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### SCHIFF'S PROPOSED GYROSCOPE EXPERIMENT AS A TEST OF THE SCALAR-TENSOR THEORY OF GENERAL RELATIVITY

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We compare an explicit expression for the precession of a gyroscope in the Brans-Dicke scalar-tensor general relativity theory with the result derived by Schiff using Einstein's theory, and suggest that the gyroscope experiment offers the best possibility for testing the Brans-Dicke theory. Further, we conclude that the gyroscope in a satellite offers a more sensitive test than the earth-bound gyroscope.

Einstein's general theory of relativity is generally acclaimed as the correct theory of gravitation. Perhaps its only serious challenger is the scalar-tensor theory of Brans-Dicke (BD).<sup>1</sup> In the latter theory the gravitational constant is normalized to give the well-known red-shift result, and the dimensionless coupling constant  $\omega$  is selected to be  $\geq 6$  to ensure that the result for the precession of the perihelion of Mercury agrees, with an accuracy of 8% or less, with the computed value predicted by Einstein's theory. For  $\omega = 6$ , the BD theory gives a precession of 39.6'' arc/century which is about 3.43'' arc/century less than Einstein's value. The recent work of Dicke and Goldenberg<sup>2</sup> on the contribution of solar oblateness to the precession of the perihelion seems to favor the BD theory, but there is considerable controversy surrounding both the measurement itself and the relation<sup>4</sup> between the surface oblateness and the interior oblateness (the latter being the source of the quadrupole moment). It has recently been shown<sup>5</sup> that the rate of gravitational radiation from a system of binary stars in BD theory is smaller than the value predicted by Einstein's theory by a factor of  $(2\omega + 3)/(2\omega + 4)$ ; however, it seems that it will be a considerable time before this test is experimentally feasible. Cosmological tests<sup>6</sup> have likewise been unable to resolve the question. It is our purpose in this communication to suggest that perhaps the best test is the gyroscope experiment proposed by Schiff.<sup>7,8</sup> In particular, we write down an explicit expression for the precession of the gyroscope in BD theory for comparison with the Einstein value.

The angular velocity of precession in Einstein theory,  $\vec{\Omega}_E$  say, may be written as<sup>8</sup>

$$\vec{\Omega}_E = \vec{\Omega}_T + \vec{\Omega}_{DS} + \vec{\Omega}_{LT}, \quad (1)$$

where  $\vec{\Omega}_T$ ,  $\vec{\Omega}_{DS}$  and  $\vec{\Omega}_{LT}$  are the so-called Thomas, de Sitter, and Lense-Thirring contributions, respectively. Explicitly,<sup>8</sup>

$$\Omega_T = \frac{1}{2}(\vec{f} \times \vec{v}), \quad (2a)$$

$$\Omega_{DS} = (3m/2r^3)(\vec{r} \times \vec{v}), \quad (2b)$$

$$\Omega_{LT} = (I/r^3)[(3\vec{r}/r^2)(\vec{\omega} \cdot \vec{r}) - \vec{\omega}], \quad (2c)$$

where  $\vec{f}$  is the acceleration arising from any nongravitational constraint,  $m$  is the mass of the gyroscope ( $c = G = 1$ ),  $\vec{r}$  its position vector with respect to the center of the earth,  $\vec{v}$  is its velocity vector, and  $I$  and  $\omega$  are the moment of inertia and rotational angular velocity of the earth, respectively.

Following Eddington<sup>9</sup> and Robertson<sup>10</sup>, Schiff<sup>11</sup> has written the metric for the nonrotating earth in its most general isotropic form:

$$ds^2 = [1 - 2\alpha(m/r) + 2\beta(m/r)^2 + \dots] dt^2 - [1 + 2\gamma(m/r) + \dots](dx^2 + dy^2 + dz^2), \quad (3)$$

and deduces that the de Sitter term is modified by a factor  $(\alpha + 2\gamma)/3$ . For the particular case of the BD theory it is easy to show that this factor is  $(3\omega + 4)/(3\omega + 6)$ . Being a special-relativistic effect only, the Thomas precession remains unchanged in the BD theory. However, there is a change in the Lense-Thirring effect which is deduced quite easily from an observation made by the present author and

Salmona<sup>5</sup> to the effect that, in the weak-field limit, the solutions of the transformed<sup>12</sup> BD equations (the so-called barred system) are exactly the same as the solutions to Einstein's equation except for a factor. Explicitly,<sup>5</sup>

$$(\bar{h}_{\mu\nu})_{\text{BD}} = [(2\omega + 3)/(2\omega + 4)](h_{\mu\nu})_{\text{E}} \quad (4)$$

in the weak-field limit. This immediately enables us to conclude that the Lense-Thirring precession is reduced by a similar factor.<sup>13</sup> Thus, the angular velocity of precession in BD theory,  $\Omega_{\text{BD}}$  say, may be written as

$$\begin{aligned} \vec{\Omega}_{\text{BD}} = & \vec{\Omega}_{\text{T}} + [(4 + 3\omega)/(6 + 3\omega)]\vec{\Omega}_{\text{DS}} \\ & + [(3 + 2\omega)/(4 + 2\omega)]\vec{\Omega}_{\text{LT}}. \end{aligned} \quad (5)$$

Note that the factor modifying the de Sitter term turns out to be identical to the factor appearing in the perihelion precession angle<sup>1</sup> in BD theory (this is not true in general, of course). For a value of  $\omega = 6$  we obtain

$$\vec{\Omega}_{\text{BD}} = \vec{\Omega}_{\text{T}} + (11/12)\vec{\Omega}_{\text{DS}} + \frac{15}{16}\vec{\Omega}_{\text{LT}}. \quad (6)$$

It is clear that the most sensitive test of the BD theory occurs when  $\Omega_{\text{DS}}$  and  $\Omega_{\text{LT}}$  attain their maximum possible values. Now, for a gyroscope in a satellite at moderate altitude (orbiting the earth's equatorial plane and with the gyroscope spin axis normal to the earth's axis),  $\Omega_{\text{DS}}$  is about<sup>11,14</sup> 7"/yr,  $\Omega_{\text{LT}}$  is about 0.1"/yr, and  $\Omega_{\text{T}}$  is practically zero ( $\Omega_{\text{T}}$  can be made exactly zero if the slave satellite idea of Pugh<sup>15</sup> is adopted). For a gyroscope in an earth-bound laboratory<sup>11,14</sup> (with spin axis normal to the earth's axis),  $\Omega_{\text{E}}$  is roughly 0.4"/yr with  $\Omega_{\text{T}}$ ,  $\Omega_{\text{DS}}$ , and  $\Omega_{\text{LT}}$  contributing to the same order of magnitude. It is thus clear that the gyroscope in a satellite offers the most sensitive test of the BD theory (particularly of the terms in the BD metric which contribute to the de Sitter effect); it is fortunate that this is also the most convenient experimental arrangement.<sup>7,8,11,14</sup> With regard to experimental accuracy, Schiff<sup>11</sup> states that the direction of the spin axis can be read out with an accuracy of 0.1"; an accuracy of 0.0" now appears possible.<sup>17</sup> Thus we see the favorable possibilities that exist for distinguishing between  $\Omega_{\text{E}}$  and  $\Omega_{\text{BD}}$ . As a further refinement, we note that it will also be possible to test separately the  $(4 + 3\omega)/(6 + 3\omega)$  and  $(3 + 2\omega)/(4 + 2\omega)$  terms appearing in Eq. (5). This arises because

of the different angular dependences. For example,<sup>8</sup> at a laboratory latitude of 35°16' the secular precession arising from the Lense-Thirring effect is zero.

To summarize, we consider that the Schiff gyroscope experiment offers the best possibility for testing the BD theory and that the gyroscope in a satellite offers a more sensitive test than the earth-bound gyroscope.

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$K^+d$  STRUCTURE IN  $I=0$  AT 1.2 GeV/c  
AS A RESULT OF S-STATE  $K-K^*(890)$  CHANNEL COUPLING

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A well-defined peak has been reported<sup>1</sup> in the  $K^+d$  total cross section at 1.2 GeV/c  $K^+$  laboratory momentum. Subtraction of the  $K^+p$  cross section and making the Glauber correction for screening leaves an  $I=0$  peak of approximately 6 mb above background and a width of about 150 MeV.<sup>1</sup>

We wish to report that a previous fit to the data<sup>2</sup> below 800 MeV/c contains a prediction of this experimental peak. The purpose of Ref. 2 was the explanation of the rise in the  $I=0$ , S-state phase shift between 400 and 800 MeV/c. The successful mechanism was the coupling of this  $K^+N$  state to the S-state  $K^*(890)N$  channel whose threshold is at  $\sim 1000$  MeV/c. A simple boundary condition model was used. The boundary radius  $r_0$ , in the theoretically indicated range, was taken from an earlier  $I=1$  fit.<sup>3</sup> The three homogeneous boundary condition parameters were fitted to the scattering length and phase shifts below  $K^*$  threshold. The amount of  $K^*$  production predicted at higher energies was comparable with that observed in the  $I=1$  channel and was quantitatively related assuming isovector exchange.

In Ref. 2 the  $K^*$  width was ignored, as all detailed comparisons were at energies more than 200 MeV/c below  $K^*$  threshold and thus insensitive to its width. We have now included the effect of the  $K^*$  width and recalculated results in the region 0-1.5 GeV/c. The nonvanishing width requires that Eq. (4) of Ref. 2 be replaced by<sup>4-6</sup>

$$f_{\text{eff}}^0 = f^0 - (f_c^0)^2 \int \frac{\rho(m)}{f_*^0 - i r_0 K(m)} dm, \quad (1)$$

where  $K(m)$  is the relativistic momentum of a  $K^*$  of mass  $m$  in the center-of-momentum system;  $f^0$ ,  $f_c^0$ , and  $f_*^0$  are constants; and  $\rho(m)$  is the resonance shape in high-energy  $K^*$  production,

$$\rho(m) = N \frac{\gamma(q/q_*)^3}{(m^2 - m_*^2)^2 + (m_*^4/m^2)\gamma^2(q/q_*)^6} \quad (2)$$

for real pion momentum  $q(m)$  in the  $K^*$  rest system, and vanishes for imaginary  $q(m)$ . The normalization is  $\int \rho(m) dm = 1$ .  $q(m)$  is given by

$$(q^2 + m_K^2)^{1/2} + (q^2 + m_\pi^2)^{1/2} = m. \quad (3)$$

$\gamma$  is the reduced width and  $m_*$  and  $q^*$  are the values of  $m$  and  $q$  at the  $K^*$  peak.

With the above  $f_{\text{eff}}^0$  the complex amplitude is computed as in Ref. 2. Using  $m_* = 891$  MeV and  $\gamma = 50$  MeV the phase-shift fit of Ref. 2 was restored by small variations of the boundary conditions. The choice  $f^0 = 4.1$ ,  $(f_c^0)^2 = 10.2$ , and  $f_*^0 = 1.3$  at  $r_0 = 0.45 m_\pi^{-1}$  fits the phase shift  $\delta_{00}$  below threshold as shown in Fig. 1. The predicted  $K^+N$  scattering length is  $a_0 = 0.036 m_\pi^{-1}$ . The same figure shows the calculated values of  $\delta_{00}$  and  $\eta_{00}$  above threshold. The resulting  $\sigma_{\text{tot}}^0(S_{\frac{1}{2}})$  is peaked as shown in Fig. 2. The experimental  $\sigma_{\text{tot}}^0$  is obtained from  $\sigma_{\text{tot}}^{K^+d}$  and  $\sigma_{\text{tot}}^{K^+p}$  as in Ref. 1 and the error corridor obtained from

$$\Delta \sigma_{\text{tot}}^0 = \{ [2 \Delta \sigma_{\text{tot}}^0(K^+d)]^2 + [3 \Delta \sigma_{\text{tot}}^0(K^+p)]^2 \}^{1/2}, \quad (4)$$

neglecting any error in the calculation of deuteron effects. On subtracting  $\sigma_{\text{tot}}^0(S_{\frac{1}{2}})$  from the experimental cross section the peak is re-