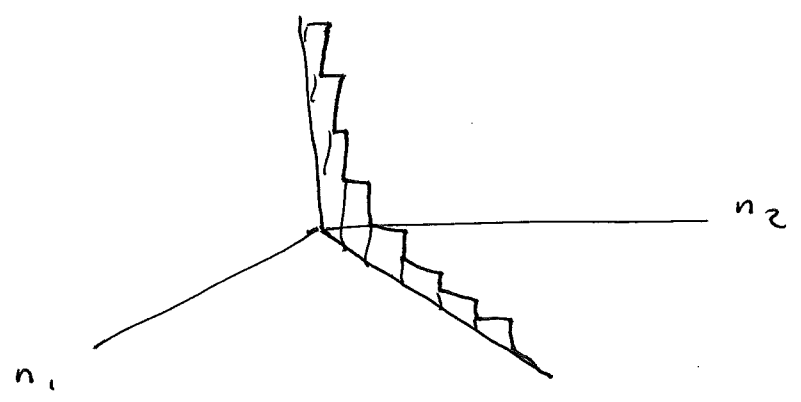
$$J[\hat{n}_a, \hat{n}_b] = \frac{\text{cov}[\hat{n}_a, \hat{n}_b]}{\Delta n_a \cdot \Delta n_b} = \begin{cases} 1 & r > 0 \\ 0 & r = 0 \end{cases}$$

so maximally correlated

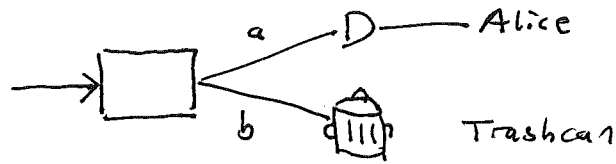


$$\begin{aligned}
 P_{n_1}^{n_2} &\equiv \left| \langle n_1 | \langle n_2 | \bar{\rho} \rangle_{ab} \right|^2 \\
 &= \left| \sum_n C_n \underbrace{\langle n_1 | n \rangle_a}_{\delta_{n_1}^n} \underbrace{\langle n_2 | n \rangle_b}_{\delta_{n_2}^n} \right|^2 \\
 &= \left| C_{n_1} \langle n_2 | n_1 \rangle_b \right|^2 \\
 &= |C_{n_1}|^2 \delta_{n_1}^{n_2} \\
 &= \boxed{\tanh^{2n}(\kappa) \cdot \text{sech}^2(\kappa) \delta_{n_1}^{n_2}}
 \end{aligned}$$

Hence this joint probability is zero unless $n_1 = n_2$
 a measurement by Alice collapses both modes
 into $n \rightarrow n_1$ and hence Bob can only measure n_1
 as well. Since $0 < \tanh^2(\kappa) < 1$ then $P_{n_1}^{n_2}$
 is diagonal



Different set up



Suppose I place a beam dump in mode b — what does Alice see? Alice may not look in trash can (environment) so she must express her ignorance by tracing out mode b . Alice then has a mixed state. Start with pure state density matrix

$$\hat{\rho}_{ab} \equiv |\Xi\rangle_{ab} \langle \Xi|$$

$$= \sum_{nm} C_n^* C_m |n\rangle_a |n\rangle_b \langle m|_b \langle m|_a$$

$$\hat{\rho}_a = \text{tr}_b [\hat{\rho}_{ab}]$$

$$= \sum_{\ell} \langle \ell | \hat{\rho}_{ab} | \ell \rangle_b$$

$$= \sum_{\ell mn} C_n^* C_m |n\rangle_a \underbrace{\langle \ell | n \rangle_b}_{\delta_n^\ell} \underbrace{\langle m | \ell \rangle_b}_{\delta_m^\ell}$$

$$= \boxed{\sum_n |C_n|^2 |n\rangle_a} \quad \text{Alice ignores Bob}$$

By symmetry $a \leftrightarrow b$

$$\boxed{\hat{\rho}_b = \sum_n |C_n|^2 |n\rangle_b} \quad \text{Bob ignores Alice}$$

Recall $|C_n|^2 \equiv P_n = \frac{\tanh^{2n} r}{\text{sech}^2 r}$

We know $\sum_n P_n = \text{sech}^2 r \frac{1}{1 - \tanh^2 r} = 1$

So P_n is a normalized probability distribution.

Also Eq. 7.171 $\Rightarrow \langle \hat{n}_a \rangle = \langle \hat{n}_b \rangle = \sum_n n P_n = \sinh^2 r$

$P_n^{a,b} = \frac{\tanh^{2n} r}{\text{sech}^2 r} \quad \cosh^2 - \sinh^2 = 1$

$= \frac{\tanh^{2n} r}{\cosh^2 r} = \frac{\cancel{\tanh^{2n} r}}{1 + \cancel{\sinh^2 r}} = \frac{[\sinh^2 r]^n}{[\cosh^2 r]^{n+1}} = \frac{[\sinh^2 r]^n}{[1 + \sinh^2 r]^{n+1}}$

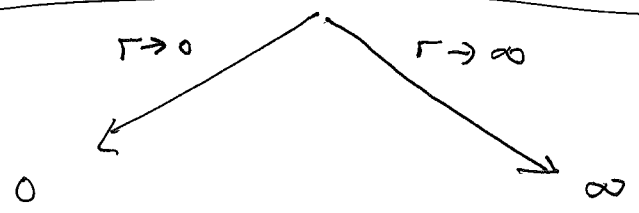
$= \boxed{\frac{\bar{n}^n}{(\bar{n} + 1)^{n+1}}} \quad \text{where } \bar{n} = \bar{n}_a = \bar{n}_b$

Hence Alice E.O. sees a distribution Eq. 2.145 which is a thermal state!

~~$\bar{n} = \sinh^2 r = e^{-\frac{\hbar\omega}{kT}}$~~ $\bar{n} = \sinh^2 r = \frac{1}{e^x - 1}$

$x = \frac{\hbar\omega}{kT} \Rightarrow e^x - 1 = \coth^2 r \text{ csch}^2 r$
 $\Rightarrow e^x = \coth^2 r \text{ csch}^2 r + 1 = \coth^2 r$
 $\Rightarrow x = \ln[\coth^2 r]$

$T = \frac{\hbar\omega}{2k \coth^2 r \ln[\coth^2 r]}$



So squeezing \Leftrightarrow Temperature Effective.

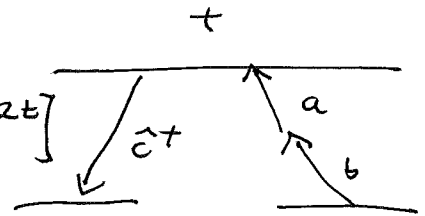
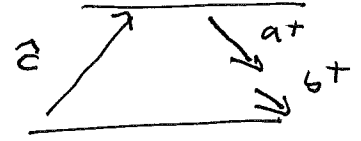
The full hamiltonian is (did this)!

$$\hat{H} = \hbar\omega_a a^\dagger a + \hbar\omega_b b^\dagger b + \hbar\omega_c c^\dagger c + \overbrace{i\hbar\chi_2 [\hat{a}^\dagger \hat{b} \hat{c}^\dagger - \hat{a}^\dagger \hat{b}^\dagger \hat{c}]}^{\hat{H}_{INT}}$$

$$\hat{c} \rightarrow \gamma e^{-i\omega_c t}$$

$$\hat{a} \rightarrow \hat{a} e^{-i\omega_a t}$$

$$\hat{b} \rightarrow \hat{b} e^{-i\omega_b t}$$



$$\Rightarrow \hat{H}_{INT} = i\hbar\chi_2 [\gamma \hat{a} \hat{b} e^{+i\Omega t} - \gamma^* \hat{a}^\dagger \hat{b}^\dagger e^{-i\Omega t}]$$

$$\Omega = \omega_c - \omega_a - \omega_b = 0 \quad \text{cons. Eng.}$$

$$\Rightarrow \hat{H}_{INT} = i\hbar [\zeta^\dagger a b - \zeta a^\dagger b^\dagger]$$

$$\zeta = \chi_2 \gamma$$

$$\Rightarrow U(t) = e^{-\frac{i}{\hbar} \hat{H}_{INT} t} = \hat{S}_{ab}(\bar{\zeta})$$

$$\bar{\zeta} = \zeta t = \chi_2 \gamma t$$