

# Phys 582 – General Field Theory

## Problem Set No.1 Solutions

Di Zhou

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### 1 The Landau Theory of Phase Transitions as a Classical Field Theory

1. Due to the free energy density  $\varepsilon$ , the free energy is:

$$F = \int \varepsilon d^3x = \int \left( \frac{1}{2} (\nabla\phi(\vec{x}))^2 + U(\phi(\vec{x})) \right) d^3x \quad (1)$$

To find the saddle point of this free energy, we set the variation of free energy to be zero:

$$\delta F = \int \left( \frac{\delta\varepsilon}{\delta\nabla\phi} \delta\nabla\phi + \frac{\delta\varepsilon}{\delta\phi} \delta\phi \right) d^3x = \int \left( -\nabla \left( \frac{\delta\varepsilon}{\delta\nabla\phi} \right) + \frac{\delta\varepsilon}{\delta\phi} \right) \delta\phi d^3x \quad (2)$$

Here we have used: (1)  $\nabla\delta\phi = \delta\nabla\phi$ ; (2) the variation vanishes at the boundaries. The saddle-point equation is, to let the variation of  $\varepsilon$  to be 0, i.e.,

$$-\nabla \left( \frac{\delta\varepsilon}{\delta\nabla\phi} \right) + \frac{\delta\varepsilon}{\delta\phi} = 0 \Rightarrow -\nabla^2\phi + \frac{\delta U}{\delta\phi} = 0 \quad (3)$$

Plugging in the form of  $U(\phi)$  in this problem, we have:

$$-\nabla^2\phi + m_0^2\phi + \frac{\lambda_4}{3!}\phi^3 + \frac{\lambda_6}{5!}\phi^5 = 0 \quad (4)$$

2. The constant field  $\phi$  implies that  $\nabla\phi = 0$ . Eq. (4) reduces into:

$$m_0^2\phi + \frac{\lambda_4}{3!}\phi^3 + \frac{\lambda_6}{5!}\phi^5 = 0 \Rightarrow \bar{\phi}^2 = 0 \quad \text{or} \quad \bar{\phi}^2 = -10 \frac{\lambda_4}{\lambda_6} \left( 1 \mp \sqrt{1 - \frac{6m_0^2\lambda_6}{5\lambda_4^2}} \right) \quad (5)$$

Since  $\lambda_4 < 0, \lambda_6 > 0$  and the field  $\bar{\phi}$  is assumed to be a real field, imaginary  $\bar{\phi}$  is meaningless. This implies that for  $T < T_0$  and  $T > T_0$  the number of meaningful solutions is different. What is more, above  $T_0$  also exists a critical transition temperature  $T^* > T_0$  that, the "nontrivial free energy" is lower than "trivial

free energy”:  $F(\bar{\phi} \neq 0) < F(\bar{\phi} = 0)$ . The system tends to be in the lower free energy state, equivalently speaking, the symmetry breaking state. The critical temperature can be determined by  $F(\bar{\phi}_c \neq 0) = F(\bar{\phi} = 0)$ . Therefore we need to consider the corresponding free energy density:

$$\varepsilon(\bar{\phi}_c \neq 0) = \varepsilon(\bar{\phi} = 0) \Rightarrow \left( m_0^2 + \frac{\lambda_4}{12} \bar{\phi}_c^2 + \frac{\lambda_6}{360} \bar{\phi}_c^4 \right) \bar{\phi}_c^2 = 0 \quad (6)$$

$\bar{\phi}_c^2 \neq 0$  due to  $T^* > T_0 \Rightarrow m_0^2 > 0$ , we can solve the above equation as:

$$m_0^2 + \frac{\lambda_4}{12} \bar{\phi}_c^2 + \frac{\lambda_6}{360} \bar{\phi}_c^4 = 0 \Rightarrow \bar{\phi}_c^2 = 180 \left( -\frac{\lambda_4}{12\lambda_6} \pm \sqrt{\frac{\lambda_4^2}{144\lambda_6^2} - \frac{m_0^2}{90\lambda_6}} \right) \quad (7)$$

On the other hand, recall the nontrivial solution in Eq.(5), we have:

$$\bar{\phi}_c^2 = 180 \left( -\frac{\lambda_4}{12\lambda_6} \pm \sqrt{\frac{\lambda_4^2}{144\lambda_6^2} - \frac{m_0^2}{90\lambda_6}} \right) = -10 \frac{\lambda_4}{\lambda_6} \left( 1 \mp \sqrt{1 - \frac{6m_0^2\lambda_6}{5\lambda_4^2}} \right) \quad (8)$$

Let us denote  $\frac{m_0^2\lambda_6}{\lambda_4^2} = x$  and NOTE: the above equation corresponds to 4 equations!

$$\frac{1}{2} \pm \sqrt{\frac{9}{4} - \frac{18}{5}x} = \pm \sqrt{1 - \frac{6}{5}x} \quad \text{and} \quad \frac{1}{2} \pm \sqrt{\frac{9}{4} - \frac{18}{5}x} = \mp \sqrt{1 - \frac{6}{5}x} \quad (9)$$

Although there are 4 equations, the Physical solution (i.e., real field solution and lower free energy requirements) is only one:

$$x = \frac{5}{8} \Rightarrow T^* = \frac{5}{8} \frac{\lambda_4^2}{a\lambda_6} + T_0 \quad (10)$$

At the transition temperature  $T = T^* + \epsilon \rightarrow T = T^* - \epsilon$ , the mean-field solution switches from  $\bar{\phi}^2 = 0 \rightarrow \bar{\phi}^2 = -\frac{15\lambda_4}{\lambda_6}$ : this is the incontinuous phase transition, the first-order phase transition.

Up to now, we still haven't decided which  $\bar{\phi}^2$  of the two solutions in Eq.(5) should be the real ground state. In the two states, the mean-field free energy is given by:

$$\varepsilon = \frac{-\lambda_4^3}{\lambda_6^2} \left( \frac{5}{9} + \frac{5}{3} \frac{m_0^2\lambda_6}{\lambda_4^2} \mp \frac{5}{9} \sqrt{1 - \frac{6m_0^2\lambda_6}{5\lambda_4^2}} \right) \text{ for } \bar{\phi}^2 = \frac{-10\lambda_4}{\lambda_6} \left( 1 \mp \sqrt{1 - \frac{6m_0^2\lambda_6}{5\lambda_4^2}} \right) \quad (11)$$

Of course the minus-sign free-energy is smaller, which corresponds to the mean-field solution also with the minus sign. The plot for  $U(\bar{\phi})$  to  $\bar{\phi}$  is from fig.1 to fig.6.

3. Now the four-point coupling constant  $\lambda_4 > 0$ ,

$$T > T_0, m_0^2 > 0 \Rightarrow \min(\varepsilon) = 0 \quad \text{when} \quad \bar{\phi}^2 = 0$$

$$T < T_0, m_0^2 < 0 \Rightarrow \min(\varepsilon) < 0 \quad \text{when} \quad \bar{\phi}^2 = 10 \frac{\lambda_4}{\lambda_6} \left( -1 + \sqrt{1 - \frac{6m_0^2\lambda_6}{5\lambda_4^2}} \right) \quad (12)$$

This implies that the transition temperature occurs at temperature  $T = T_0$ . At  $T = T_0$  the mean-field solution is exactly 0. Therefore the order parameter,  $\bar{\phi}$  is continuous. This is the second-order phase transition. (Moreover, you can prove the derivative of order parameter to temperature is discontinuous, of order  $(T - T_0)^{-1}$ ).

4. Recall Eq.(10) we have the expression for  $\lambda_4 < 0$  gives the behavior of the phase boundary as a function of  $T^* - T_0$ , which is the first order phase transition.

$$\lambda_4 = -\sqrt{\frac{8a\lambda_6}{5} (T^* - T_0)} \quad (13)$$

For  $\lambda_4 > 0$ , the phase boundary always occurs at  $T = T_0$ , and it corresponds to the second order phase transition. The plot for the boundary is fig.9.

## 2 Scalar Electrodynamics

1. For a local transformation,

$$A'_\mu(x) = A_\mu(x) + \partial_\mu \Lambda(x) \Rightarrow D'_\mu = \partial_\mu + ieA_\mu + ie\partial_\mu \Lambda \quad (14)$$

The complex scalar field potential terms are of course invariant:

$$\begin{aligned} |\phi'(x)|^2 &= \phi'^*(x)\phi'(x) = \phi^*(x)e^{ie\Lambda(x)}\phi(x)e^{-ie\Lambda(x)} = \phi^*(x)\phi(x) = |\phi(x)|^2 \\ (|\phi'(x)|^2)^2 &= (|\phi(x)|^2)^2 \end{aligned} \quad (15)$$

The EM field stress tensor transforms like:

$$F^{\mu\nu} = \partial^\mu A^\nu - \partial^\nu A^\mu = \partial^\mu (A^\nu + \partial^\nu \Lambda) - \partial^\nu (A^\mu + \partial^\mu \Lambda) = F^{\mu\nu} \quad (16)$$

Finally the scalar-gauge field dynamic term transforms as:

$$\begin{aligned} D'_\mu \phi'(x) &= (\partial_\mu + ieA_\mu + ie\partial_\mu \Lambda) (\phi(x)e^{-ie\Lambda(x)}) = D_\mu \phi(x)e^{-ie\Lambda(x)} \\ \Rightarrow (D'_\mu \phi')^* (D'^\mu \phi') &= (D_\mu \phi e^{-ie\Lambda})^* (D^\mu \phi e^{-ie\Lambda}) = (D_\mu \phi)^* (D^\mu \phi) \end{aligned} \quad (17)$$

In conclusion, the total Lagrangian Density is invariant under gauge transformation.

2. The classical equations of motion comes from the Lagrange Equation,

$$\frac{\delta \mathcal{L}}{\delta \phi} - \partial_\mu \left( \frac{\delta \mathcal{L}}{\delta \partial_\mu \phi} \right) = 0, \quad \frac{\delta \mathcal{L}}{\delta \phi^*} - \partial_\mu \left( \frac{\delta \mathcal{L}}{\delta \partial_\mu \phi^*} \right) = 0, \quad \frac{\delta \mathcal{L}}{\delta A_\nu} - \partial_\mu \left( \frac{\delta \mathcal{L}}{\delta \partial_\mu A_\nu} \right) = 0 \quad (18)$$

Solving these three equations, we obtain the more ugly form:

$$\begin{aligned} -\frac{1}{2}\partial^2 \phi^* + ieA^\mu \partial_\mu \phi^* + \frac{1}{2}ie(\partial_\mu A^\mu) \phi^* + \frac{1}{2}e^2 A^2 \phi^* - \frac{m^2}{2}\phi^* - \frac{\lambda}{12}|\phi|^2 \phi^* &= 0 \\ -\frac{1}{2}\partial^2 \phi - ieA^\mu \partial_\mu \phi - \frac{1}{2}ie(\partial_\mu A^\mu) \phi + \frac{1}{2}e^2 A^2 \phi - \frac{m^2}{2}\phi - \frac{\lambda}{12}|\phi|^2 \phi &= 0 \\ \frac{1}{2}(\partial^\nu \phi)^* ie\phi - \frac{1}{2}(ie\phi^*)\partial^\nu \phi + e^2 A^\nu \phi^* \phi + \partial_\mu F^{\mu\nu} &= 0 \end{aligned} \quad (19)$$

However, due to the newly defined  $D_\mu = \partial_\mu + ieA_\mu$  and  $D^\mu = \partial^\mu + ieA^\mu$ , one can expect the simplified version of the above equations by using  $D_\mu$ :

$$\begin{aligned}
-\frac{1}{2}(D^\mu D_\mu \phi)^* - \frac{m^2}{2}\phi^* - \frac{\lambda}{12}|\phi|^2\phi^* &= 0 \\
-\frac{1}{2}D^\mu D_\mu \phi - \frac{m^2}{2}\phi - \frac{\lambda}{12}|\phi|^2\phi &= 0 \\
\frac{1}{2}ie[\phi(D^\nu \phi)^* - \phi^*(D^\nu \phi)] + \partial_\mu F^{\mu\nu} &= 0
\end{aligned} \tag{20}$$

3. The canonical momentum is:

$$\Pi = \frac{\delta \mathcal{L}}{\delta \partial_0 \phi} \Rightarrow H = \frac{\delta \mathcal{L}}{\delta \partial_0 \phi} \phi - \mathcal{L} \tag{21}$$

The scalar field and EM field momentums are:

$$\begin{aligned}
\Pi &= \frac{1}{2}(\partial_0 \phi^* - ieA_0 \phi^*) = \frac{1}{2}(D_0 \phi)^* \\
\Pi^* &= \frac{1}{2}(\partial_0 \phi + ieA_0 \phi) = \frac{1}{2}(D_0 \phi) \\
\Pi^i &= F^{i0}
\end{aligned} \tag{22}$$

The Hamiltonian Density by definition, is

$$\begin{aligned}
h &= \Pi \partial_0 \phi + \Pi^* \partial_0 \phi^* + \Pi^\mu \partial_0 A_\mu - \mathcal{L} \\
&= \frac{1}{2}(D_0 \phi)^* \partial_0 \phi + \frac{1}{2}(D_0 \phi) \partial_0 \phi^* + F^{\mu 0} \partial_0 A_\mu - \mathcal{L} \\
&= (D_0 \phi)^* (D_0 \phi) + \frac{1}{2}(D_0 \phi)^* (-ieA_0) \phi + \frac{1}{2}D_0 \phi (ieA_0) \phi^* + E_i \partial_0 A_i - \mathcal{L} \\
&= \frac{1}{2}(D_0 \phi)^* (D_0 \phi) + \frac{1}{2}(D_i \phi)^* (D_i \phi) + \frac{1}{2}ieA_0 [\phi^* (D_0 \phi) - \phi (D_0 \phi)^*] \\
&\quad + U(\phi) + E_i \partial_0 A_i + \frac{1}{4}F^2
\end{aligned} \tag{23}$$

Note from Eq.(19): the Conserved Quantity, Charge, is the special case of that of the third equation of Eq.(19), and note:  $D^0 \phi = 2\Pi^*$ :

$$ie(\phi \Pi - \phi^* \Pi^*) + \partial_\mu F^{\mu 0} = 0 \tag{24}$$

Therefore we have:

$$-\int d^3x \frac{ieA_0}{2} (\phi \Pi - \phi^* \Pi^*) = \int d^3x A_0 \partial_\mu F^{\mu 0} = -\int d^3x (\partial_\mu A_0) F^{\mu 0} \tag{25}$$

Where we have dropped the boundary term. Combining this back into Eq.(23) with the 3rd and 5th terms, we have:

$$\begin{aligned}
\int d^3x \left( \frac{ieA_0}{2} [\phi^* (D^0 \phi) - \phi (D^0 \phi)^*] + F^{\mu 0} \partial_0 A_\mu \right) &= \int d^3x (\partial_0 A_\mu - \partial_\mu A_0) F^{\mu 0} \\
&= \int d^3x F_{0\mu} F^{\mu 0} = \int d^3x E_i^2 \tag{26}
\end{aligned}$$

Use this conclusion back into Eq.(23), with  $\frac{1}{4}F^2$ , we finally have,

$$H = \int d^3x \left( \frac{1}{2} \Pi^* \Pi + \frac{1}{2} (D_i \phi)^* (D_i \phi) + U(\phi) + \frac{1}{2} (E^2 + B^2) \right) \quad (27)$$

This, together with Gauss's Law, Eq.(24), gives the Hamiltonian of the system. Let us have a look at the third equation in Eq.(22),  $\Pi^i$ . In the case of the gauge field, there is no canonical momentum conjugate to  $A_0$  which is why it is a Lagrange multiplier field that enforces a constraint, Gauss' Law. This is why Professor Fradkin suggest me not to use the notation  $\Pi^\mu$  but to use  $\Pi^i$  since  $\Pi^0 = 0$ .

4. Using  $\rho, \theta$  instead of  $\phi^*, \phi$ ,

$$D_\mu \phi = (\partial_\mu \rho + i\rho(\partial_\mu \theta + eA_\mu)) e^{i\theta} \quad (28)$$

From Eq.(19),

$$\begin{aligned} \partial^2 \rho - \rho(\partial_\mu \theta + eA_\mu)^2 + m_0^2 \rho + \frac{\lambda}{6} \rho^3 &= 0 \\ \rho(\partial^2 \theta + e\partial^\mu A_\mu) + (\partial^\mu \rho)(2\partial_\mu \theta + eA_\mu) &= 0 \\ e\rho(\partial^\nu \theta + eA^\nu) + \partial_\mu F^{\mu\nu} &= 0 \end{aligned} \quad (29)$$

For the London Gauge  $\theta = 0$ , the equations of motions are reduced into:

$$\begin{aligned} \partial^2 \rho - \rho(eA_\mu)^2 + m_0^2 \rho + \frac{\lambda}{6} \rho^3 &= 0 \\ \rho \partial^\mu A_\mu + A_\mu \partial^\mu \rho &= 0 \\ e^2 \rho A^\nu + \partial_\mu F^{\mu\nu} &= 0 \end{aligned} \quad (30)$$

The Lagrangian Density is,

$$\mathcal{L} = \frac{1}{2} \partial^\mu \rho \partial_\mu \rho + \frac{1}{2} e^2 A^2 \rho^2 - \frac{1}{2} m_0^2 \rho^2 - \frac{\lambda}{4!} \rho^4 - \frac{1}{4} F^2 \quad (31)$$

5. If  $m_0^2 < 0$ , then  $-\frac{1}{2}m_0^2 > 0$ . For the case  $\rho = \bar{\rho}$  the effective Lagrangian Density for London Gauge is reduced from Eq.(31) that, we set  $\partial_\mu \rho$  to be 0,

$$\mathcal{L} = \left( \frac{1}{2} e^2 A^2 \rho^2 - \frac{1}{2} m_0^2 \rho^2 - \frac{\lambda}{4!} \rho^4 \right) - \frac{1}{4} F^2 \quad (32)$$

The equations of motion in Eq.(30) gives the minimization of the Lagrangian Density,

$$\rho = \sqrt{\frac{6}{\lambda} (e^2 A^2 + |m_0^2|)}, \quad \partial_\mu A^\mu = 0, \quad A_\mu = 0 \Rightarrow \bar{\rho} = \sqrt{\frac{6}{\lambda} |m_0^2|} \quad (33)$$

Fix the classical solution, and plug back into our Lagrangian Density, Eq.(32),

$$\mathcal{L} = \frac{3|m_0^2|}{2\lambda} (2e^2 A^2 + |m_0^2|) - \frac{1}{4} F^2 = \mathcal{L}_0 + \frac{1}{2} m_{\text{ph}}^2 A^2 - \frac{1}{4} F^2 \quad (34)$$

Where  $m_{\text{ph}}^2 = \bar{\rho}e^2$ . Now the equations of motion for the fluctuation fields are,

$$\frac{\delta \mathcal{L}}{\delta A_\nu} - \partial_\mu \left( \frac{\delta \mathcal{L}}{\delta \partial_\mu A_\nu} \right) = 0 \Rightarrow m_{\text{ph}}^2 A^\nu + \partial_\mu F^{\nu\mu} = 0 \quad (35)$$

From the previous equations of motion that,  $\partial_\mu A^\mu = 0$  in Eq.(33), we have  $\partial_\mu F^{\nu\mu}$  to be,  $\partial_\mu (\partial^\nu A^\mu - \partial^\mu A^\nu) = -\partial^2 A^\nu$ . Thus Eq.(35) is reduced into:

$$m_{\text{ph}}^2 A^\nu - \partial^2 A^\nu = 0 \quad (36)$$

Which is the famous Klein-Gordon equation. Here, the  $m_{\text{ph}}^2 = \bar{\rho}e^2$  is the effective photon mass.

### 3 The Dirac Equation

1. Two ways to approach this problem. First is to find the algebra properties of  $\gamma^\mu$  matrices:

$$(\gamma^0)^+ = \gamma^0, \quad (\gamma^{1,2,3})^+ = -\gamma^{1,2,3} \quad (37)$$

With these relations the Dirac Equation  $(i\partial - m)\psi = 0$  transforms as:

$$\gamma^0 [(i\partial - m)\psi]^+ = 0 \Rightarrow \gamma^0 \psi^+ \left[ -i \left( \overleftarrow{\partial} \right)^+ - m \right] = 0 \quad (38)$$

Since,

$$\begin{aligned} (\partial_0)^+ &= \partial_0 \Rightarrow (\partial_0 \gamma^0)^+ = \partial_0 \gamma^0 \\ (\partial_{1,2,3})^+ &= -\partial_{1,2,3} \Rightarrow (\partial_{1,2,3} \gamma^{1,2,3})^+ = \partial_{1,2,3} \gamma^{1,2,3} \\ &\Rightarrow (\overleftarrow{\partial})^+ = \overleftarrow{\partial} \end{aligned} \quad (39)$$

Therefore, the Dirac Equation can be written as,

$$\gamma^0 \psi^+ \left( -i \overleftarrow{\partial} - m \right) = 0 \Rightarrow \bar{\psi} \left( -i \overleftarrow{\partial} - m \right) = 0 \quad (40)$$

Now let us take care of the 4-current. To prove it is conserved, what we need is to prove  $\partial_\mu j^\mu = 0$  (below we have used Eq.(38) and the Dirac Equation):

$$\partial_\mu j^\mu = \partial_\mu (\bar{\psi} \gamma^\mu \psi) = (\partial_\mu \bar{\psi}) \gamma^\mu \psi + \bar{\psi} \gamma^\mu (\partial_\mu \psi) = (\overleftarrow{\partial} \bar{\psi}) \psi + \bar{\psi} (\overrightarrow{\partial} \psi) = 0 \quad (41)$$

Another way for this problem, is to use the Lagrangian Density  $\mathcal{L} = i\bar{\psi} \overleftarrow{\partial} \psi - m\bar{\psi} \psi$ , and use the equations of motion,

$$\partial_\mu \left( \frac{\delta \mathcal{L}}{\delta \partial_\mu \psi} \right) - \frac{\delta \mathcal{L}}{\delta \psi} = 0 \quad \partial_\mu \left( \frac{\delta \mathcal{L}}{\delta \partial_\mu \bar{\psi}} \right) - \frac{\delta \mathcal{L}}{\delta \bar{\psi}} = 0 \quad (42)$$

We can also reach the same conclusion.

2. If the spinor satisfies Dirac Equation,  $(i\partial - m)\psi = 0$ ,

$$(i\partial + m)(i\partial - m)\psi = 0 \Rightarrow (i\partial \cdot i\partial - m^2)\psi = 0 \quad (43)$$

The first term above turns out to be:

$$\begin{aligned} \partial \cdot \partial &= \partial_\mu \partial_\nu \gamma^\mu \gamma^\nu = \partial_\mu \partial_\nu \left( \frac{1}{2} \{\gamma^\mu, \gamma^\nu\} + \frac{1}{2} [\gamma^\mu, \gamma^\nu] \right) \\ \partial \cdot \partial &= \partial_\nu \partial_\mu \gamma^\nu \gamma^\mu = \partial_\nu \partial_\mu \left( \frac{1}{2} \{\gamma^\nu, \gamma^\mu\} + \frac{1}{2} [\gamma^\nu, \gamma^\mu] \right) \end{aligned} \quad (44)$$

Sum these two equations up, use the symmetric property for  $\{\gamma^\mu, \gamma^\nu\}$  and anti-symmetric for  $[\gamma^\mu, \gamma^\nu]$ ,

$$2\partial \cdot \partial = \partial_\mu \partial_\nu \{\gamma^\mu, \gamma^\nu\} = 2\partial_\mu \partial_\nu g^{\mu\nu} = 2(\partial_0^2 - \partial_1^2 - \partial_2^2 - \partial_3^2) = 2\partial^2 \quad (45)$$

Therefore, the spinor satisfies Klein-Gordon Equation:

$$(\partial^2 + m^2)\psi = 0 \quad (46)$$

3. (a) Use the properties of gamma matrices,

$$\gamma^\mu \gamma^\nu = \frac{1}{2} \{\gamma^\mu, \gamma^\nu\} + \frac{1}{2} [\gamma^\mu, \gamma^\nu] = g^{\mu\nu} - i\sigma^{\mu\nu} \quad (47)$$

Thus the inner product of  $\mathcal{A}\mathcal{B}$  is,

$$\mathcal{A}\mathcal{B} = A_\mu B_\nu \gamma^\mu \gamma^\nu = A_\mu B_\nu (g^{\mu\nu} - i\sigma^{\mu\nu}) = A \cdot B - i\sigma_{\mu\nu} A^\mu B^\nu \quad (48)$$

(b) The trace take over the spinor indices, thus

$$\text{Tr}(\mathcal{A}\mathcal{B}) = \text{Tr}(A_\mu B_\nu g^{\mu\nu} - i\sigma_{\mu\nu} A^\mu B^\nu) = A_\mu B_\nu g^{\mu\nu} \text{Tr}(1_4) = 4A \cdot B \quad (49)$$

(c) Use  $\frac{1}{2} \{\gamma^\mu, \gamma^\nu\} = g^{\mu\nu}$ ,

$$\gamma^\lambda \gamma^\mu \gamma_\lambda = \gamma^\lambda \gamma^\mu \gamma^\lambda g_{\lambda\lambda} = 2g^{\lambda\mu} \gamma^\lambda g_{\lambda\lambda} - \gamma^\mu \gamma^\lambda \gamma^\lambda g_{\lambda\lambda} \quad (50)$$

Since

$$\gamma^\mu \gamma^\lambda \gamma^\lambda g_{\lambda\lambda} = 4\gamma^\mu, \quad g^{\lambda\mu} g_{\lambda\lambda} = g_\lambda^\mu = \delta_\lambda^\mu \quad (51)$$

Therefore we have

$$\gamma^\lambda \gamma^\mu \gamma_\lambda = 2\gamma^\mu - 4\gamma^\mu = -2\gamma^\mu \quad (52)$$

## 4 Transformation Properties of Field Bilinears in the Dirac Theory

(a) The spinor transformation is,  $\psi'(x') = S(\Lambda)\psi(x)$ , with  $S(\Lambda) = \exp\left(-\frac{i}{4}\sigma_{\mu\nu}\omega^{\mu\nu}\right)$ , therefore,

$$S(\Lambda)^+ = \left[ \exp\left(-\frac{i}{4}\sigma_{\mu\nu}\omega^{\mu\nu}\right) \right]^+ = \exp\left(\frac{i}{4}(\omega^{\mu\nu})^+ \sigma_{\mu\nu}^+\right) = \exp\left(\frac{i}{4}\omega^{\mu\nu}\sigma_{\mu\nu}^+\right) \quad (53)$$

Note  $(\omega^{\mu\nu})^+ = \omega^{\mu\nu}$  because  $\omega^{\mu\nu}$  is a number, and '+' has nothing to do with a number. Therefore, what we want to prove is,

$$\bar{\psi}'(x')\psi'(x') = \bar{\psi}(x)(\gamma^0)^{-1}S^+\gamma^0S\psi(x) = \bar{\psi}(x)\gamma^0S^+\gamma^0S\psi(x) \quad (54)$$

To prove the  $\gamma^0S^+\gamma^0S = I$ , use the Taylor Series,

$$\gamma^0S^+\gamma^0 = \gamma^0 \left[ \sum_{n=0}^{\infty} \left( \frac{i}{4}\omega^{\mu\nu}\sigma_{\mu\nu}^+ \right)^n \right] \gamma^0 = \sum_{n=0}^{\infty} \left( \frac{i}{4}\omega^{\mu\nu}\gamma^0\sigma_{\mu\nu}^+\gamma^0 \right)^n \quad (55)$$

Since  $(\sigma_{\mu\nu})^+ = -\frac{i}{2}[\gamma^{\nu+}, \gamma^{\mu+}]$ , and  $(\gamma^0)^+ = \gamma^0$ ,  $(\gamma^{1,2,3})^+ = -\gamma^{1,2,3}$ , we can prove the identity,

$$\gamma^0(\gamma^{\nu+}\gamma^{\mu+} - \gamma^{\mu+}\gamma^{\nu+})\gamma^0 = -[\gamma^\mu, \gamma^\nu] \Rightarrow \gamma^0\sigma_{\mu\nu}^+\gamma^0 = \sigma_{\mu\nu} \quad (56)$$

Therefore we reach the conclusion that,

$$\gamma^0S^+\gamma^0 = S^{-1} \Rightarrow \bar{\psi}'(x')\psi'(x') = \bar{\psi}(x)\psi(x) \quad (57)$$

(b) Continue from above,

$$\bar{\psi}'(x')\gamma_5\psi'(x') = \bar{\psi}(x)S^{-1}\gamma_5S\psi(x) = \bar{\psi}(x)S^{-1}i\gamma^0\gamma^1\gamma^2\gamma^3S\psi(x) \quad (58)$$

Use  $S^{-1}\gamma^\mu S (\Lambda^{-1})_\mu^\nu = \gamma^\nu$ , or, equivalently  $S^{-1}\gamma^\mu S = \gamma^\nu \Lambda_\nu^\mu$ ,

$$\begin{aligned} S^{-1}i\gamma^0\gamma^1\gamma^2\gamma^3S &= i(S^{-1}\gamma^0S)(S^{-1}\gamma^1S)(S^{-1}\gamma^2S)(S^{-1}\gamma^3S) \\ &= i(\gamma^\alpha\Lambda_\alpha^0)(\gamma^\beta\Lambda_\beta^1)(\gamma^\lambda\Lambda_\lambda^2)(\gamma^\delta\Lambda_\delta^3) \\ &= i\gamma^\alpha\gamma^\beta\gamma^\lambda\gamma^\delta\Lambda_\alpha^0\Lambda_\beta^1\Lambda_\lambda^2\Lambda_\delta^3 \\ &= i\epsilon^{\alpha\beta\lambda\delta}\gamma^0\gamma^1\gamma^2\gamma^3\Lambda_\alpha^0\Lambda_\beta^1\Lambda_\lambda^2\Lambda_\delta^3 \\ &= \gamma^5\epsilon^{\alpha\beta\lambda\delta}\Lambda_\alpha^0\Lambda_\beta^1\Lambda_\lambda^2\Lambda_\delta^3 = \gamma^5\det\Lambda \end{aligned} \quad (59)$$

This is what we need for this proof.

$$\bar{\psi}'(x')\gamma_5\psi'(x') = \bar{\psi}(x)\gamma_5\psi(x)\det\Lambda \quad (60)$$

(c) This is just the same as the proof in (b) problem:

$$\bar{\psi}'(x')\gamma^\mu\psi'(x') = \bar{\psi}(x)S^{-1}\gamma^\mu S\psi(x) = \bar{\psi}(x)\gamma^\nu\Lambda_\nu^\mu\psi(x) \quad (61)$$



(d) Again duplicate the process of (b),

$$\bar{\psi}'(x')\gamma_5\gamma^\mu\psi'(x') = \bar{\psi}(x)(S^{-1}\gamma_5S)(S^{-1}\gamma^\mu S)\psi(x) = \bar{\psi}(x)\gamma_5\gamma^\nu\psi(x)\Lambda_\nu^\mu\det\Lambda \quad (62)$$

(e) The proof for a combination of  $\gamma^\mu$  matrices is the same as that of one  $\gamma^\mu$  matrix:

$$\begin{aligned} \bar{\psi}'(x')\sigma^{\mu\nu}\psi'(x') &= \bar{\psi}(x)\left(S^{-1}\frac{i}{2}(\gamma^\mu\gamma^\nu - \gamma^\nu\gamma^\mu)S\right)\psi(x) \\ &= \bar{\psi}(x)\left(\frac{i}{2}(\gamma^\alpha\gamma^\beta - \gamma^\beta\gamma^\alpha)\right)\psi(x)\Lambda_\alpha^\mu\Lambda_\beta^\nu = \bar{\psi}(x)\sigma^{\mu\nu}\psi(x)\Lambda_\alpha^\mu\Lambda_\beta^\nu \end{aligned} \quad (63)$$

## 5 Chiral Symmetry

1. Using the new form of slash operators in Chiral Representation, since:

$$\not{\partial} = \partial_0\gamma^0 - \vec{\sigma} \cdot \vec{\partial} \quad (64)$$

The new Dirac Equation is written as:

$$\begin{pmatrix} -mI & -i\partial_0I - i\vec{\sigma} \cdot \vec{\partial} \\ -i\partial_0I + i\vec{\sigma} \cdot \vec{\partial} & -mI \end{pmatrix} \begin{pmatrix} \phi \\ \chi \end{pmatrix} = 0 \quad (65)$$

or, equivalently,

$$\begin{aligned} -m\phi - \left(i\partial_0I + i\vec{\sigma} \cdot \vec{\partial}\right)\chi &= 0 \\ -m\chi - \left(i\partial_0I - i\vec{\sigma} \cdot \vec{\partial}\right)\phi &= 0 \end{aligned} \quad (66)$$

2. The massless Dirac Equation with  $m = 0$ , reduces to:

$$\begin{aligned} \left(\partial_0I + \vec{\sigma} \cdot \vec{\partial}\right)\chi &= 0 \\ \left(\partial_0I - \vec{\sigma} \cdot \vec{\partial}\right)\phi &= 0 \end{aligned} \quad (67)$$

Hence  $\phi$  and  $\chi$  decouples. Let us denote  $\chi = (\chi_1, \chi_2)^T$  and  $\phi = (\phi_1, \phi_2)^T$ ,

$$\begin{aligned} \begin{pmatrix} \partial_0 + \partial_3 & \partial_1 - i\partial_2 \\ \partial_1 + i\partial_2 & \partial_0 - \partial_3 \end{pmatrix} \begin{pmatrix} \chi_1 \\ \chi_2 \end{pmatrix} &= 0 \\ \begin{pmatrix} -\partial_0 + \partial_3 & \partial_1 - i\partial_2 \\ \partial_1 + i\partial_2 & -\partial_0 - \partial_3 \end{pmatrix} \begin{pmatrix} \phi_1 \\ \phi_2 \end{pmatrix} &= 0 \end{aligned} \quad (68)$$

Plugging into the plane wave solution, i.e.,  $\chi_i = \chi_{i0}e^{ik_\mu x^\mu}$ ,  $\phi_i = \phi_{i0}e^{ip_\nu x^\nu}$ , the above two matrices are:

$$\begin{aligned} \begin{pmatrix} k_0 - k_3 & -k_1 + ik_2 \\ -k_1 - ik_3 & k_0 + k_3 \end{pmatrix} \begin{pmatrix} \chi_1 \\ \chi_2 \end{pmatrix} &= 0 \\ \begin{pmatrix} -p_0 - p_3 & -p_1 + ip_2 \\ -p_1 - ip_2 & -p_0 + p_3 \end{pmatrix} \begin{pmatrix} \phi_1 \\ \phi_2 \end{pmatrix} &= 0 \end{aligned} \quad (69)$$

Solving these two matrix equations, we get,

$$k_0^2 = \vec{k}^2, \quad p_0^2 = \vec{p}^2 \quad (70)$$

Therefore the dispersion relation is  $k_0 = \omega = \pm|\vec{k}|$ ,  $p_0 = \omega = \pm|\vec{p}|$ . The relations between two components of that of 2-spinors are:

$$\chi_2 = \frac{k_0 - k_3}{k_1 - ik_2} \chi_1, \quad \phi_2 = \frac{-p_0 - p_3}{p_1 - ip_2} \phi_1 \quad (71)$$

By definition  $\phi$  lives in the upper part of the spinor wave function, and  $\gamma_5$  matrix acting on it resulting the +1 eigenvalue, and hence  $\phi$  wave has the +1 chirality. Similar argument shows  $\chi$  has the -1 chirality.

3. First of all let us calculate what  $e^{i\gamma_5\theta}$  is (to use Taylor Expansion):

$$e^{i\gamma_5\theta} = \sum_{n=0}^{\infty} \frac{(i\theta)^n}{n!} (\gamma_5)^n = \sum_{n=2k}^{\infty} \frac{(i\theta)^{2k}}{(2k)!} + \sum_{n=2k+1}^{\infty} \frac{(i\theta)^{2k+1}}{(2k+1)!} \gamma_5 = \cos\theta + i\gamma_5 \sin\theta \quad (72)$$

Write in the matrix form:

$$e^{i\gamma_5\theta} = \begin{pmatrix} e^{i\theta} & 0 \\ 0 & e^{-i\theta} \end{pmatrix} \quad (73)$$

(a)  $\phi$  and  $\chi$  transforms as:

$$\begin{pmatrix} e^{i\theta} & 0 \\ 0 & e^{-i\theta} \end{pmatrix} \begin{pmatrix} \phi \\ \chi \end{pmatrix} = \begin{pmatrix} e^{i\theta} \phi \\ e^{-i\theta} \chi \end{pmatrix} \quad (74)$$

(b) By definition  $\bar{\psi} = \psi^+ \gamma^0$ ,  $\psi'^+ = \psi^+ e^{-i\gamma_5\theta}$ , we get  $\bar{\psi}' = \bar{\psi} \gamma^0 e^{-i\gamma_5\theta} \gamma^0 = \bar{\psi} e^{i\gamma_5\theta}$ .

(c) Using the conclusions obtained above,  $\bar{\psi}' \psi' = \bar{\psi} e^{2i\gamma_5\theta} \psi$ ;

$$\bar{\psi}' \gamma^\mu \psi' = \bar{\psi} e^{i\gamma_5\theta} \gamma^\mu e^{i\gamma_5\theta} \psi \quad (75)$$

Since,

$$\begin{pmatrix} e^{i\theta} & 0 \\ 0 & e^{-i\theta} \end{pmatrix} \gamma^\mu \begin{pmatrix} e^{i\theta} & 0 \\ 0 & e^{-i\theta} \end{pmatrix} = \gamma^\mu \quad (76)$$

We find the quantity is conserved,  $\bar{\psi}' \gamma^\mu \psi' = \bar{\psi} \gamma^\mu \psi$ .

(d) The form of Dirac Equation is written in Eq.(64). If  $m \neq 0$ , the coupling between  $\phi$  and  $\chi$  is non-zero. However, under a Chiral Transformation  $\phi$  and  $\chi$  rotate in the opposite angle, Dirac Equation cannot be invariant. Recall Eq.(64), if the Dirac Equation acts on the transformed spinor wave function,

$$\begin{aligned} & \begin{pmatrix} -m_{11}I & -i\partial_0 I - i\vec{\sigma} \cdot \vec{\partial} \\ -i\partial_0 I + i\vec{\sigma} \cdot \vec{\partial} & -m_{22}I \end{pmatrix} \begin{pmatrix} e^{i\theta} \phi \\ e^{-i\theta} \chi \end{pmatrix} = 0 \\ \Rightarrow & \begin{pmatrix} -m_{11}I e^{2i\theta} & -i\partial_0 I - i\vec{\sigma} \cdot \vec{\partial} \\ -i\partial_0 I + i\vec{\sigma} \cdot \vec{\partial} & -m_{22}I e^{-2i\theta} \end{pmatrix} \begin{pmatrix} \phi \\ \chi \end{pmatrix} = 0 \end{aligned} \quad (77)$$

The mass term breaks the Chiral Symmetry, the newly defined mass is,

$$m \begin{pmatrix} Ie^{-2i\theta} & 0 \\ 0 & Ie^{2i\theta} \end{pmatrix} = me^{-2i\gamma_5\theta} \quad (78)$$

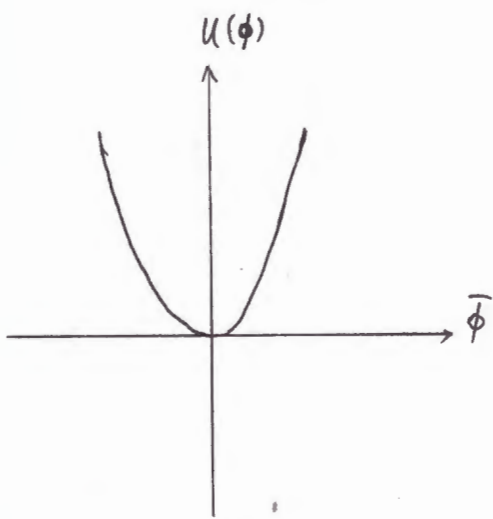


Fig. 1.  $x > \frac{5}{6}$

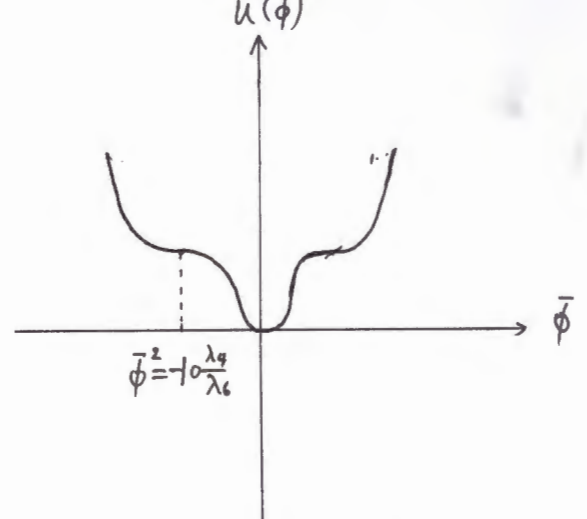


Fig 2.  $x = \frac{5}{6}$

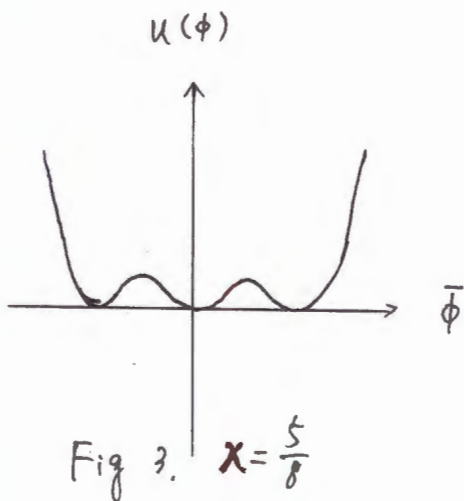


Fig 3.  $x = \frac{5}{8}$

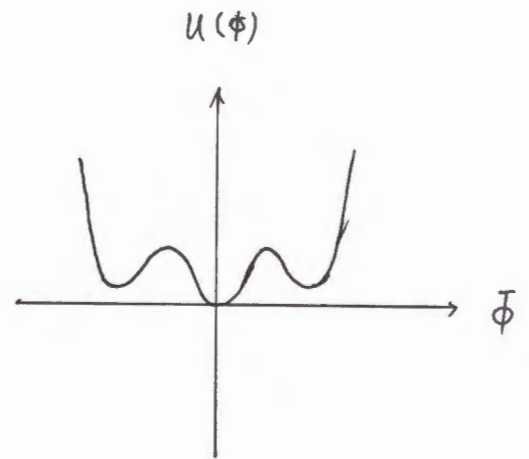


Fig 4.  $\frac{5}{6} > x > \frac{5}{8}$

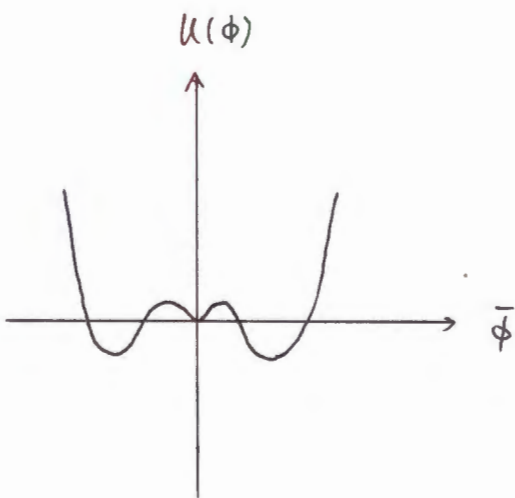


Fig 5.  $\frac{5}{8} > x > 0$

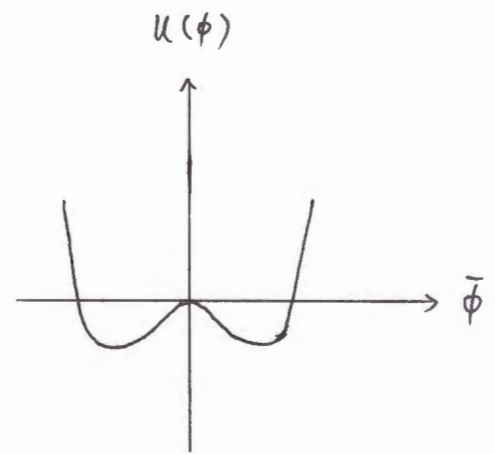


Fig 6.  $x < 0$

$\lambda_4 < 0$ : First-Order  
Phase Transition

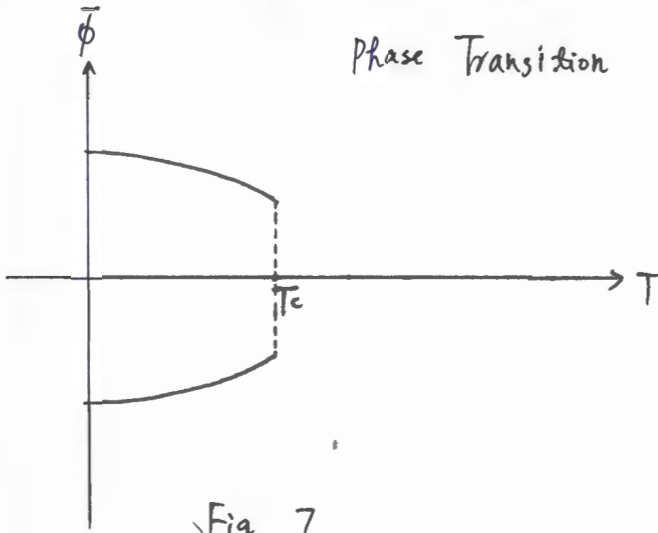


Fig. 7.

$\lambda_4 > 0$ : Second-Order  
Phase Transition

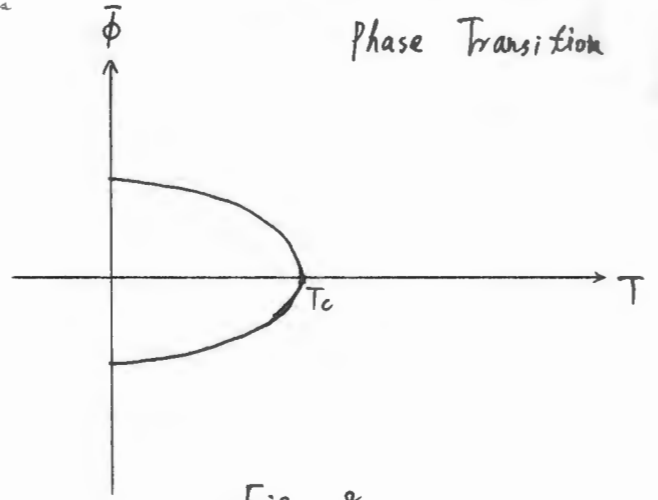


Fig. 8.

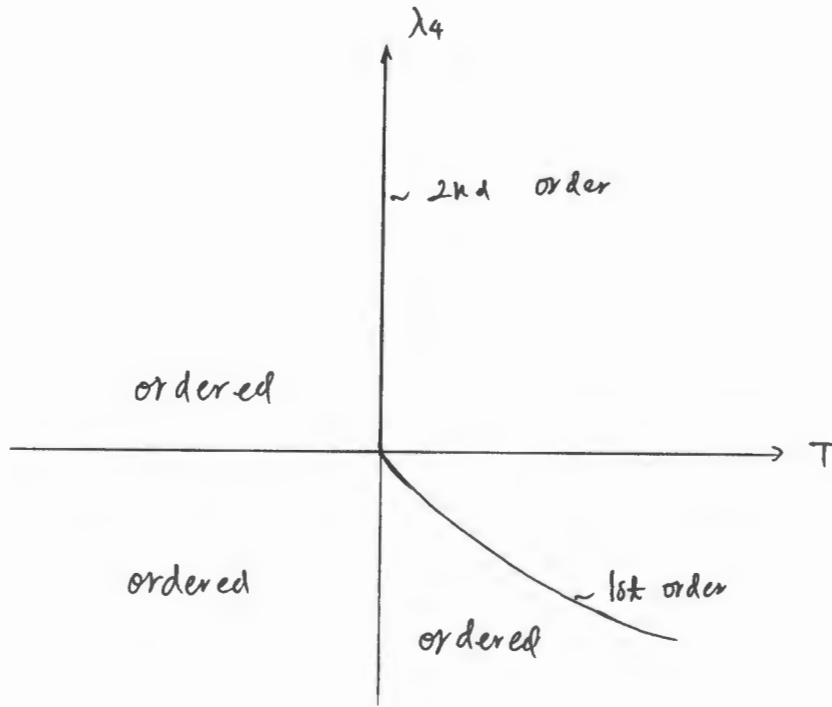


Fig. 9.