## 9 Quantization of Gauge Fields

We will now turn to the problem of the quantization of the simplest gauge theory, the free electromagnetic field. This is an abelian gauge theory. In Physics 583 we will discuss at length the quantization of non-abelian gauge fields. Unlike abelian theories, such as the free electromagnetic field, even in the absence of matter fields, non-abelian gauge theories are not free fields and have highly non-trivial dynamics. In Physics 582 we will discuss canonical quantization and path-integral quantization of the free electromagnetic gauge field.

### 9.1 Canonical Quantization of the Free Electromagnetic Field

Maxwell's theory was the first field theory to be quantized. However, the quantization procedure involves a number of subtleties not shared by the other problems that we have considered so far. The issue is the fact that this theory has a local gauge invariance. Unlike systems which only have global symmetries, not all the classical configurations of vector potentials represent physically distinct states. It could be argued that one should abandon the picture based on the vector potential and go back to a picture based on electric and magnetic fields instead. However, there is no local Lagrangian that can describe the time evolution of the system now. Furthermore is not clear which fields, $\vec{E}$ or $\vec{B}$ (or some other field) plays the role of coordinates and which can play the role of momenta. For that reason, one sticks to the Lagrangian formulation with the vector potential $A_{\mu}$ as its independent coordinate-like variable.

The Lagrangian for Maxwell's theory

$$
\begin{equation*}
\mathcal{L}=-\frac{1}{4} F_{\mu \nu} F^{\mu \nu} \tag{1}
\end{equation*}
$$

where $F_{\mu \nu}=\partial_{\mu} A_{\nu}-\partial_{\nu} A_{\mu}$, can be written in the form

$$
\begin{equation*}
\mathcal{L}=\frac{1}{2}\left(\vec{E}^{2}-\vec{B}^{2}\right) \tag{2}
\end{equation*}
$$

where

$$
\begin{align*}
E_{j} & =-\partial_{0} A_{j}-\partial_{j} A_{0} \\
B_{j} & =-\epsilon_{j k \ell} \partial_{k} A_{\ell} \tag{3}
\end{align*}
$$

The electric field $E_{j}$ and the space components of the vector potential $A_{j}$ form a canonical pair since, by definition, the momentum $\Pi_{j}$ conjugate to $A_{j}$ is

$$
\begin{equation*}
\Pi_{j}(x)=\frac{\partial \mathcal{L}}{\delta \partial_{0} A_{j}(x)}=\partial_{0} A_{j}+\partial_{j} A_{0}=-E_{j} \tag{4}
\end{equation*}
$$

Notice that since $\mathcal{L}$ does not contain any terms which include $\partial_{0} A_{0}$, the momentum $\Pi_{0}$, conjugate to $A_{0}$, vanishes

$$
\begin{equation*}
\Pi_{0}=\frac{\delta \mathcal{L}}{\delta \partial_{0} A_{0}}=0 \tag{5}
\end{equation*}
$$

A consequence of this result is that $A_{0}$ is essentially arbitrary and it plays the role of a Lagrange multiplier. Indeed it is always possible to find a gauge transformation $\phi$

$$
\begin{equation*}
A_{0}^{\prime}=A_{0}+\partial_{0} \phi \quad A_{j}^{\prime}=A_{j}-\partial_{j} \phi \tag{6}
\end{equation*}
$$

such that $A_{0}^{\prime}=0$. The solution is

$$
\begin{equation*}
\partial_{0} \phi=-A_{0} \tag{7}
\end{equation*}
$$

which is consistent provided that $A_{0}$ vanishes both in the remote part and in the remote future $(t \rightarrow \pm \infty)$.

The canonical formalism can be applied to Maxwell's electrodynamics if we notice that the fields $A_{j}(\vec{x})$ and $\Pi_{j^{\prime}}\left(\vec{x}^{\prime}\right)$ obey the equal-time Poisson Brackets

$$
\begin{equation*}
\left\{A_{j}(\vec{x}), \Pi_{j^{\prime}}\left(\vec{x}^{\prime}\right)\right\}_{P B}=\delta_{j j^{\prime}} \delta^{3}\left(\vec{x}-\vec{x}^{\prime}\right) \tag{8}
\end{equation*}
$$

or, in terms of the electric field $\vec{E}$,

$$
\begin{equation*}
\left\{A_{j}(\vec{x}), E_{j^{\prime}}\left(\vec{x}^{\prime}\right)\right\}_{P B}=-\delta_{j j^{\prime}} \delta^{3}\left(\vec{x}-\vec{x}^{\prime}\right) \tag{9}
\end{equation*}
$$

The classical Hamiltonian density is defined in the usual manner

$$
\begin{equation*}
\mathcal{H}=\Pi_{j} \partial_{0} A_{j}-\mathcal{L} \tag{10}
\end{equation*}
$$

We find

$$
\begin{equation*}
\mathcal{H}(x)=\frac{1}{2}\left(\vec{E}^{2}+\vec{B}^{2}\right)-A_{0}(x) \vec{\nabla} \cdot \vec{E}(x) \tag{11}
\end{equation*}
$$

Except for the last term, this is the usual answer. It is easy to see that the last term is a constant of motion. Indeed the equal-time Poisson Bracket between the Hamiltonian density $\mathcal{H}(\vec{x})$ and $\vec{\nabla} \cdot \vec{E}(\vec{y})$ is zero. By explicit calculation, we get

$$
\begin{equation*}
\{\mathcal{H}(\vec{x}), \vec{\nabla} \cdot \vec{E}(\vec{y})\}_{P B}=\int d^{3} z\left[-\frac{\delta \mathcal{H}(\vec{x})}{\delta A_{j}(\vec{z})} \frac{\delta \vec{\nabla} \cdot \vec{E}(\vec{y})}{\delta E_{j}(\vec{z})}+\frac{\delta \mathcal{H}(\vec{x})}{\delta E_{j}(\vec{z})} \frac{\delta \vec{\nabla} \cdot \vec{E}(\vec{y})}{\delta A_{j}(\vec{z})}\right] \tag{12}
\end{equation*}
$$

But

$$
\begin{align*}
\frac{\delta \mathcal{H}(\vec{x})}{\delta A_{j}(\vec{z})}=\int d^{3} w \frac{\delta \mathcal{H}(\vec{x})}{\delta B_{k}(\vec{w})} \frac{\delta B_{k}(\vec{w})}{\delta A_{j}(\vec{z})} & =\int d^{3} w B_{k}(\vec{w}) \delta(\vec{x}-\vec{w}) \epsilon_{k \ell j} \nabla_{\ell}^{w} \delta(\vec{w}-\vec{z}) \\
& =-\epsilon_{k \ell j} \nabla_{\ell}^{z} \int d^{3} w B_{k}(\vec{w}) \delta(\vec{x}-\vec{w}) \delta(\vec{w}-\vec{z}) \tag{13}
\end{align*}
$$

Hence

$$
\begin{equation*}
\frac{\delta \mathcal{H}(\vec{x})}{\delta A_{j}(\vec{z})}=\epsilon_{j \ell k} \nabla_{\ell}^{z}\left(B_{k}(\vec{x}) \delta(\vec{x}-\vec{z})\right)=\epsilon_{j \ell k} B_{k}(\vec{x}) \nabla_{\ell}^{x} \delta(\vec{x}-\vec{z}) \tag{14}
\end{equation*}
$$

Similarly, we get

$$
\begin{equation*}
\frac{\delta \vec{\nabla} \cdot \vec{E}(\vec{y})}{\delta E_{j}(\vec{z})}=\nabla_{j}^{y} \delta(\vec{y}-\vec{z}) \quad \frac{\delta \vec{\nabla} \cdot \vec{E}(\vec{y})}{\delta A_{j}(\vec{z})}=0 \tag{15}
\end{equation*}
$$

Thus, the Poisson Bracket is

$$
\begin{align*}
\{\mathcal{H}(\vec{x}), \vec{\nabla} \cdot \vec{E}(\vec{y})\}_{P B} & =\int d^{3} z\left[-\epsilon_{j \ell k} B_{k}(\vec{x}) \nabla_{\ell}^{x} \delta(\vec{x}-\vec{z}) \nabla_{j}^{y} \delta(\vec{y}-\vec{z})\right] \\
& =-\epsilon_{j \ell k} B_{k}(\vec{x}) \nabla_{\ell}^{x} \nabla_{j}^{y} \delta(\vec{x}-\vec{y}) \\
& =\epsilon_{j \ell k} B_{k}(\vec{x}) \nabla_{\ell}^{x} \nabla_{j}^{x} \delta(\vec{x}-\vec{y})=0 \tag{16}
\end{align*}
$$

provided that $\vec{B}(\vec{x})$ is non-singular. Thus, $\vec{\nabla} \cdot \vec{E}(\vec{x})$ is a constant of motion. It is easy to check that $\vec{\nabla} \cdot \vec{E}$ generates infinitesimal gauge transformations. We will prove this statement directly in the quantum theory.

Since $\vec{\nabla} \cdot \vec{E}(\vec{x})$ is a constant of motion, if we pick a value for it at some initial time $x_{0}=t_{0}$, it will remain constant in time. Thus we can write

$$
\begin{equation*}
\vec{\nabla} \cdot \vec{E}(\vec{x})=\rho(\vec{x}) \tag{17}
\end{equation*}
$$

which we recognize to be Gauss's Law. Naturally, an external charge distribution may be explicitly time dependent and then

$$
\begin{equation*}
\frac{d}{d t}(\vec{\nabla} \cdot \vec{E})=\frac{\partial}{\partial t}(\vec{\nabla} \cdot \vec{E})=\frac{\partial}{\partial t} \rho_{e x t}(\vec{x}, t) \tag{18}
\end{equation*}
$$

Before turning to the quantization of this theory, we must notice that $A_{0}$ plays the role of a Lagrange multiplier field whose variation forces Gauss's Law, $\vec{\nabla} \cdot \vec{E}=$ 0. Hence Gauss's Law should be regarded as a constraint rather than an equation of motion. This issue becomes very important in the quantum theory. Indeed, without the constraint $\vec{\nabla} \cdot \vec{E}=0$, the theory is absolutely trivial, and wrong.

Constraints impose very severe restrictions on the allowed states of a quantum theory. Consider for instance a particle of mass $m$ moving freely in three dimensional space. Its stationary states have wave functions $\Psi_{\vec{p}}(\vec{r}, t)$

$$
\begin{equation*}
\Psi_{\vec{p}}(\vec{r}, t) \sim e^{i\left(\frac{\vec{p} \cdot \vec{r}-E(\vec{p}) t}{\hbar}\right)} \tag{19}
\end{equation*}
$$

with an energy $E(\vec{p})=\frac{\vec{p}^{2}}{2 m}$. If we constrain the particle to move only on the surface of a sphere of radius $R$, it becomes equivalent to a rigid rotor of moment of inertia $I=m R^{2}$ and energy eigenvalues $\epsilon_{\ell m}=\frac{\hbar^{2}}{2 I} \ell(\ell+1)$ where $\ell=0,1,2, \ldots$,
and $|m| \leq \ell$. Thus, even the simple constraint $\vec{r}^{2}=R^{2}$, does have non-trivial effects.

The constraints that we have to impose when quantizing Maxwell's electrodynamics do not change the energy spectrum. This is so because we can reduce the number of degrees of freedom to be quantized by taking advantage of the gauge invariance of the classical theory. This procedure is called gauge fixing. For example, the classical equation of motion

$$
\begin{equation*}
\square A^{\mu}-\partial^{\mu}\left(\partial_{\nu} A^{\nu}\right)=0 \tag{20}
\end{equation*}
$$

becomes, in the Coulomb gauge $A_{0}=0$ and $\vec{\nabla} \cdot \vec{A}=0$,

$$
\begin{equation*}
\square A_{j}=0 \tag{21}
\end{equation*}
$$

However the Coulomb gauge is not compatible with the Poisson Bracket

$$
\begin{equation*}
\left\{A_{j}(\vec{x}), \Pi_{j},\left(\vec{x}^{\prime}\right)\right\}_{P B}=\delta_{j j^{\prime}} \delta\left(\vec{x}-\vec{x}^{\prime}\right) \tag{22}
\end{equation*}
$$

since the spatial divergence of the $\delta$-function does not vanish. It will follow that the quantization of the theory in the Coulomb gauge is achieved at the price of a modification of the commutation relations. Because the classical theory is gauge invariant we can always fix the gauge without any loss of physical content. The procedure of gauge fixing has the attractive that the number of independent variables is greatly reduced. A standard approach to the quantization of a gauge theory is to fix the gauge first, at the classical level, and to quantize later. However, a number of problems arise immediately. For instance, in most gauges symmetries such as Lorentz invariance are lost, at least manifestly so. Thus, the Coulomb gauge, also known as the radiation or transverse gauge, spoils Lorentz invariance, but it has the attractive that the nature of the physical states (the photons) is quite transparent. We will see below that the quantization of the theory in this gauge has some peculiarities.

Another standard choice is the Lorentz gauge

$$
\begin{equation*}
\partial_{\mu} A^{\mu}=0 \tag{23}
\end{equation*}
$$

whose main appeal is its manifest covariance. The quantization of the system is this gauge follows the method developed by and Gupta and Bleuer. While highly successful, it requires the introduction of states with negative norm (known as ghosts) which cancel-out all the gauge-dependent contributions to physical quantities.

More general covariant gauges can also be defined. A general approach consists not on imposing a rigid restriction on the degrees of freedom, but to add new terms to the Lagrangian which eliminate the gauge freedom. For instance, the modified Lagrangian

$$
\begin{equation*}
\mathcal{L}=-\frac{1}{4} F^{2}+\frac{\alpha}{2}\left(\partial_{\mu} A^{\mu}(x)\right)^{2} \tag{24}
\end{equation*}
$$

is not gauge invariant because of the presence of the last term. We can easily see that this term weighs gauge equivalent configurations differently and the $\alpha$
plays the role of a Lagrange multiplier In fact, in the limit $\alpha \rightarrow \infty$ we recover the Lorentz gauge condition. In the path integral quantization of Maxwell's theory it is proven that this approach is equivalent to an average over gauges of the physical quantities. If $\alpha=1$, the equations of motion become very simple, i.e., $\partial^{2} a_{\mu}=0$. This is the Feynman gauge. This is the gauge in which the calculations are simplest. Still within the Hamiltonian or canonical quantization method, a third approach has been developed. In this approach one fixes the gauge $A_{0}=0$. This condition is not enough to eliminate the gauge freedom. In this gauge a residual set of gauge transformations are still allowed, the timeindependent ones. In this approach quantization is achieved by replacing the Poisson Brackets by commutators and Gauss' Law condition becomes now a constraint on the space of physical quantum states. So, we quantize first and constrain later.

In general, it is a non-trivial task to prove that all the different quantizations yield a theory with the same physical properties. In practice what one has to prove is that these different gauge choices yield theories whose states differ from each other at most by a unitary transformation. Otherwise, the quantized theories would be physically inequivalent. In addition, the recovery of Lorentz invariance may be a bit tedious in some cases. There is however, an alternative, complementary, approach to the quantum theory in which most of these issues become very transparent. This is the path-integral approach. This method has the advantage that all the symmetries are taken care of from the out set. In addition, the canonical methods encounter very serious difficulties in the treatment of the non-abelian generalizations of Maxwell's electrodynamics.

We will consider here two canonical approaches: 1) quantization in the Coulomb gauge and 2) canonical quantization in the $A_{0}=0$ gauge in the Schr'odinger picture.

### 9.2 Coulomb Gauge

Quantization in the Coulomb gauge follows the methods developed for the scalar field very closely. Indeed the classical constraints $A_{0}=0$ and $\vec{\nabla} \cdot \vec{A}=0$ allow for a Fourier expansion of the vector potential $\vec{A}\left(\vec{x}, x_{0}\right)$. In Fourier space we write

$$
\begin{equation*}
\vec{A}\left(\vec{x}, x_{0}\right)=\int \frac{d^{3} p}{(2 \pi)^{3} 2 p_{0}} \vec{A}\left(\vec{p}, x_{0}\right) \exp (i \vec{p} \cdot \vec{x}) \tag{25}
\end{equation*}
$$

where $\vec{A}\left(\vec{p}, x_{0}\right)=\vec{A}^{*}\left(-\vec{p}, x_{0}\right)$. Maxwell's equations yield the classical equation of motion

$$
\begin{equation*}
\square \vec{A}\left(\vec{x}, x_{0}\right)=0 \tag{26}
\end{equation*}
$$

The Fourier expansion is consistent only if $\vec{A}\left(\vec{p}, x_{0}\right)$ satisfies

$$
\begin{equation*}
\partial_{0}^{2} \vec{A}\left(\vec{P}, x_{0}\right)+\vec{p}^{2} \vec{A}\left(\vec{p}, x_{0}\right)=0 \tag{27}
\end{equation*}
$$

The constraint $\vec{\nabla} \cdot \vec{A}=0$ now becomes the transversality condition

$$
\begin{equation*}
\vec{p} \cdot \vec{A}\left(\vec{p}, x_{0}\right)=0 \tag{28}
\end{equation*}
$$

Thus, $\vec{A}\left(\vec{p}, x_{0}\right)$ has the time dependence

$$
\begin{equation*}
\vec{A}\left(\vec{p}, x_{0}\right)=\vec{A}(\vec{p}) e^{i p_{0} x_{0}}+\vec{A}(-\vec{p}) e^{-i p_{0} x_{0}} \tag{29}
\end{equation*}
$$

where $p_{0}=|\vec{p}|$. Then, the expansion takes the form

$$
\begin{equation*}
\vec{A}\left(\vec{x}, x_{0}\right)=\int \frac{d^{3} p}{(2 \pi)^{3} 2 p_{0}}\left[\overrightarrow{A^{*}}(\vec{p}) e^{i p \cdot x}+\vec{A}(\vec{p}) e^{-i p \cdot x}\right] \tag{30}
\end{equation*}
$$

where $p \cdot x=p_{\mu} x^{\mu}$. The transversality condition is satisfied by introducing two polarization unit vectors $\vec{\epsilon}(\vec{p})$ and $\vec{\epsilon}_{2}(\vec{p})$ such that $\vec{\epsilon}_{1} \cdot \vec{\epsilon}_{2}=\vec{\epsilon}_{1} \cdot \vec{p}=\vec{\epsilon}_{2} \cdot \vec{p}=0$ and $\vec{\epsilon}_{1}^{2}=\vec{\epsilon}_{2}^{2}=1$. Hence if $\vec{A}$ has to be orthogonal to $\vec{p}$, it must be a linear combination of $\vec{\epsilon}_{1}$ and $\vec{\epsilon}_{2}$, i.e.,

$$
\begin{equation*}
\vec{A}(\vec{p})=\sum_{\alpha=1,2} \vec{\epsilon}_{\alpha}(\vec{p}) a_{\alpha}(\vec{p}) \tag{31}
\end{equation*}
$$

where the factors $a_{\alpha}(\vec{p})$ are complex amplitudes. In terms of $a_{\alpha}(\vec{p})$ and $a_{\alpha}^{*}(\vec{p})$ the Hamiltonian looks like a sum of oscillators.

The passage to the quantum theory is achieved by assigning to each amplitude $a_{\alpha}(\vec{p})$ a Heisenberg operator $\hat{a}(\vec{p})$. Similarly $a_{\alpha}^{*}(\vec{p})$ maps onto the adjoint operator $\hat{a}_{\alpha}^{\dagger}(\vec{p})$. The expansion of the vector potential now is

$$
\begin{equation*}
\hat{\vec{A}}(x)=\int \frac{d^{3} p}{(2 \pi)^{3} 2 p_{0}} \sum_{\alpha=1,2} \vec{\epsilon}_{\alpha}(p)\left[\hat{a}_{\alpha}(\vec{p}) e^{-i p \cdot x}+\hat{a}_{\alpha}^{\dagger}(\vec{p}) e^{i p \cdot x}\right] \tag{32}
\end{equation*}
$$

with $p^{2}=0$ and $p_{0}=|\vec{p}|$. The operators $\hat{a}_{\alpha}(\vec{p})$ and $\hat{a}_{\alpha}^{\dagger}(\vec{p})$ satisfy commutation relations

$$
\begin{align*}
{\left[\hat{a}_{\alpha}(\vec{p}), \hat{a}_{\alpha^{\prime}}^{\dagger}\left(\vec{p}^{\prime}\right)\right] } & =2 p_{0}(2 \pi)^{3} \delta\left(\vec{p}-\vec{p}^{\prime}\right) \\
{\left[\hat{a}_{\alpha}(\vec{p}), \hat{a}_{\alpha^{\prime}}\left(\vec{p}^{\prime}\right)\right] } & =\left[\hat{a}_{\alpha}^{\dagger}(\vec{p}), \hat{a}_{\alpha^{\prime}}^{\dagger}\left(\vec{p}^{\prime}\right)\right]=0 \tag{33}
\end{align*}
$$

It is straightforward to check that the vector potential $\vec{A}(\vec{x})$ and the electric field $\vec{E}(\vec{x})$ obey the (unconventional) equal-time commutation relation

$$
\begin{equation*}
\left[A_{j}(\vec{x}), E_{j^{\prime}}\left(\vec{x}^{\prime}\right)\right]=-i\left(\delta_{j j^{\prime}}-\frac{\nabla j \nabla j^{\prime}}{\nabla^{2}}\right) \delta^{3}\left(\vec{x}-\vec{x}^{\prime}\right) \tag{34}
\end{equation*}
$$

where the symbol $1 / \nabla^{2}$ represents the inverse of the Laplacian, i.e., the Laplacian Green function. In the derivation of this relation, the following identity was used

$$
\begin{equation*}
\sum_{\alpha=1,2} \epsilon_{\alpha}^{j}(\vec{p}) \epsilon_{\alpha}^{j^{\prime}}(\vec{p})=\delta_{j j^{\prime}}-\frac{p_{j} p_{j^{\prime}}}{\vec{p}^{2}} \tag{35}
\end{equation*}
$$

These commutation relations are an extension of the canonical ones and it is consistent with the transversality condition $\vec{\nabla} \cdot \vec{A}=0$

In this gauge the (normal-ordered) Hamiltonian is

$$
\begin{equation*}
\hat{H}=\int \frac{d^{3} p}{(2 \pi)^{3} 2 p_{0}} p_{0} \sum_{\alpha=1,2} \hat{a}_{\alpha}^{\dagger}(\vec{p}) \hat{a}_{\alpha}(\vec{p}) \tag{36}
\end{equation*}
$$

The ground state $|0\rangle$ is annihilated by both polarizations $\hat{a}_{\alpha}(\vec{p})|0\rangle=0$. The single-particle states are $\hat{a}_{\alpha}^{\dagger}(\vec{p})|0\rangle$ and represent photons with momentum $\vec{p}$, energy $p_{0}=|\vec{p}|$ and with the two possible linear polarizations labelled by $\alpha=$ 1,2 . Circularly polarized photons can be constructed in the usual manner.

### 9.3 The Gauge $A_{0}=0$

In this gauge we will apply directly the canonical formalism. In what follows we will fix $A_{0}=0$ and associate to the three spatial components $A_{j}$ of the vector potential an operator, $\hat{A}_{j}$ which acts on a Hilbert space of states. Similarly, to the canonical momentum $\Pi_{j}=-E_{j}$, we assign an operator $\widehat{\Pi}_{j}$. These operators obey the equal-time commutation relations

$$
\begin{equation*}
\left[\widehat{A}_{j}(\vec{x}), \widehat{\Pi}_{j^{\prime}}\left(\vec{x}^{\prime}\right)\right]=i \delta\left(\vec{x}-\vec{x}^{\prime}\right) \delta_{j j^{\prime}} \tag{37}
\end{equation*}
$$

Hence the vector potential $\vec{A}$ and the electric field $\vec{E}$ do not commute since they are canonically conjugate operators

$$
\begin{equation*}
\left[\hat{A}_{j}(\vec{x}), \hat{E}_{j^{\prime}}\left(\vec{x}^{\prime}\right)\right]=-i \delta_{j j^{\prime}} \delta\left(\vec{x}-\vec{x}^{\prime}\right) \tag{38}
\end{equation*}
$$

Let us now specify the Hilbert space to be the space of states $|\Psi\rangle$ with wave functions which, in the field representation, have the form $\Psi\left(\left\{A_{j}(\vec{x})\right\}\right)$. When acting on these states, the electric field is the functional differential operator

$$
\begin{equation*}
\hat{E}_{j}(\vec{x}) \equiv i \frac{\delta}{\delta A_{j}(\vec{x})} \tag{39}
\end{equation*}
$$

In this Hilbert space, the inner product is

$$
\begin{equation*}
\left.\left\langle\left\{A_{j}(\vec{x})\right\} \mid\left\{A_{j}(\vec{x})\right\}\right\rangle\right\rangle \equiv \Pi_{\vec{x}, j} \delta\left(A_{j}(\vec{x})-A_{j}(\vec{x})\right) \tag{40}
\end{equation*}
$$

This Hilbert space is actually much too large. Indeed states with wave functions that differ by time-independent gauge transformations

$$
\begin{equation*}
\Psi_{\phi}\left(\left\{A_{j}(\vec{x})\right\}\right) \equiv \Psi\left(\left\{A_{j}(\vec{x})-\nabla_{j} \phi(\vec{x})\right\}\right) \tag{41}
\end{equation*}
$$

are physically equivalent since the matrix elements of the electric field operator $\hat{E}_{j}(\vec{x})$ and magnetic field operator $\hat{B}_{j}(\vec{x})=\epsilon_{j k \ell} \nabla_{k} \hat{A}_{\ell}(\vec{x})$ are the same for all gauge-equivalent states, i.e.,

$$
\begin{align*}
\left\langle\Psi_{\phi^{\prime}}^{\prime}\left(\left\{A_{j}(\vec{x})\right\}\right)\right| \hat{E}_{j}(\vec{x})\left|\Psi_{\phi}\left(\left\{A_{j}(\vec{x})\right\}\right)\right\rangle & =\left\langle\Psi^{\prime}\left(\left\{A_{j}(\vec{x})\right\}\right)\right| \hat{E}_{j}(\vec{x})\left|\Psi\left(\left\{A_{j}(\vec{x})\right\}\right)\right\rangle \\
\left\langle\Psi_{\phi^{\prime}}^{\prime}\left(\left\{A_{j}(\vec{x})\right\}\right)\right| \hat{B}_{j}(\vec{x})\left|\Psi_{\phi}\left(\left\{A_{j}(\vec{x})\right\}\right)\right\rangle & =\left\langle\Psi^{\prime}\left(\left\{A_{j}(\vec{x})\right\}\right)\right| \hat{B}_{j}(\vec{x})\left|\Psi\left(\left\{A_{j}(\vec{x})\right\}\right)\right\rangle \tag{42}
\end{align*}
$$

The (local) operator $\hat{Q}(\vec{x})$

$$
\begin{equation*}
\hat{Q}(\vec{x})=\nabla{ }_{j} \hat{E}_{j}(\vec{x}) \tag{43}
\end{equation*}
$$

commutes locally with the Hamiltonian

$$
\begin{equation*}
[\hat{Q}(\vec{x}), \hat{H}]=0 \tag{44}
\end{equation*}
$$

and, hence, it can be diagonalized simultaneously with $\hat{H}$. Let us show now that $\hat{Q}(\vec{x})$ generates local infinitesimal time-independent gauge transformations. From the canonical commutation relation

$$
\begin{equation*}
\left[\hat{A}_{j}(\vec{x}), \hat{E}_{j^{\prime}}\left(\vec{x}^{\prime}\right)\right]=-i \delta_{j j^{\prime}} \delta\left(\vec{x}-\vec{x}^{\prime}\right) \tag{45}
\end{equation*}
$$

we get (by differentiation)

$$
\begin{equation*}
\left[\hat{A}_{j}(\vec{x}), \hat{Q}\left(\vec{x}^{\prime}\right)\right]=\left[\hat{A}_{j}(\vec{x}), \nabla_{j} \hat{E}_{j^{\prime}}\left(\vec{x}^{\prime}\right)\right]=i \nabla_{j}^{x} \delta\left(\vec{x}-\vec{x}^{\prime}\right) \tag{46}
\end{equation*}
$$

Hence, we also find

$$
\begin{equation*}
\left[i \int d \vec{z} \phi(\vec{z}) \hat{Q}(\vec{z}), \hat{A}_{j}(\vec{x})\right]=-\int d \vec{z} \phi(\vec{z}) \nabla_{j}^{z} \delta(\vec{z}-\vec{x})=\nabla_{j} \phi(\vec{x}) \tag{47}
\end{equation*}
$$

and

$$
\begin{align*}
& e^{i \int d \vec{z} \phi(\vec{z}) \hat{Q}(\vec{z})} \hat{A}_{j}(\vec{x}) e^{-i \int d \vec{z} \phi(\vec{z}) \hat{Q}(\vec{z})}= \\
& \quad=e^{-i \int d \vec{z} \nabla_{k} \phi(\vec{z}) \hat{E}_{k}(\vec{z})} \hat{A}_{j}(\vec{x}) e^{\left.i \int d \vec{z}\right) \nabla_{k} \phi(\vec{z}) \hat{E}_{k}(\vec{z})} \\
& \quad=\hat{A}_{j}(\vec{x})+\nabla_{j} \phi(\vec{x}) \tag{48}
\end{align*}
$$

The physical requirement that states that differ by time-independent gauge transformations be equivalent to each other leads to the demand that we should restrict the Hilbert space to the space of gauge-invariant states. These states, which we will denote by $\mid$ Phys $\rangle$, satisfy

$$
\begin{equation*}
\hat{Q}(\vec{x}) \mid \text { Phys }\rangle \equiv \vec{\nabla} \cdot \hat{\vec{E}}(\vec{x}) \mid \text { Phys }\rangle=0 \tag{49}
\end{equation*}
$$

Thus, the constraint means that only the states which obey Gauss' law are in the physical Hilbert space. Unlike the quantization in the Coulomb gauge, in the $A_{0}=0$ gauge the commutators are canonical and the states are constrained to obey Gauss' law.

In the Schrödinger picture, the eigenstates of the system obey the Schrödinger equation

$$
\begin{equation*}
\int d \vec{x} \frac{1}{2}\left[-\frac{\delta^{2}}{\delta A_{j}(\vec{x})^{2}}+B_{j}(\vec{x})^{2}\right] \Psi[A]=\mathcal{E} \Psi[A] \tag{50}
\end{equation*}
$$

where $\Psi[A]$ is a shorthand for the wave functional $\Psi\left(\left\{A_{j}(\vec{x})\right\}\right)$. In this notation, the constraint of Gauss' law is

$$
\begin{equation*}
\nabla_{j}^{x} \hat{E}_{j}(\vec{x}) \Psi[A] \equiv i \nabla_{j}^{x} \frac{\delta}{\delta A_{j}(\vec{x})} \Psi[A]=0 \tag{51}
\end{equation*}
$$

This constraint can be satisfied by separating the real field $A_{j}(\vec{x})$ into longitudinal $A_{j}^{L}(\vec{x})$ and transverse $A_{j}^{T}(\vec{x})$ parts

$$
\begin{equation*}
A_{j}(\vec{x})=A_{j}^{L}(\vec{x})+A_{j}^{T}(\vec{x})=\int \frac{d^{3} p}{(2 \pi)^{3}}\left(A_{j}^{L}(\vec{p})+A_{j}^{T}(\vec{p})\right) e^{i \vec{p} \cdot \vec{x}} \tag{52}
\end{equation*}
$$

where $A_{j}^{L}(\vec{x})$ and $A_{j}^{T}(\vec{x})$ satisfy

$$
\begin{equation*}
\nabla_{j} A_{j}^{T}(\vec{x})=0 \quad A_{j}^{L}(\vec{x})=\nabla_{j} \phi(\vec{x}) \tag{53}
\end{equation*}
$$

and $\phi(\vec{x})$ is, for the moment, arbitrary. In terms of $A_{j}^{L}$ and $A_{j}^{T}$ the constraint of Gauss' law simply becomes

$$
\begin{equation*}
\nabla_{j}^{x} \frac{\delta}{\delta A_{j}^{L}(\vec{x})} \Psi[A]=0 \tag{54}
\end{equation*}
$$

and the Hamiltonian now is

$$
\begin{equation*}
\hat{H}=\int d^{3} p \frac{1}{2}\left\{-\frac{\delta^{2}}{\delta A_{j}^{T}(\vec{p}) \delta A_{j}^{T}(\overrightarrow{-} p)}-\frac{\delta^{2}}{\delta A_{j}^{L}(\vec{p}) \delta A_{j}^{L}(\overrightarrow{-} p)}+\vec{p}^{2} A_{j}^{T}(\vec{p}) A_{j}^{T}(\overrightarrow{-} p)\right\} \tag{55}
\end{equation*}
$$

We satisfy the constraint by looking only at gauge-invariant states. Their wave functions do not depend on the longitudinal components of $\vec{A}(\vec{x})$. Hence, $\Psi[A]=$ $\Psi\left[A^{T}\right]$. When acting on those states, the Hamiltonian is

$$
\begin{equation*}
H \Psi=\int d^{3} p \frac{1}{2}\left[-\frac{\delta^{2}}{\delta A_{j}^{T}(\vec{p}) \delta A_{j}^{T}(\overrightarrow{-} p)}+\vec{p}^{2} A_{j}^{T}(\vec{p}) A_{j}^{T}(\overrightarrow{-} p)\right] \Psi=\mathcal{E} \Psi \tag{56}
\end{equation*}
$$

Let $\vec{\epsilon}_{1}(\vec{p})$ and $\vec{\epsilon}_{2}(\vec{p})$ be two vectors which together with the unit vector $\vec{n}_{p}=\vec{p} /|\vec{p}|$ form an orthonormal basis. Let us define the operators ( $\alpha=1,2 ; j=1,2,3$ )

$$
\begin{align*}
\hat{a}(\vec{p}, \alpha) & =\frac{1}{\sqrt{2|\vec{p}|}} \epsilon_{j}^{\alpha}(\vec{p})\left[\frac{\delta}{\delta A_{j}^{T}(-\vec{p})}+|\vec{p}| A_{j}^{T}(\vec{p})\right] \\
\hat{a}^{\dagger}(\vec{p}, \alpha) & =\frac{1}{\sqrt{2|\vec{p}|}} \epsilon_{j}^{\alpha}(\vec{p})\left[-\frac{\delta}{\delta A_{j}^{T}(\vec{p})}+|\vec{p}| A_{j}^{T}(-\vec{p})\right. \tag{57}
\end{align*}
$$

These operators satisfy the commutation relations

$$
\begin{equation*}
\left[\hat{a}(\vec{p}, \alpha), \hat{a}^{\dagger}\left(\vec{p}^{\prime}, \alpha^{\prime}\right)\right]=\delta_{\alpha \alpha^{\prime}} \delta^{3}\left(\vec{p}-\vec{p}^{\prime}\right) \tag{58}
\end{equation*}
$$

In terms of these operators, the Hamiltonian $\hat{H}$ and the expansion of the transverse part of the vector potential are

$$
\begin{align*}
\hat{H} & =\int d^{3} p \frac{|\vec{p}|}{2} \sum_{\alpha=1,2}\left[\hat{a}^{\dagger}(\vec{p}, \alpha) \hat{a}(\vec{p}, \alpha)+\hat{a}(\vec{p}, \alpha) \hat{a}^{\dagger}(\vec{p}, \alpha)\right] \\
A_{j}^{T}(\vec{x}) & =\int \frac{d^{3} p}{\sqrt{(2 \pi)^{3} 2|\vec{p}|}} \sum_{\alpha=1,2} \epsilon_{\alpha}^{j}(\vec{p})\left[\hat{a}(\vec{p}, \alpha) e^{i \vec{p} \cdot \vec{x}}+\hat{a}^{\dagger}(\vec{p}, \alpha) e^{-i \vec{p} \cdot \vec{x}}\right] \tag{59}
\end{align*}
$$

We recognize these expressions to be the same ones that we obtained before in the Coulomb gauge (except for the normalization factors).

It is instructive to derive the wave functional for the ground state. The ground state $|0\rangle$ is the state annihilated by all the oscillators $\hat{a}(\vec{p}, \alpha)$. Hence its wave function $\Psi_{0}[A]$ satisfies

$$
\begin{equation*}
\left\langle\left\{A_{j}(\vec{x})\right\}\right| \hat{a}(\vec{p}, \alpha)|0\rangle=0 \tag{60}
\end{equation*}
$$

This equation is the functional differential equation

$$
\begin{equation*}
\sum_{\alpha}^{j}(\vec{p})\left[\frac{\delta}{\delta A_{j}^{T}(-\vec{p})}+|\vec{p}| A_{j}^{T}(\vec{p})\right] \Psi_{0}\left(\left\{A_{j}^{T}(\vec{p})\right\}\right)=0 \tag{61}
\end{equation*}
$$

It is easy to check that the unique solution of this equation is

$$
\begin{equation*}
\Psi_{0}[A]=N \exp \left[-\frac{1}{2} \int d^{3} p|\vec{p}| A_{j}^{T}(\vec{p}) A_{j}^{T}(-\vec{p})\right] \tag{62}
\end{equation*}
$$

Since the transverse components of $A_{j}(\vec{p})$ satisfy

$$
\begin{equation*}
A_{j}^{T}(\vec{p})=\epsilon_{j k \ell} \frac{p_{k} A_{\ell}(\vec{p}}{|\vec{p}|}=\left(\frac{\vec{p} \times \vec{A}(\vec{p})}{|\vec{p}|}\right)_{j} \tag{63}
\end{equation*}
$$

we can write $\Psi_{0}[A]$ in the form

$$
\begin{equation*}
\Psi_{0}[A]=\mathcal{N} \exp \left[-\frac{1}{2} \int \frac{d^{3} p}{|\vec{p}|}(\vec{p} \times \vec{A}(\vec{p})) \cdot(\vec{p} \times \vec{A}(-\vec{p}))\right] \tag{64}
\end{equation*}
$$

It is instructive to write this wave function in position space, i.e., as a functional of the configuration of magnetic fields $\{\vec{B}(\vec{x})\}$. Clearly, we have

$$
\begin{align*}
\vec{p} \times \vec{A}(\vec{p}) & =-i \int \frac{d^{3} x}{(2 \pi)^{3 / 2}}\left(\vec{\nabla}_{x} \times \vec{A}(\vec{x})\right) e^{-i \vec{p} \cdot \vec{x}} \\
\vec{p} \times \vec{A}(-\vec{p}) & =i \int \frac{d^{3} x}{(2 \pi)^{3 / 2}}\left(\vec{\nabla}_{x} \times \vec{A}(\vec{x})\right) e^{i \vec{p} \cdot \vec{x}} \tag{65}
\end{align*}
$$

By substitution of these identities back into the exponent of the wave function, we get

$$
\begin{equation*}
\Psi_{0}[A]=\mathcal{N} \exp \left\{-\frac{1}{2} \int d^{3} x \int d^{3} x^{\prime} \vec{B}(\vec{x}) \cdot \vec{B}\left(\vec{x}^{\prime}\right) G\left(\vec{x}, \vec{x}^{\prime}\right)\right\} \tag{66}
\end{equation*}
$$

where $G\left(\vec{x}, \vec{x}^{\prime}\right)$ is given by

$$
\begin{equation*}
G\left(\vec{x}, \vec{x}^{\prime}\right)=\int \frac{d^{3} p}{(2 \pi)^{3}} \frac{e^{-i \vec{p} \cdot\left(\vec{x}-\vec{x}^{\prime}\right)}}{|\vec{p}|} \tag{67}
\end{equation*}
$$

This function has a singular behavior at large values of $|\vec{p}|$. We will define a smoothed version $G_{\Lambda}\left(\vec{x}, \vec{x}^{\prime}\right)$ to be

$$
\begin{equation*}
G_{\Lambda}\left(\vec{x}, \vec{x}^{\prime}\right)=\int \frac{d^{3} p}{(2 \pi)^{3}} \frac{e^{-i \vec{p} \cdot\left(\vec{x}-\vec{x}^{\prime}\right)}}{|\vec{p}|} e^{-|\vec{p}| / \Lambda} \tag{68}
\end{equation*}
$$

which cuts off the contributions with $|\vec{p}| \gg \Lambda$. Also, $G_{\Lambda}\left(\vec{x}, \vec{x}^{\prime}\right)$ formally goes back to $G\left(\vec{x}, \vec{x}^{\prime}\right)$ as $\Lambda \rightarrow \infty . G_{\Lambda}\left(\vec{x}, \vec{x}^{\prime}\right)$ can be evaluated explicitly to give

$$
\left.\begin{array}{rl}
G_{\Lambda}\left(\vec{x}, \vec{x}^{\prime}\right) & =\frac{1}{2 \pi^{2}\left|\vec{x}-\vec{x}^{\prime}\right|^{2}} \int_{0}^{\infty} d t \sin t e^{-t / \Lambda\left|\vec{x}-\vec{x}^{\prime}\right|} \\
& =\frac{1}{2 \pi^{2}\left|\vec{x}-\vec{x}^{\prime}\right|^{2}} \Im\left[\frac{1}{\Lambda\left|\vec{x}-\vec{x}^{\prime}\right|}-i\right. \tag{69}
\end{array}\right]
$$

Thus,

$$
\begin{equation*}
\lim _{\Lambda \rightarrow \infty} G_{\Lambda}\left(\vec{x}, \vec{x}^{\prime}\right)=\frac{1}{2 \pi^{2}\left|\vec{x}-\vec{x}^{\prime}\right|^{2}} \tag{70}
\end{equation*}
$$

Hence, the ground state wave functional $\Psi_{0}[A]$ is

$$
\begin{equation*}
\Psi_{0}[A]=\mathcal{N} \exp \left\{-\frac{1}{4 \pi^{2}} \int d^{3} x \int d^{3} x \frac{\vec{B}(\vec{x}) \cdot \vec{B}\left(\vec{x}^{\prime}\right)}{\left|\vec{x}-\vec{x}^{\prime}\right|^{2}}\right\} \tag{71}
\end{equation*}
$$

which is only a functional of the configuration of magnetic fields.

### 9.4 Path Integral Quantization of Gauge Theories

We have discuss at length the quantization of the abelian gauge theory (i.e., Maxwell's electromagnetism) within canonical quantization in the $A_{0}=0$ and in the Coulomb gauges. Conceptually what we have done is perfectly correct although it poses a number of problems.

1. The canonical formalism is natural in the gauge $A_{0}=0$ and it can be generalized to other gauge theories. However, this gauge is highly noncovariant and it is necessary to prove covariance of physical observables at the end. In addition the gauge field propagator in this gauge is very complicated.
2. The particle spectrum is most transparent in the transverse (or Coulomb) gauge. However, in addition of being non-covariant, it is not possible to generalize this gauge to non-Abelian theories due to subtle topological problems (they will be discussed in Physics 583). The propagator is equally awful in this gauge. The commutation relations in real space look quite different from those in scalar field theory.
3. In non-Abelian theories, even in the absence of matter fields, the theory is already non-linear and needs to be regularized in a manner that gauge invariance is preserved.
4. Although it is possible to use covariant gauges, such as the Lorentz gauge $\partial_{\mu} A^{\mu}=0$, the quantization of the theory is these gauges requires a laborious approach (known as Gupta-Bleuer) of difficult generalization.

At the root of this problems is the issue of quantizing a theory which has a local (or gauge) symmetry in a manner that both Lorentz and gauge invariance are kept explicitly. It turns out that path-integral quantization is the most direct approach to deal with these problems.

Let us construct the path integral for the free electromagnetic field. However, formally the procedure that we will use can be applied to any gauge theory. We will begin with the theory quantized canonically in the gauge $A_{0}=0$.

We saw above that, in the gauge $A_{0}=0$, the electric field $\vec{E}$ is (minus) the momentum canonically conjugate to $\vec{A}$, the spatial components of the gauge field, and obey the equal-time canonical commutation relations

$$
\begin{equation*}
\left[E_{j}(\vec{x}), A_{k}\left(\vec{x}^{\prime}\right)\right]=i \delta^{3}\left(\vec{x}-\vec{x}^{\prime}\right) \tag{72}
\end{equation*}
$$

In addition, in this gauge Gauss' Law becomes a constraint on the space of states, i.e.,

$$
\begin{equation*}
\left.\vec{\nabla} \cdot \vec{E}(\vec{x}) \mid \text { Phys }\rangle=J_{0}(\vec{x}) \mid \text { Phys }\right\rangle \tag{73}
\end{equation*}
$$

which defines the physical Hilbert space. Here $J_{0}(x)$ is a charge density distribution. In the presence of a set of conserved sources $J_{\mu}(x)\left(\right.$ i.e., $\left.\partial_{\mu} J^{\mu}=0\right)$ the Hamiltonian of the free field theory is

$$
\begin{equation*}
\hat{H}=\int d^{3} x \frac{1}{2}\left(\vec{E}^{2}+\vec{B}^{2}\right)+\int d^{3} x \vec{J} \cdot \vec{A} \tag{74}
\end{equation*}
$$

We will construct the path-integral in this space.
Let su denote by $Z\left[J_{\mu}\right]$ the quantity

$$
\begin{equation*}
Z[J]=\operatorname{tr}^{\prime} T e^{-i \int d x_{0} \hat{H}} \equiv \operatorname{tr}\left(T e^{-i \int d x_{0} \hat{H}} \hat{P}\right) \tag{75}
\end{equation*}
$$

where $\mathrm{tr}^{\prime}$ means a trace (or sum) over the space of states that satisfy the constraint of Gauss' Law. We implement this constraint by means of the operator
$\hat{P}$ which projects onto these states

$$
\begin{equation*}
\hat{P}=\prod_{\vec{x}} \delta\left(\vec{\nabla} \cdot \vec{E}(x)-J_{0}(x)\right) \tag{76}
\end{equation*}
$$

We will now follow the standard construction of the path integral but making sure that we only sum over histories that are consistent with the constraint. In principle all we need to do is to insert complete sets of states which are eigenstates of the field operator $\vec{A}(x)$ at all intermediate times. These states, denoted by $\left|\left\{\vec{A}\left(\vec{x}, x_{0}\right)\right\}\right\rangle$, are not gauge invariant (i.e., they do not satisfy the constraint). However, the projection operator $\hat{P}$ weeds out the unphysical components of these states. Hence, if the projection operator is included in the evolution operator, the inserted states actually are gauge-invariant. Thus, to insert at every intermediate time $x_{0}^{k}\left(k=1, \ldots, N\right.$ with $N \rightarrow \infty$ and $\left.\Delta x_{0} \rightarrow 0\right)$ a complete set of gauge-invariant eigenstates amounts to writing $Z[J]$ as

$$
\begin{align*}
& Z[J]=\prod_{k=1}^{N} \int \mathcal{D} A_{j}\left(\vec{x}, x_{0}^{k}\right) \\
& \quad\left\langle\left\{A_{j}\left(\vec{x}, x_{0}^{k}\right)\right\}\right|\left(1-i \Delta x_{0} \hat{H}\right) \prod_{\vec{x}} \delta\left(\vec{\nabla} \cdot \vec{E}\left(\vec{x}, x_{0}^{k}\right)-J_{0}\left(\vec{x}, x_{0}^{k}\right)\right)\left|\left\{A_{j}\left(\vec{x}, x_{0}^{k+1}\right)\right\}\right\rangle \tag{77}
\end{align*}
$$

As an operator, the projection operator $\hat{P}$ is naturally spanned by the eigenstates of the electric field operator $\left|\left\{\vec{E}\left(\vec{x}, x_{0}\right)\right\}\right\rangle$, i.e.,

$$
\begin{align*}
& \prod_{\vec{x}} \delta\left(\vec{\nabla} \cdot \vec{E}\left(\vec{x}, x_{0}\right)-J_{0}\left(\vec{x}, x_{0}\right)\right) \equiv \\
& \int \mathcal{D} \vec{E}\left(\vec{x}, x_{0}\right)\left|\left\{\vec{E}\left(\vec{x}, x_{0}\right)\right\}\right\rangle\left\langle\left\{\vec{E}\left(\vec{x}, x_{0}\right)\right\}\right| \prod_{\vec{x}} \delta\left(\vec{\nabla} \cdot \vec{E}\left(\vec{x}, x_{0}\right)-J_{0}\left(\vec{x}, x_{0}\right)\right) \tag{78}
\end{align*}
$$

The delta function has the integral representation

$$
\begin{align*}
& \prod_{\vec{x}} \delta\left(\vec{\nabla} \cdot \vec{E}\left(\vec{x}, x_{0}\right)-J_{0}\left(\vec{x}, x_{0}\right)\right)= \\
& =\mathcal{N} \int \mathcal{D} A_{0}\left(\vec{x}, x_{0}\right) e^{i \Delta x_{0} \int d^{3} x A_{0}\left(\vec{x}, x_{0}\right)\left(\vec{\nabla} \cdot \vec{E}\left(\vec{x}, x_{0}\right)-J_{0}\left(\vec{x}, x_{0}\right)\right)} \tag{79}
\end{align*}
$$

Hence, the matrix elements of interest become

$$
\begin{array}{r}
\int \mathcal{D} \vec{A} \prod_{x_{0}}\left\langle\left\{\vec{A}\left(\vec{x}, x_{0}\right)\right\}\right|\left(1-i \Delta x_{0} \hat{H}\right) \prod_{\vec{x}} \delta\left(\nabla_{j} \hat{E}_{j}-J_{0}\right)\left|\left\{\vec{A}\left(\vec{x}, x_{0}+\Delta x_{0}\right)\right\}\right\rangle \\
=\int \mathcal{D} A_{0} \mathcal{D} \overrightarrow{A D} \vec{E} \prod_{x_{0}}\left\langle\left\{\vec{A}\left(\vec{x}, x_{0}\right)\right\} \mid\left\{\vec{E}\left(\vec{x}, x_{0}\right)\right\}\right\rangle\left\langle\left\{\vec{E}\left(\vec{x}, x_{0}\right)\right\} \mid\left\{\vec{A}\left(\vec{x}, x_{0}+\Delta x_{0}\right)\right\}\right\rangle \\
\times e^{i \Delta x_{0}\left[\int d^{3} x A_{0}\left(\vec{x}, x_{0}\right)\left(\vec{\nabla} \cdot \vec{E}\left(\vec{x}, x_{0}\right)-J_{0}\left(\vec{x}, x_{0}\right)\right)-\frac{\left\langle\left\{\vec{A}\left(\vec{x}, x_{0}\right)\right\}\right| \hat{H}\left|\left\{\vec{E}\left(\vec{x}, x_{0}\right)\right\}\right\rangle}{\left\langle\left\{\vec{A}\left(\vec{x}, x_{0}\right)\right\} \mid\left\{\vec{E}\left(\vec{x}, x_{0}\right)\right\}\right\rangle}\right]} \tag{80}
\end{array}
$$

The overlaps are equal to

$$
\begin{equation*}
\left\langle\left\{\vec{A}\left(\vec{x}, x_{0}\right)\right\} \mid\left\{\vec{E}\left(\vec{x}, x_{0}\right)\right\}\right\rangle=e^{i \int d^{3} x \vec{A}\left(\vec{x}, x_{0}\right) \cdot \vec{E}\left(\vec{x}, x_{0}\right)} \tag{81}
\end{equation*}
$$

Hence, we find that the product of the overlaps is given by

$$
\begin{align*}
\prod_{x_{0}} \quad & \left\langle\left\{\vec{A}\left(\vec{x}, x_{0}\right)\right\} \mid\left\{\vec{E}\left(\vec{x}, x_{0}\right)\right\}\right\rangle\left\langle\left\{\vec{E}\left(\vec{x}, x_{0}\right)\right\} \mid\left\{\vec{A}\left(\vec{x}, x_{0}+\Delta x_{0}\right)\right\}\right\rangle= \\
& =e^{-i \int d x_{0} \int d^{3} x \vec{E}\left(\vec{x}, x_{0}\right) \cdot \partial_{0} \vec{A}\left(\vec{x}, x_{0}\right)} \tag{82}
\end{align*}
$$

The matrix elements of the Hamiltonian are

$$
\begin{equation*}
\frac{\left\langle\left\{\vec{A}\left(\vec{x}, x_{0}\right)\right\}\right| \hat{H}\left|\left\{\vec{E}\left(\vec{x}, x_{0}\right)\right\}\right\rangle}{\left\langle\left\{\vec{A}\left(\vec{x}, x_{0}\right)\right\} \mid\left\{\vec{E}\left(\vec{x}, x_{0}\right)\right\}\right\rangle}=\int d^{3} x\left[\frac{1}{2}\left(\vec{E}^{2}+\vec{B}^{2}\right)+\vec{J} \cdot \vec{A}\right] \tag{83}
\end{equation*}
$$

Putting everything together we find that the path integral expression for $Z[J]$ has the form

$$
\begin{equation*}
Z[J]=\int \mathcal{D} A_{\mu} \mathcal{D} \vec{E} e^{i S\left[A_{\mu}, \vec{E}\right]} \tag{84}
\end{equation*}
$$

where

$$
\begin{equation*}
\mathcal{D} A_{\mu}=\mathcal{D} \overrightarrow{A D} A_{0} \tag{85}
\end{equation*}
$$

and the action $S\left[A_{\mu}, \vec{E}\right]$ is given by

$$
\begin{equation*}
S\left[A_{\mu}, \vec{E}\right]=\int d^{4} x\left[-\vec{E} \cdot \partial_{0} \vec{A}-\frac{1}{2}\left(\vec{E}^{2}+\vec{B}^{2}\right)-\vec{J} \cdot \vec{A}+A_{0}\left(\vec{\nabla} \cdot \vec{E}-J_{0}\right)\right] \tag{86}
\end{equation*}
$$

Notice that the Lagrange multiplier field $A_{0}$, which appeared when we introduced the integral representation of the delta function, has become the time component of the vector potential (that is the reason why I called it $A_{0}$ ).

Since the action is quadratic in the electric fields, we can integrate them out explicitly to find

$$
\begin{align*}
& \int \mathcal{D} \vec{E} e^{i \int d^{4} x\left(-\frac{1}{2} \vec{E}^{2}+\vec{E} \cdot\left(\partial_{0} \vec{A}-\vec{\nabla} A_{0}\right)\right)}= \\
& \text { const. } e^{i \int d^{4} x \frac{1}{2}\left(\partial_{0} \vec{A}-\vec{\nabla} A_{0}\right)^{2}} \tag{87}
\end{align*}
$$

We now collect everything and find that the path integral is

$$
\begin{equation*}
Z[J]=\int \mathcal{D} A_{\mu} e^{i \int d^{4} x \mathcal{L}} \tag{88}
\end{equation*}
$$

where the Lagrangian is

$$
\begin{equation*}
\mathcal{L}=-\frac{1}{4} F_{\mu \nu} F^{\mu \nu}+J_{\mu} A^{\mu} \tag{89}
\end{equation*}
$$

which is what we should have expected. We should note here that this results is valid for all Gauge Theories, Abelian or non-Abelian. In other words, the path integral is always the sum over the histories of the field $A_{\mu}$ with a weight factor which is the exponential of $i / \hbar$ times the action $S$ of the Gauge Theory.

Therefore we found that, at least formally, we can write a functional integral which will play the role of the generating functional of the $N$-point functions of this theories,

$$
\begin{equation*}
-i\langle 0| T A_{\mu_{1}}\left(x_{1}\right) \ldots A_{\mu_{N}}\left(x_{N}\right)|0\rangle \tag{90}
\end{equation*}
$$

### 9.5 Path Integrals and Gauge Fixing

We must emphasize that the expression for the path integral in Eq. (88) is formal because we are summing over all histories of the field without restriction. In fact, since the action $S$ and the integration measure $\mathcal{D} A_{\mu}$ are both gauge invariant, histories that differ by gauge transformations have the same weight and the partition function has an apparent divergence of the form $v(G)^{V}$, where $v(G)$ is the volume of the gauge group $G$ and $V$ is the (infinite) volume of spacetime. In order to avoid this problem we must implement some sort of gauge fixing condition on the sum over histories. We will do so by means of a method introduced by L. Faddeev and V. Popov. Although the method works for all Gauge Theories, the non-Abelian theories have subtleties and technical issues that we will discus below. We will begin with a general discussion of the method and then we will specialize it for the case of Maxwell's theory, the $U(1)$ gauge theory without matter fields.

Let the vector potential $A_{\mu}$ be a field which takes values in the algebra of a gauge group $G$, i.e., $A_{\mu}$ is a linear combination of the group generators, and let $U(x)$ be an unitary-matrix field that takes values on a representation of the
group $G$ (please recall our earlier discussion on this subject). For the Abelian group $U(1)$, we have

$$
\begin{equation*}
U(x)=e^{i \phi(x)} \tag{91}
\end{equation*}
$$

where $\phi(x)$ is a real (scalar) field. A gauge transformation is, for a group $G$

$$
\begin{equation*}
A_{\mu}^{U}=U A_{\mu} U^{\dagger}-i U \partial_{\mu} U^{\dagger} \tag{92}
\end{equation*}
$$

For the Abelian group $U(1)$ we have

$$
\begin{equation*}
A_{\mu}^{U}=A_{\mu}+\partial_{\mu} \phi \tag{93}
\end{equation*}
$$

In order to avoid infinities in $Z[J]$ we must impose restrictions on the sum over histories such that histories that are related via a gauge transformation are counted exactly once. In order to do that we must find a way to classify the vector potentials in classes. We will do this by defining gauge fixing conditions. Each class is labelled by a representative configuration and other elements in the class are related to it by smooth gauge transformations. Hence, all configurations in a given class are characterized by a set of gauge invariant data (such as field strengths in the Abelian theory). We must choose gauge conditions such that the theory remains local and, if possible, Lorentz covariant. It is essential that, whatever gauge condition we use that each class is counted exactly once by the gauge condition. It turns out that for the Abelian theory this is trivially the case but in non-Abelian theories there are many gauges (such as the Coulomb gauge) in which, for topological reasons, a class may be counted more than once. (This question is known as the Gribov problem.) Finally we must also keep in mind that we are only fixing the local gauge invariance but we should not alter the boundary conditions since they represent physical degrees of freedom.


Figure 1: The gauge fixing condition selects a manifold of configurations.

How do we impose a gauge condition consistently? We will do it in the following way. Let us denote by

$$
\begin{equation*}
g\left(A_{\mu}\right)=0 \tag{94}
\end{equation*}
$$

the gauge condition we wish to impose, where $g\left(A_{\mu}\right)$ is a local differentiable function of the gauge fields and/or their derivatives. e. g. $g\left(A_{\mu}\right)=\partial_{\mu} A^{\mu}$ for the Lorentz gauge or $g\left(A_{\mu}\right)=n_{\mu} A^{\mu}$ for an axial gauge. Note that the discussion that follows is valid for all compact Lie groups $G$ of volume $v(G)$. For the special case of Maxwell's gauge theory, the gauge group is $U(1)$. Up to topological considerations, the group $U(1)$ is isomorphic to the real numbers $\mathbb{R}$, even though $v(U(1))=2 \pi$ and $v(\mathbb{R})=\infty$.

Naively, to impose a gauge condition would mean to restrict the path integral by inserting Eq. (94) as a delta function in the integrand,

$$
\begin{equation*}
Z[J] \sim \int \mathcal{D} A_{\mu} \delta\left(g\left(A_{\mu}\right)\right) e^{i S[A, J]} \tag{95}
\end{equation*}
$$

We will see below that in general this is an inconsistent (and wrong) prescription.
Following Faddeev and Popov we begin by considering the following integral

$$
\begin{equation*}
\Delta_{g}^{-1}\left[A_{\mu}\right] \equiv \int \mathcal{D} U \delta\left(g\left(A_{\mu}^{U}\right)\right) \tag{96}
\end{equation*}
$$

where $A_{\mu}^{U}(x)$ are the configurations of gauge fields related by the gauge transformation $U(x)$ to the configuration $A_{\mu}(x)$, i.e., we move inside one class.

Let us show that $\Delta_{g}^{-1}\left[A_{\mu}\right]$ is gauge invariant. We now observe that the integration measure $\mathcal{D} U$, usually called the Haar measure, is invariant under the composition rule $U \rightarrow U U^{\prime}$,

$$
\begin{equation*}
\mathcal{D} U=\mathcal{D}\left(U U^{\prime}\right) \tag{97}
\end{equation*}
$$

where $U^{\prime}$ is and arbitrary but fixed element of $G$. For the case of $G=U(1)$, $U=\exp (i \phi)$ and $\mathcal{D} U \equiv \mathcal{D} \phi$.

Using the invariance of the measure, Eq. (97) we can write

$$
\begin{equation*}
\Delta_{g}^{-1}\left[A_{\mu}^{U^{\prime}}\right]=\int \mathcal{D} U \delta\left(g\left(A_{\mu}^{U^{\prime} U}\right)\right)=\int \mathcal{D} U^{\prime \prime} \delta\left(g\left(A_{\mu}^{U^{\prime \prime}}\right)\right)=\Delta_{g}^{-1}\left[A_{\mu}\right] \tag{98}
\end{equation*}
$$

where we have set $U^{\prime} U=U^{\prime \prime}$. Therefore $\Delta_{g}^{-1}\left[A_{\mu}\right]$ is gauge invariant, i.e., it is a function of the class and not of the configuration $A_{\mu}$ itself. Obviously we can also write Eq. (96) in the form

$$
\begin{equation*}
1=\Delta_{g}\left[A_{\mu}\right] \int \mathcal{D} U \delta\left(g\left(A_{\mu}^{U}\right)\right) \tag{99}
\end{equation*}
$$

We will now insert the number 1, as given by Eq. (99), in the path integral for a general Gauge Theory and find

$$
\begin{align*}
Z[J] & =\int \mathcal{D} A_{\mu} \times 1 \times e^{i S[A, J]} \\
& =\int \mathcal{D} A_{\mu} \Delta_{g}\left[A_{\mu}\right] \int \mathcal{D} U \delta\left(g\left(A_{\mu}^{U}\right)\right) e^{i S[A, J]} \tag{100}
\end{align*}
$$

We now make the change of variables

$$
\begin{equation*}
A_{\mu} \rightarrow A_{\mu}^{V} \tag{101}
\end{equation*}
$$

where $V=V(x)$ is an arbitrary gauge transformation, and find

$$
\begin{equation*}
Z[J]=\iint \mathcal{D} U \mathcal{D} A_{\mu}^{V} e^{i S\left[A^{V}, J\right]} \Delta_{g}\left[A_{\mu}^{V}\right] \delta\left(g\left(A_{\mu}^{V U}\right)\right) \tag{102}
\end{equation*}
$$

(Notice that we have changed the order of integration.) We now choose $V=$ $U^{-1}$, and use the gauge invariance of the action $S[A, J]$, of the measure $\mathcal{D} A_{\mu}$ and of $\Delta_{g}[A]$ to write the partition as

$$
\begin{equation*}
Z[J]=\left[\int \mathcal{D} U\right] \int \mathcal{D} A_{\mu} \Delta_{g}\left[A_{\mu}\right] \delta\left(g\left(A_{\mu}\right)\right) e^{i S[A, J]} \tag{103}
\end{equation*}
$$

The factor in brackets in Eq. (103) is the infinite constant

$$
\begin{equation*}
\int \mathcal{D} U=v(G)^{V} \tag{104}
\end{equation*}
$$

where $v(G)$ is the volume of the gauge group and $V$ is the (infinite) volume of space-time. This infinite constant is nothing but the result of summing over gauge-equivalent states.

Thus, provided the quantity $\Delta_{g}\left[A_{\mu}\right]$ is finite and it does not vanish identically, we find that the consistent rule for fixing the gauge consists in dividing out the (infinite) factor of the volume of the group element but, more importantly, to insert together with the constraint $\delta\left(g\left(A_{\mu}\right)\right)$ the factor $\Delta_{g}\left[A_{\mu}\right]$ in the integrand of $Z[J]$,

$$
\begin{equation*}
Z[J] \sim \int \mathcal{D} A_{\mu} \Delta_{g}\left[A_{\mu}\right] \delta\left(g\left(A_{\mu}\right)\right) e^{i S[A, J]} \tag{105}
\end{equation*}
$$

We are only left to compute $\Delta_{g}\left[A_{\mu}\right]$. We will show now that $\Delta_{g}\left[A_{\mu}\right]$ is a determinant of a certain operator. The quantity $\Delta_{g}\left[A_{\mu}\right]$ is known as the Faddeev-Popov determinant. We will only compute first this determinant for the case of the Abelian theory $U(1)$. Below we will also discuss the non-Abelian case, relevant for Yang-Mills gauge theories.

We will compute $\Delta_{g}\left[A_{\mu}\right]$ by using the fact that $g\left[A_{\mu}^{U}\right]$ can be regarded as a function of $U(x)$ (for $A_{\mu}(x)$ fixed). We will now change variables from $U$ to $g$. The price we pay is a Jacobian factor since

$$
\begin{equation*}
\mathcal{D} U=\mathcal{D} g \operatorname{Det}\left|\frac{\delta U}{\delta g}\right| \tag{106}
\end{equation*}
$$

where the determinant is the Jacobian of the change of variables. Since this is a non-linear change of variables, we expect a non-trivial Jacobian. Therefore we can write

$$
\begin{equation*}
\Delta_{g}^{-1}\left[A_{\mu}\right]=\int \mathcal{D} U \delta\left(g\left(A_{\mu}^{U}\right)\right)=\int \mathcal{D} g \operatorname{Det}\left|\frac{\delta U}{\delta g}\right| \delta(g) \tag{107}
\end{equation*}
$$

and we find

$$
\begin{equation*}
\Delta_{g}^{-1}\left[A_{\mu}\right]=\operatorname{Det}\left|\frac{\delta U}{\delta g}\right|_{g=0} \tag{108}
\end{equation*}
$$

or, conversely

$$
\begin{equation*}
\Delta_{g}\left[A_{\mu}\right]=\operatorname{Det}\left|\frac{\delta g}{\delta U}\right|_{g=0} \tag{109}
\end{equation*}
$$

Thus far all we have done holds for all gauge theories (with a compact gauge group). We will specialize our discussion first for the case of the $U(1)$ gauge theory, Maxwell's electromagnetism. We will discuss how this applies to nonAbelian Yang-Mill gauge theories below. For example, for the particular case of the Abelian $U(1)$ gauge theory, the Lorentz gauge condition is obtained by the choice $g\left(A_{\mu}\right)=\partial_{\mu} A^{\mu}$. Then, for $U(x)=\exp (i \phi(x))$, we get

$$
\begin{equation*}
g\left(A_{\mu}^{U}\right)=\partial_{\mu}\left(A^{\mu}+\partial^{\mu} \phi\right)=\partial_{\mu} A^{\mu}+\partial^{2} \phi \tag{110}
\end{equation*}
$$

Hence,

$$
\begin{equation*}
\frac{\delta g(x)}{\delta \phi(y)}=\partial^{2} \delta(x-y) \tag{111}
\end{equation*}
$$

Thus, for the Lorentz gauge of the Abelian theory, the Faddeev-Popov determinant is given by

$$
\begin{equation*}
\Delta_{g}\left[A_{\mu}\right]=\operatorname{Det} \partial^{2} \tag{112}
\end{equation*}
$$

which is a constant independent of $A_{\mu}$. This is a peculiarity of the Abelian theory and, as we will see below, it is not true in the non-Abelian case.

Let us return momentarily to the general case of Eq. (105), and modify the gauge condition from $g\left(A_{\mu}\right)=0$ to $g\left(A_{\mu}\right)=c(x)$, where $c(x)$ is some arbitrary function of $x$. The partition function now reads

$$
\begin{equation*}
Z[J] \sim \int \mathcal{D} A_{\mu} \Delta_{g}\left[A_{\mu}\right] \delta\left(g\left(A_{\mu}\right)-c(x)\right) e^{i S[A, J]} \tag{113}
\end{equation*}
$$

We will now average over the arbitrary functions with a Gaussian weight (properly normalized to unity)

$$
\begin{align*}
Z_{\alpha}[J] & =\mathcal{N} \int \mathcal{D} A_{\mu} \mathcal{D} c e^{-i \int d^{4} x \frac{c(x)^{2}}{2 \alpha}} \Delta_{g}\left[A_{\mu}\right] \delta\left(g\left(A_{\mu}\right)-c(x)\right) e^{i S[A, J]} \\
& =\mathcal{N} \int \mathcal{D} A_{\mu} \Delta_{g}\left(A_{\mu}\right) e^{+i \int d^{4} x}\left[\mathcal{L}[A, J]-\frac{1}{2 \alpha}\left(g\left(A_{\mu}\right)\right)^{2}\right] \tag{114}
\end{align*}
$$

From now on we will restrict our discussion to the $U(1)$ Abelian gauge theory (the electromagnetic field) and $g\left(A_{\mu}\right)=\partial_{\mu} A^{\mu}$. From Eq. (114) we find that in this gauge the Lagrangian is

$$
\begin{equation*}
\mathcal{L}_{\alpha}=-\frac{1}{4} F_{\mu \nu}^{2}-J_{\mu} A^{\mu}-\frac{1}{2 \alpha}\left(\partial_{\mu} A^{\mu}\right)^{2} \tag{115}
\end{equation*}
$$

The parameter $\alpha$ labels a family of gauge fixing conditions known as the Feynman' $t$ Hooft gauges. For $\alpha \rightarrow 0$ we recover the strong constraint $\partial_{\mu} A^{\mu}=0$, the Lorentz gauge. From the point of view of doing calculations the simplest is the gauge $\alpha=1$ (the Feynman gauge) as we will see now. After some algebra is straightforward to see that, up to surface terms, the Lagrangian is equal to

$$
\begin{equation*}
\mathcal{L}_{\alpha}=\frac{1}{2} A_{\mu}\left[g^{\mu \nu} \partial^{2}-\frac{\alpha-1}{\alpha} \partial^{\mu} \partial^{\nu}\right] A_{\nu}-J_{\mu} A^{\mu} \tag{116}
\end{equation*}
$$

and the partition function reduces to

$$
\begin{equation*}
Z[J]=\mathcal{N} \operatorname{Det}\left[\partial^{2}\right] \int \mathcal{D} A_{\mu} e^{i \int d^{4} x \mathcal{L}_{\alpha}[A, J]} \tag{117}
\end{equation*}
$$

Hence, in a general gauge labelled by $\alpha$, we get

$$
\begin{align*}
Z[J]= & \mathcal{N} \operatorname{Det}\left[\partial^{2}\right] \operatorname{Det}\left[g^{\mu \nu} \partial^{2}-\frac{\alpha-1}{\alpha} \partial^{\mu} \partial^{\nu}\right]^{-1 / 2} \\
& \times e^{-\frac{i}{2} \int d^{4} x \int d^{4} y J_{\mu}(x) G^{\mu \nu}(x-y) J_{\nu}(y)} \tag{118}
\end{align*}
$$

where $G^{\mu \nu}(x-y)$ is the propagator in that gauge.
The form of Eq. (118) may seem to imply that $Z[J]$ is gauge dependent. This cannot be correct since the path integral is by construction gauge-invariant. We will show in the next subsection that gauge invariance is indeed protected. This result comes about because $J_{\mu}$ is a conserved current, and as such it satisfies the continuity equation $\partial_{\mu} J^{\mu}=0$. For this family of gauges, the propagator takes the form

$$
\begin{equation*}
G_{\mu \nu}(x-y)=\left[g^{\mu \nu}+\frac{\alpha-1}{\alpha} \frac{\partial^{\mu} \partial^{\nu}}{\partial^{2}}\right] G(x-y) \tag{119}
\end{equation*}
$$

where $G(x-y)$ is the propagator of the scalar field.
Thus, as expected for a free field theory, $Z[J]$ is a product of two factors: a functional (or fluctuation) determinant, and a factor that depends solely on the sources $J_{\mu}$ which contains all the information on the correlation functions. For the case of a single scalar field we also found a contribution in the form of a determinant factor but its power was $-1 / 2$. Here there are two such factors. The first one is the Faddeev-Popov determinant. The second one is the determinant of the fluctuation operator for the gauge field. However, in the Feynman gauge, $\alpha=1$, this operator is just $g^{\mu \nu} \partial^{2}$, and its determinant has the same form as the Faddeev-Popov determinant except that it has a power $-4 / 2$. This is what one would have expected for a theory with four independent fields (one for each component of $A_{\mu}$ ). The Faddeev-Popov determinant has power +1 . Thus the total power is just $1-4 / 2=-1$, which is the correct answer for a theory with only two independent fields.

### 9.6 The Propagator

For general $\alpha, G_{\mu \nu}(x-y)$ is the solution of the Green's function equation

$$
\begin{equation*}
\left[g^{\mu \nu} \partial^{2}-\frac{\alpha-1}{\alpha} \partial^{\mu} \partial^{\nu}\right] G_{\nu \lambda}(x-y)=g_{\lambda}^{\mu} \delta^{4}(x-y) \tag{120}
\end{equation*}
$$

Notice that in the special case of the Feynman gauge, $\alpha=1$, this equation becomes

$$
\begin{equation*}
\partial^{2} G^{\mu \nu}(x-y)=g^{\mu \nu} \delta^{4}(x-y) \tag{121}
\end{equation*}
$$

Hence, in the Feynman gauge, $G_{\mu \nu}(x-y)$ takes the form

$$
\begin{equation*}
G^{\mu \nu}(x-y)=g^{\mu \nu} G(x-y) \tag{122}
\end{equation*}
$$

where $G(x-y)$ is just the propagator of a free scalar field, i.e.,

$$
\begin{equation*}
\partial^{2} G(x-y)=\delta^{4}(x-y) \tag{123}
\end{equation*}
$$

However, in a general gauge the propagator

$$
\begin{equation*}
G_{\mu \nu}(x-y)=-i\langle 0| T A_{\mu}(x) A_{\nu}(y)|0\rangle \tag{124}
\end{equation*}
$$

does not coincide with the propagator of a scalar field. Therefore, $G_{\mu \nu}(x-y)$, as expected, is a gauge dependent quantity. In spite of that it does contain physical information. Let us examine this issue by calculating the propagator in a general gauge $\alpha$.

The Fourier transform of $G_{\mu \nu}(x-y)$ in $D$ space-time dimensions is

$$
\begin{equation*}
G_{\mu \nu}(x-y)=\int \frac{d^{D} p}{(2 \pi)^{D}} \widetilde{G}_{\mu \nu}(p) e^{i p \cdot x} \tag{125}
\end{equation*}
$$

This a solution of Eq. (120) provided $\widetilde{G}_{\mu \nu}(p)$ satisfies

$$
\begin{equation*}
\left[-g^{\mu \nu} p^{2}+\frac{\alpha-1}{\alpha} p^{\mu} p^{\nu}\right] \widetilde{G}_{\nu \lambda}(p)=g_{\lambda}^{\mu} \tag{126}
\end{equation*}
$$

The formal solution is

$$
\begin{equation*}
\widetilde{G}_{\mu \nu}(p)=-\frac{1}{p^{2}}\left[g^{\mu \nu}+\frac{\alpha-1}{\alpha} \frac{p^{\mu} p^{\nu}}{p^{2}}\right] \tag{127}
\end{equation*}
$$

In space-time the form of this (still formal) solution is given by Eq. (119).
In particular, in the Feynman gauge $\alpha=1$, we get

$$
\begin{equation*}
\widetilde{G}_{\mu \nu}^{F}(p)=-\frac{1}{p^{2}} g^{\mu \nu} \tag{128}
\end{equation*}
$$

whereas in the Lorentz gauge we find instead

$$
\begin{equation*}
\widetilde{G}_{\mu \nu}^{L}(p)=-\frac{1}{p^{2}}\left[g^{\mu \nu}-\frac{p^{\mu} p^{\nu}}{p^{2}}\right] \tag{129}
\end{equation*}
$$

Hence, in all cases there is a pole in $p^{2}$ in front of the propagator and a matrix structure that depends on the gauge choice. Notice that the matrix in brackets in the Lorentz gauge, known as the transverse projection operator, satisfies

$$
\begin{equation*}
p_{\mu}\left[g^{\mu \nu}-\frac{p^{\mu} p^{\nu}}{p^{2}}\right]=0 \tag{130}
\end{equation*}
$$

which follows from the gauge condition $\partial_{\mu} A^{\mu}=0$.
The physical information of this propagator is condensed in its analytic structure. It has a pole at $p^{2}=0$ which implies that $p_{0}=\sqrt{\vec{p}^{2}}=|\vec{p}|$ is the singularity of $\widetilde{G}_{\mu \nu}(p)$. Hence the pole in the propagator tells us that this theory has a massless particle, the photon.

To actually compute the propagator in space-time from $\widetilde{G}_{\mu \nu}(p)$ requires that we define the integrals in momentum space carefully. As it stands, the Fourier integral Eq. (125) is ill defined due to the pole in $\widetilde{G}_{\mu \nu}(p)$ at $p^{2}=0$. A proper definition requires that we move the pole into the complex plane by shifting $p^{2} \rightarrow p^{2}+i \epsilon$, where $\epsilon$ is real and $\epsilon \rightarrow 0^{+}$. This prescription yields the Feynman propagator. We will see in the next section that this rule applies to any theory and that it always yields the vacuum expectation value of the time ordered product of fields. For the rest of this section we will use the propagator in the Feynman gauge which reduces to the propagator of a scalar field. This is a quantity we know quite well, both in Euclidean and Minkowski space-times.

### 9.7 Physical meaning of $Z[J]$ and the Wilson loop operator

We discussed before that a general property of the path integral of any theory is that, in imaginary time, $Z[0]$ is just

$$
\begin{equation*}
Z[0]=\langle 0 \mid 0\rangle \sim e^{-T E_{0}} \tag{131}
\end{equation*}
$$

where $T$ is the time span (i.e., $\quad T \rightarrow \infty$ ) (watch out, here $T$ is not the temperature!), and $E_{0}$ is the vacuum energy. Thus, if the sources $J_{\mu}$ are static (or quasi-static) we get instead

$$
\begin{equation*}
\frac{Z[J]}{Z[0]} \sim e^{-T\left[E_{0}(J)-E_{0}\right]} \tag{132}
\end{equation*}
$$

Thus, the change in the vacuum energy due to the presence of the sources is

$$
\begin{equation*}
U(J)=E_{0}(J)-E_{0}=\lim _{T \rightarrow \infty} \ln \frac{Z[J]}{Z[0]} \tag{133}
\end{equation*}
$$

As we will see, the behavior of this quantity has a lot of information about the physical properties of the vacuum (i.e., the ground state) of a theory. Quite generally, if the quasi-static sources $J_{\mu}$ are well separated from each other, $U(J)$ can be split into two terms: a self-energy of the sources, and an interaction energy, i.e.,

$$
\begin{equation*}
U(J)=E_{\text {self-energy }}[J]+V_{\text {int }}[J] \tag{134}
\end{equation*}
$$

As an example, we will now compute the expectation value of the Wilson loop operator,

$$
\begin{equation*}
W_{\Gamma}=\langle 0| T e^{i e} \oint_{\Gamma} d x_{\mu} A^{\mu}|0\rangle \tag{135}
\end{equation*}
$$

where $\Gamma$ is the closed path in space-time shown in the figure. Physically, what we are doing is looking at the electromagnetic field created by the current

$$
\begin{equation*}
J_{\mu}(x)=e \delta\left(x_{\mu}-s_{\mu}\right) \hat{s}_{\mu} \tag{136}
\end{equation*}
$$

where $s_{\mu}$ is the set of points of space-time on the loop $\Gamma$, and $\hat{s}_{\mu}$ is a unit vector field tangent to $\Gamma$. The loop $\Gamma$ has time span $T$ and spatial width $R$. We will be interested in loops such that $T \gg R$ so that the sources are turned on adiabatically in the remote past and switched off also adiabatically in the remote future. By current conservation the loop must be oriented. Thus, at a fixed time $x_{0}$ the loop looks like a pair of static sources with charges $\pm e$ at $\pm R / 2$. In other words we are looking of the affects of a particle-antiparticle pair which is created at rest in the remote past, the members of the pair are slowly separated (to avoid bremsstrahlung radiation) and live happily apart from each other, at a prudent distance $R$, for a long time $T$, and are finally (and adiabatically) annihilated in the remote future. Thus, we are in the quasistatic regime described above and $Z[J] / Z[0]$ should tell us what is the effective interaction between this pair of sources ("electrodes").


Figure 2: The Wilson loop operator can be viewed as representing a pair of quasi-static sources of charge $\pm e$ separated a distance $R$ from each other.

What are the possible behaviors of the Wilson loop operator in general (that is for any gauge theory)? The answer to this depends on the nature of the vacuum state. We will se that a given theory may have different vacua or phases (as in thermodynamic phases) and that the behavior of the physical observables is different in different vacua (or phases). Here we will do an explicit computation for the case of Maxwell's $U(1)$ gauge theory. However the behavior that will find only holds for a free field and it is not generic. What are the possible behaviors, then? A loop is an extended object (as opposed to a local field operator) characterized by its geometric properties: its area, perimeter, aspect ratio, and so on. We will show later on that these geometric properties of the loop to characterize the behavior of the Wilson loop operator. Here are the generic cases:

1. Area Law: Let $A=R T$ be the area of the loop. One possible behavior of the Wilson loop operator is the area law

$$
\begin{equation*}
W_{\Gamma} \sim e^{-\sigma R T} \tag{137}
\end{equation*}
$$

w We will show later on that this is the fastest possible decay of the Wilson loop operator as a function of size. The quantity $\sigma$ is known as the string tension. If the area law is obeyed the effective potential for $R$ large (but still small compared to $T$ ) behaves as

$$
\begin{equation*}
V_{\mathrm{int}}(R)=\lim _{T \rightarrow \infty} \frac{-1}{T} \ln W_{\Gamma}=\sigma R \tag{138}
\end{equation*}
$$

Hence, in this case the energy to separate a pair of sources grows linearly with distance and the sources are confined. We will say that in this case the the theory is in confined phase.
2. Perimeter Law: Another possible decay behavior (weaker than the area law) is a perimeter law

$$
\begin{equation*}
W_{\Gamma} \sim e^{-\rho(R+T)}+O\left(e^{-R / \xi}\right) \tag{139}
\end{equation*}
$$

where $\rho$ is a constant with units of energy, and $\xi$ is a length scale. This decay law implies that in this case

$$
\begin{equation*}
V_{\mathrm{int}} \sim \rho+\text { const. } e^{-R / \xi} \tag{140}
\end{equation*}
$$

Thus, in this case the energy to separate two sources is finite. This is a deconfined phase. However since it is massive (with a mass scale $m \sim \xi^{-1}$ ) there are no long range gauge bosons. This is the Higgs phase. It bears a close analogy with a superconductor.
3. Scale Invariant: Yet another possibility is that the Wilson loop behavior is determined by the aspect ratio $R / T$ or $T / R$,

$$
\begin{equation*}
W_{\Gamma} \sim e^{-\alpha\left(\frac{R}{T}+\frac{T}{R}\right)} \tag{141}
\end{equation*}
$$

where $\alpha$ is a dimensionless constant. This behavior leads to an interaction

$$
\begin{equation*}
V_{\mathrm{int}} \sim-\frac{\alpha}{R} \tag{142}
\end{equation*}
$$

which coincides with the Coulomb law in 4 dimensions. We will see that this is a deconfined phase with massless gauge bosons (photons).

We will now compute the expectation value of the Wilson loop operator in Maxwell's $U(1)$ gauge theory. We will return to the general problem when we discuss the strong coupling behavior of gauge theories. We begin by using the analytic continuation of Eq. (118) to imaginary time, i.e.,

$$
\begin{align*}
Z[J] & =\mathcal{N} \operatorname{Det}\left[\partial^{2}\right]^{-1} e^{-\frac{1}{2} \int d^{4} x \int d^{4} y J_{\mu}(x)\left\langle A_{\mu}(x) A_{\nu}(y)\right\rangle J_{\nu}(y)} \\
& =\mathcal{N} \operatorname{Det}\left[\partial^{2}\right]^{-1} e^{-\frac{e^{2}}{2} \oint_{\Gamma} d x_{\mu} \oint_{\Gamma} d y_{\nu}\left\langle A_{\mu}(x) A_{\nu}(y)\right\rangle} \tag{143}
\end{align*}
$$

where $\left\langle A_{\mu}(x) A_{\nu}(y)\right\rangle$ is the Euclidean propagator of the gauge fields in the family of gauges labelled by $\alpha$. In the Feynman gauge $\alpha=1$ the propagator is given by the expression

$$
\begin{equation*}
\left\langle A_{\mu}(x) A_{\nu}(y)\right\rangle=\delta_{\mu \nu} \int \frac{d^{D} p}{(2 \pi)^{D}} \frac{1}{p^{2}} e^{i p_{\mu} \cdot\left(x_{\mu}-y_{\mu}\right)} \tag{144}
\end{equation*}
$$

where $\mu=1, \ldots, D$. After doing the integral we find that the Euclidean propagator (the correlation function) in the Feynman gauge is

$$
\begin{equation*}
\left\langle A_{\mu}(x) A_{\nu}(y)\right\rangle=\delta_{\mu \nu} \frac{\Gamma\left(\frac{D}{2}-1\right)}{4 \pi^{D / 2}|x-y|^{D-2}} \tag{145}
\end{equation*}
$$

Therefore $E[J]-E_{0}$ is equal to

$$
\begin{align*}
E[J]-E_{0} & =\lim _{T \rightarrow \infty} \frac{e^{2}}{2 T} \oint_{\Gamma} \oint_{\Gamma} d \vec{x} \cdot d \vec{x} \frac{\Gamma\left(\frac{D}{2}-1\right)}{4 \pi^{D / 2}|x-y|^{D-2}} \\
=2 \times \text { self }- \text { energy } & -\frac{e^{2}}{2 T} 2 \int_{-T / 2}^{+T / 2} d x_{D} \int_{-T / 2}^{+T / 2} d y_{D} \frac{\Gamma\left(\frac{D}{2}-1\right)}{4 \pi^{D / 2}|x-y|^{D-2}} \tag{146}
\end{align*}
$$

where $|x-y|^{2}=\left(x_{D}-y_{D}\right)^{2}+R^{2}$. The integral in Eq. (146) is equal to

$$
\begin{align*}
& \int_{-T / 2}^{+T / 2} d x_{D} \int_{-T / 2}^{+T / 2} d y_{D} \frac{\Gamma\left(\frac{D}{2}-1\right)}{4 \pi^{D / 2}|x-y|^{D-2}} \\
& =\int_{-T / 2}^{+T / 2} d s \int_{-(T / 2+s) / R}^{+(T / 2-s) / R} \frac{d t}{\left(t^{2}+1\right)^{(D-2) / 2}} \frac{1}{R^{D-3}} \times \frac{\Gamma\left(\frac{D-2}{2}\right)}{4 \pi^{D / 2}} \\
& \approx \frac{1}{R^{D-3}} \int_{-T / 2}^{+T / 2} d s \int_{-\infty}^{+\infty} \frac{d t}{\left(t^{2}+1\right)^{(D-2) / 2}} \times \frac{\Gamma\left(\frac{D-2}{2}\right)}{4 \pi^{D / 2}} \\
& =\frac{T \sqrt{\pi}}{R^{D-3}} \frac{\Gamma\left(\frac{D-3}{D}\right)}{\Gamma\left(\frac{D-2}{2}\right)} \times \frac{\Gamma\left(\frac{D-2}{2}\right)}{4 \pi^{D / 2}} \tag{147}
\end{align*}
$$

where

$$
\begin{equation*}
\Gamma(\nu)=\int_{0}^{\infty} d t t^{\nu-1} e^{-t} \tag{148}
\end{equation*}
$$

is the Euler Gamma function. In Eq. (147) we already took the limit $T / R \rightarrow \infty$. Putting it all together we find that the interaction energy of a pair of static sources of charges $\pm e$ separated a distance $R$ in $D$ dimensional space-time is given by

$$
\begin{equation*}
V_{\mathrm{int}}(R)=-\frac{\Gamma\left(\frac{D-1}{2}\right)}{2 \pi^{(D-1) / 2}(D-3)} \frac{e^{2}}{R^{D-3}} \tag{149}
\end{equation*}
$$

This is the Coulomb potential in $D$ space-time dimensions. In particular, in $D=4$ dimensions we find

$$
\begin{equation*}
V_{\mathrm{int}}(R)=-\left(\frac{e^{2}}{4 \pi}\right) \frac{1}{R} \tag{150}
\end{equation*}
$$

where the quantity $e^{2} / 4 \pi$ is the fine structure constant.
Therefore we find that, even at the quantum level, the effective interaction between a pair of static sources is the Coulomb interaction. This is true because Maxwell's theory is a free field theory. It is also true in Quantum Electrodynamics (QED), the Quantum Field Theory of electrons and photons, at distances $R$ much greater than the Compton wavelength of the electron. However it is not true at short distances where the effective charge is screened by fluctuations of the Dirac field and the potential becomes exponentially suppressed. In contrast, in Quantum Chromodynamics (QCD) the situation is quite different: even in the absence of matter, for $R$ large compared with a scale $\xi$ determined by the dynamics, the effective potential grows linearly with $R$. This effect in known as confinement, and the scale $\xi$ is the confinement scale. Conversely, the potential is Coulomb-like at short distances, a behavior known as asymptotic freedom.

### 9.8 Path-Integral Quantization of Non-Abelian Gauge Theories

Most of what we did for the Abelian case, carries over to non-Abelian gauge theories where, as we will see, it plays a much more central role. In this section we will discuss the general properties of the path-integral quantization of nonAbelian gauge theories, but we will not deal with the non-linearities here.

The path integral $Z[J]$ for a non-Abelian gauge field $A_{\mu}=A_{\mu}^{a} \lambda^{a}$ in the algebra of a simply connected compact Lie group $G$, whose generators are the Hermitian matrices $\lambda^{a}$, with gauge condition(s) $g^{a}[A]$ is

$$
\begin{equation*}
Z[J]=\int \mathcal{D} A_{\mu}^{a} e^{i S[A, J]} \delta(g[A]) \Delta_{\mathrm{FP}}[A] \tag{151}
\end{equation*}
$$

where we will use the family of covariant gauge conditions

$$
\begin{equation*}
g^{a}[A]=\partial^{\mu} A_{\mu}^{a}(x)+c^{a}(x)=0 \tag{152}
\end{equation*}
$$

and where $\Delta_{\mathrm{FP}}[A]$ is the Faddeev-Popov determinant. Notice that we impose one gauge condition for each direction in the algebra of the gauge group $G$. We will proceed as we did in the Abelian case and consider an average over gauges. In other words we will work in the manifestly covariant Feynman-‘t Hooft gauges.

Let us work out the structure of the Faddeev-Popov determinant for a general gauge fixing condition $g^{a}[A]$. Let $U$ be an infinitesimal gauge transformation,

$$
\begin{equation*}
U \simeq 1+i \epsilon^{a}(x) \lambda^{a}+\ldots \tag{153}
\end{equation*}
$$

Under a gauge transformation the vector field $A_{\mu}$ becomes

$$
\begin{equation*}
A_{\mu}^{U}=U A_{\mu} U^{-1}+i\left(\partial_{\mu} U\right) U^{-1} \equiv A_{\mu}+\delta A_{\mu} \tag{154}
\end{equation*}
$$

For an infinitesimal transformation the change of $A_{\mu}$ is

$$
\begin{equation*}
\delta A_{\mu}=i \epsilon^{a}\left[\lambda^{a}, A_{\mu}\right]-\partial_{\mu} \epsilon^{a} \lambda^{a}+O\left(\epsilon^{2}\right) \tag{155}
\end{equation*}
$$

where $\lambda^{a}$ are the generators of the algebra of the gauge group $G$.
On the other hand, since

$$
\begin{equation*}
\frac{\delta g^{a}}{\delta \epsilon^{b}}=\frac{\partial g^{a}}{\partial A_{\mu}^{c}} \frac{\delta A_{\mu}^{c}}{\delta \epsilon^{b}} \tag{156}
\end{equation*}
$$

we can also write

$$
\begin{align*}
\delta A_{\mu}^{c} & =2 i \epsilon^{b} \operatorname{tr}\left(\lambda^{c}\left[\lambda^{b}, A_{\mu}\right]\right)-2 \partial_{\mu} \epsilon^{b} \operatorname{tr}\left(\lambda^{c} \lambda^{b}\right)+O\left(\epsilon^{2}\right) \\
& =i \epsilon^{b} \operatorname{tr}\left(\lambda^{c}\left[\lambda^{b}, \lambda^{d}\right]\right) A_{\mu}^{d}-\partial_{\mu} \epsilon^{b} \delta_{b c} \\
& =-2 f^{b d e} \epsilon^{b} \operatorname{tr}\left(\lambda^{c} \lambda^{e}\right) A_{\mu}^{d}-\partial_{\mu} \epsilon^{b} \delta_{b c} \\
& =-f^{b d e} \epsilon^{b} A_{\mu}^{d}-\partial_{\mu} \epsilon^{b} \delta_{b c} \tag{157}
\end{align*}
$$

where $f^{a b c}$ are the structure constants of the Lie group $G$.
Hence, we find

$$
\begin{equation*}
\frac{\delta A_{\mu}^{c}(x)}{\delta \epsilon^{b}(y)}=-\left[\partial_{\mu} \delta_{b c}+f^{b c d} A_{\mu}^{d}\right] \delta(x-y) \equiv-D_{\mu}^{c d}[A] \delta(x-y) \tag{158}
\end{equation*}
$$

where we have denoted by $D_{\mu}[A]$ the covariant derivative in the adjoint representation, which in components is given by

$$
\begin{equation*}
D_{\mu}^{a b}[A]=\delta_{a b} \partial_{\mu}-f^{a b c} A_{\mu}^{c} \tag{159}
\end{equation*}
$$

Using these results we can put the Faddeev-Popov determinant (or Jacobian) in the form

$$
\begin{equation*}
\Delta_{F P}[A]=\operatorname{Det}\left(\frac{\delta g}{\delta \epsilon}\right)=\operatorname{Det}\left(\frac{\partial g^{a}}{\partial A_{\mu}^{c}} \frac{\delta A_{\mu}^{c}}{\delta \epsilon^{b}}\right) \tag{160}
\end{equation*}
$$

We will now define an operator $M_{F P}$ whose matrix elements are

$$
\begin{align*}
\langle x, a| M_{F P}|y, b\rangle & =\langle x, a| \frac{\partial g}{\partial A_{\mu}^{c}} \frac{\delta A_{\mu}^{c}}{\delta \epsilon}|y, b\rangle \\
& =\int_{z} \frac{\partial g^{a}(x)}{\partial A_{\mu}^{c}(z)} \frac{\delta A_{\mu}^{c}(z)}{\delta \epsilon^{b}(y)} \\
& =-\int_{z} \frac{\partial g^{a}(x)}{\partial A_{\mu}^{c}(z)} D_{\mu}^{c b} \delta(z-y) \tag{161}
\end{align*}
$$

For the case of $g^{a}[A]=\partial^{\mu} A_{\mu}^{a}(x)-c^{a}(c)$, appropriate for the Feynman-‘t Hooft gauges, we have

$$
\begin{equation*}
\frac{\partial g^{a}(x)}{\partial A_{\mu}^{c}(z)}=\delta_{a c} \partial^{\mu} \delta(x-z) \tag{162}
\end{equation*}
$$

and also

$$
\begin{align*}
\langle x, a| M_{F P}|y, b\rangle & =-\int_{z} \delta_{a c} \partial_{x}^{\mu} \delta(x-z) D_{\mu}^{c b}[A] \delta(z-y) \\
& =-\int_{z} \delta_{a c} \delta(x-z) \partial_{z}^{\mu} D_{\mu}^{c b} \delta(x-y) \\
& =-\partial^{\mu} D_{\mu}^{a b} \delta(x-y) \tag{163}
\end{align*}
$$

Thus, the Faddeev-Popov determinant now is

$$
\begin{equation*}
\Delta_{F P}=\operatorname{Det}\left(\partial^{\mu} D_{\mu}[A]\right) \tag{164}
\end{equation*}
$$

Notice that in the non-Abelian case this determinant is an explicit function of the gauge field. Since it is a determinant, it can be written as a path integral over a set of fermionic ghost fields, denoted by $\eta_{a}(x)$ and $\bar{\eta}_{a}(x)$, one per gauge condition (i.e. one per generator):

$$
\begin{equation*}
\operatorname{Det}\left[\partial^{\mu} D_{\mu}\right]=\int \mathcal{D} \eta_{a} \mathcal{D} \bar{\eta}_{a} e^{i \int d^{D} x \bar{\eta}_{a}(x) \partial^{\mu} D_{\mu}^{a b}[A] \eta_{b}(x)} \tag{165}
\end{equation*}
$$

Notice that these are fields quantized with the "wrong" statistics. In other words, these "particles" do not satisfy the general conditions for causality and unitarity to be obeyed. Hence these ghosts cannot create physical states (thereby their ghostly character).

The full form of the path integral of a Yang-Mills gauge theory with coupling constant $g$, in the Feynman-‘t Hooft covariant gauges with gauge parameter $\lambda$, is given by

$$
\begin{equation*}
Z=\int \mathcal{D} A \mathcal{D} \eta \mathcal{D} \bar{\eta} e^{i \int d^{D} x \mathcal{L}_{Y M}[A, \eta, \bar{\eta}]} \tag{166}
\end{equation*}
$$

where $\mathcal{L}_{Y M}$ is the effective Lagrangian density

$$
\begin{equation*}
\mathcal{L}_{Y M}[A, \eta, \bar{\eta}]=-\frac{1}{2 g^{2}} \operatorname{tr} F_{\mu \nu} F^{\mu \nu}+\frac{\lambda}{2 g^{2}}\left(\partial_{\mu} A^{\mu}\right)^{2}-\bar{\eta} \partial_{\mu} D^{\mu}[A] \eta \tag{167}
\end{equation*}
$$

Thus the pure gauge theory, even in the absence of matter fields, is non-linear. We will return to this problem later on when we look at both the perturbative and non-perturbative aspects of Yang-Mills gauge theories.

## 10 BRST Invariance

In the previous section we developed in detail the path-integral quantization of non-Abelian Yang-Mills gauge theories. We payed close attention to the role of gauge invariance and how to consistently fix the gauge in order to define the path-integral. Here we will show that the effective Lagrangian of a YangMills gauge field, Eq. (167), has an extended symmetry, closely related to supersymmetry. This extended symmetry plays a crucial role in proving the renormalizability of non-Abelian gauge theories.

Let us consider the QCD Lagrangian in the Feynman-'t Hooft covariant gauges (with gauge parameter $\lambda$ and coupling constant $g$ ). The Lagrangian density $\mathcal{L}$ of this theory is

$$
\begin{equation*}
\mathcal{L}=\bar{\psi}(i D-m) \psi-\frac{1}{4} F_{\mu \nu}^{a} F_{a}^{\mu \nu}-\frac{1}{2 \lambda} B_{a} B_{a}+B_{a} \partial^{\mu} A_{\mu}^{a}-\bar{\eta}^{a} \partial^{\mu} D_{\mu}^{a b} \eta^{b} \tag{168}
\end{equation*}
$$

Here $\psi$ is a Dirac Fermi field, which represents quarks and it transforms under the fundamental representation of the gauge group $G$; the "Hubbard-Stratonovich" field $B_{a}$ is an auxiliary field which has no dynamics of its own and it transforms as a vector in the adjoint representation of $G$.

Becchi, Rouet, Stora and Tyutin realized that this gauge-fixed Lagrangian has the following ("BRST") symmetry, where $\epsilon$ is an infinitesimal anti-commuting
parameter:

$$
\begin{align*}
\delta A_{\mu}^{a} & =\epsilon D_{\mu}^{a b} \eta_{b}  \tag{169}\\
\delta \psi & =i g \epsilon \eta^{a} t^{a} \psi  \tag{170}\\
\delta \eta^{a} & =-\frac{1}{2} g \epsilon f^{a b c} \eta_{b} \eta_{c}  \tag{171}\\
\delta \bar{\eta}^{a} & =\epsilon B^{a}  \tag{172}\\
\delta B^{a} & =0 \tag{173}
\end{align*}
$$

Here Eq. (169) and Eq. (170) are local gauge transformations and as such leave invariant the first two terms of the effective Lagrangian $\mathcal{L}$ of Eq. (168). The third term of Eq. (168) is trivial. The invariance of the fourth and fifth terms holds because the change of $\delta A$ in the fourth term cancels against the change of $\bar{\eta}$ in the fifth term. Finally, it remains to see that the changes of the fields $A_{\mu}$ and $\eta$ in the fifth term of Eq. (168) cancel out. To see that this is the case we check that
$\delta\left(D_{\mu}^{a b} \eta^{b}\right)=D_{\mu}^{a b} \delta \eta^{b}+g f^{a b c} \delta A_{\mu}^{b} \eta^{c}=-\frac{1}{2} g^{2} \epsilon f^{a b c} f^{c d e}\left(A_{\mu}^{b} \eta^{d} \eta^{e}+A_{\mu}^{d} \eta^{e} \eta^{b}+A_{\mu}^{e} \eta^{b} \eta^{d}\right)$
which vanishes due to the Jacobi identity for the structure constants

$$
\begin{equation*}
f^{a d e} f^{b c d}+f^{b d e} f^{c a d}+f^{c d e} f^{a b d}=0 \tag{175}
\end{equation*}
$$

or, equivalently, from the nested commutators of the generators $t^{a}$ :

$$
\begin{equation*}
\left[t^{a},\left[t^{b}, t^{c}\right]\right]+\left[t^{b},\left[t^{c}, t^{a}\right]\right]+\left[t^{c},\left[t^{a}, t^{b}\right]\right]=0 \tag{176}
\end{equation*}
$$

Hence, BRST is at least a global symmetry of the gauge-fixed action with gauge fixing parameter $\lambda$.

This symmetry has a remarkable property which follows from its fermionic nature. Let $\phi$ be any of the fields of the Lagrangian and $Q \phi$ be the BRST transformation of the field,

$$
\begin{equation*}
\delta \phi=\epsilon Q \phi \tag{177}
\end{equation*}
$$

For instance,

$$
\begin{equation*}
Q^{a} A_{\mu}^{a}=D_{\mu}^{a b} \eta^{b} \tag{178}
\end{equation*}
$$

and so on. It follows that for any field $\phi$

$$
\begin{equation*}
Q^{2} \phi=0 \tag{179}
\end{equation*}
$$

i.e. the BRST transformation of $Q \phi$ vanishes. This rule works for the field $A_{\mu}$ due to the transformation property of $\delta\left(D_{\mu}^{a b} \eta^{b}\right)$. It also holds for the ghosts since

$$
\begin{equation*}
Q^{2} \eta^{a}=\frac{1}{2} g^{2} f^{a b c} f^{b d e} \eta^{c} \eta^{d} \eta^{e}=0 \tag{180}
\end{equation*}
$$

which holds due to the Jacobi identity.

What are the implications of the existence of BRST as a continuous symmetry? To begin with it implies that there is a conserved self-adjoint charge $Q$ that must necessarily commute with the Hamiltonian $H$ of the Yang-Mills gauge theory. Above we saw how $Q$ acts on the fields, $Q^{2} \phi=0$, for all the fields in the Lagrangian. Hence, as an operator $Q^{2}=0$, that is, the BRST charge $Q$ is nilpotent, and it commutes with $H$. Let us now show that $Q$ divides the Hilbert space of the eigenstates of $H$ is three sectors

1. Many eigenstates of $H$ must be annihilated by $Q$ for $Q^{2}=0$ to hold. Let $\mathcal{H}_{1}$ be the set of eigenstates of $H$ which are not annihilated by $Q$. Hence, if $\left|\psi_{1}\right\rangle \in \mathcal{H}_{1}$, then $Q\left|\psi_{1}\right\rangle \neq 0$. Thus, the states in $\mathcal{H}_{1}$ are not BRST invariant.
2. Let us consider the subspace of states $\mathcal{H}_{2}$ of the form $\left|\psi_{2}\right\rangle=Q\left|\psi_{1}\right\rangle$, i.e. $\mathcal{H}_{2}=Q \mathcal{H}_{1}$. Then, for these states $Q\left|\psi_{2}\right\rangle=Q^{2}\left|\psi_{1}\right\rangle=0$. Hence, the states in $\mathcal{H}_{2}$ are BRST invariant but are the BRST transform of states in $\mathcal{H}_{1}$.
3. Finally, let $\mathcal{H}_{0}$ be the set of eigenstates of $H$ that are annihilated by $Q$, $Q\left|\psi_{0}\right\rangle=0$, but which are not in $\mathcal{H}_{2}$, i.e. $\left|\psi_{0}\right\rangle \neq Q\left|\psi_{1}\right\rangle$. Hence, the states in $\mathcal{H}_{0}$ are BRST invariant and are not the BRST transform of any other state. This is the physical space of states.

It follows from the above classification that the inner product of any pair of states in $\mathcal{H}_{2},\left|\psi_{2}\right\rangle$ and $\left|\psi_{2}^{\prime}\right\rangle$, have zero inner product:

$$
\begin{equation*}
\left\langle\psi_{2} \mid \psi_{2}^{\prime}\right\rangle=\left\langle\psi_{1}\right| Q\left|\psi_{2}^{\prime}\right\rangle=0 \tag{181}
\end{equation*}
$$

where we used that $\left|\psi_{2}\right\rangle$ is the BRST transform of a state in $\mathcal{H}_{1},\left|\psi_{1}\right\rangle$. Similarly, one can show that if $\left|\psi_{0}\right\rangle \in \mathcal{H}_{0}$, then $\left\langle\psi_{2} \mid \psi_{0}\right\rangle=0$.

What is the physical meaning of BRST and of this classification? Peskin and Schroeder give a simple argument. Consider the weak coupling limit of the theory, $g \rightarrow 0$. In this limit we can find out what BRST does by looking at the transformation properties of the fields that appear in the Lagrangian of Eq. (168). In particular, $Q$ transforms a forward polarized (i.e. longitudinal) component of $A_{\mu}$ into a ghost. At $g=0$, we see that $Q \eta=0$ and that the antighost $\bar{\eta}$ transforms into the auxiliary field $B$. Also, at the classical level, $B=$ $\lambda \partial^{\mu} A_{\mu}$. Hence, the auxiliary fields $B$ are backward (longitudinally) polarized quanta of $A_{\mu}$. Thus, forward polarized gauge bosons and anti-ghosts are in $\mathcal{H}_{1}$, since they are not the BRST transform of states created by other fields. Ghosts and backward polarized gauge bosons are in $\mathcal{H}_{2}$ since they are the BRST transform of the former. Finally, transverse gauge bosons are in $\mathcal{H}_{0}$. Hence, in general, states with ghosts, anti-ghosts, and gauge bosons with unphysical polarization belong either to $\mathcal{H}_{1}$ or $\mathcal{H}_{2}$. Only the physical states belong to $\mathcal{H}_{0}$. It turns out that the $S$-matrix, when restricted to the physical space $\mathcal{H}_{0}$, is unitary (as it should).

