### FRONT-LINE RECURRENT NOVA SCIENCE REQUIRES CENTURY-OLD DATA Bradley E. Schaefer (Louisiana State University)

### Perennial problem: WHAT IS THE PROGENITOR SYSTEM FOR TYPE Ia SUPERNOVAE?





### SUDDENLY A VITAL BIG-MONEY PROBLEM:

- •Must know progenitor to calculate change in SN Ia M- $\Delta t_{15}$  relation
- •Evolution of metallicity in old Universe + change Hubble Diagram shape
- •SNAP cannot achieve goal without progenitor/evolution solution





### **PROPOSED PROGENITORS:**

•Recurrent Novae

- •Symbiotic stars
- •Super-soft sources
- •Double White Dwarf Binaries



### **RECURRENT NOVAE ARE LIKELY SOLUTION:**

To recur with  $\tau_{rec} < 100$  years, RNe must have: •High WD mass  $(1.2M_{\odot} < M_{WD} < M_{Chandra})$ •High accretion rate  $(M \sim 10^{-7} M_{\odot}/yr)$ 

M<sub>WD</sub> will exceed M<sub>Chandra</sub> Ia any year now...

### **TWO PROBLEMS:**

•Does the White Dwarf eject more mass each eruption than it gains between eruptions?  $M_{ejecta} < \tau_{rec} \dot{M}$ ?

•Are there enough RNe to produce the observed Type Ia SN rate?  $R_{RNdeath} = R_{SNIa}$ ?  $R_{RNdeath} = N_{RN} / (M0.2M_{\odot})$ 

# RN DEMOGRAPHICS REQUIRED TO ANSWER THE FRONT-LINE QUESTIONS:

τ<sub>rec</sub> - recurrence time scale
N<sub>RN</sub> - number of RNe in Milky Way
M - mass accretion rate onto white dwarf
M<sub>ejecta</sub> - mass ejected in eruption

# CAN GET THESE ONLY FROM HISTORICAL/ARCHIVAL DATA:

• $\tau_{rec}$  - can only look in archival plate collections

- $\bullet N_{RN}$  archival plates and AAVSO data only way to measure discovery efficiency
- M changes on all time scales, but we need average over the last century
- •M<sub>ejecta</sub> must have pre-eruption eclipse timings









# $\tau_{rec} = Average Recurrence Time Scale:$ Most RN eruptions are *not* discovered: Undirected searches: <efficiency> = 4% (0.6%-19% full range) Directed searches: <efficiency> = 60% (30%-100% full range) Missed eruptions make for errors of ~3X in $\tau_{rec}$ : Old $\tau_{rec}$ =83 years [2000-1917] But I found 1941 eruption Steady brightness 1941-2000 shows eruption ~1975 missed $\Rightarrow \tau_{rec}$ =27 years (3X error) U Sco: Old $\tau_{rec}$ =24 years [6 eruptions since 1863] Dut I found 1017\_1045\_% 1060 eruptions

- But I found 1917, 1945, & 1969 eruptions
  Steady eruptions every 10±2 years (lost behind Sun in 1927 and 1957)
  → τ<sub>rec</sub>=10±1 years (2.4X error)
- •V2487 Oph: Naive  $\tau_{rec}$ =98 years [1998-1900] -But <efficiency>=30% 1890-2008 with two eruptions found  $\rightarrow \tau_{rec}$ =18 years (5.4X error)
- •RS Oph: Old  $\tau_{rec}$ =23 years [5 eruptions since 1890] -But new-found eruptions in 1907 & 1945 plus new eruption in 2006  $\Rightarrow \tau_{rec}$ =14 years (1.6X error)



 $N_{RN} = 10 / (0.2/200) = 10,000$  RNe in our galaxy!

•TEST: Seek old eruptions of 'Classical Novae': Success - Nova Oph 1998 erupted in 1900

(Pagnotta, Schaefer, & Xiao; 2008)











**RESULTS:** [see also my Poster 491.04 on Wednesday]  $\cdot R_{RNdeath} \sim R_{SNIa}$  for Milky Way, M31, & LMC

 $\label{eq:massed_ejecta} \bullet M_{ejecta} \, <\!\! < \tau_{rec} \, \Bar{M} \quad \mbox{ for CI Aql and U Sco}$ 

### **LESSONS**:

Most nova events are missed → Great opportunity for observers
 → ~30% of 'Classical Novae' are RNe with multiple eruptions in the last century

•Harvard plates should be digitized [see Poster 427.07 by Tang, Grindlay, Los, & Laycock]

Front-line science is laying around in archival data
Science unique and unobtainable with modern telescopes [cf. Special Session on Monday]