

Movies

Produced by LSU's Astrophysics Theory Group

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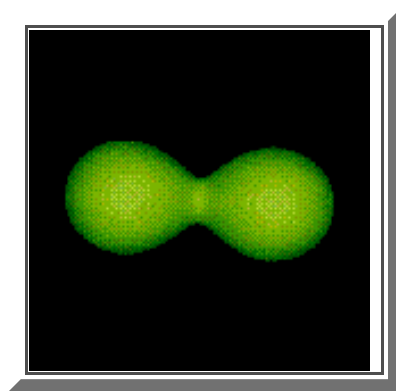


NOTE: One frame from the "Binary Merger" movie shown below has been featured on the [cover of the 1998 July/August issue of *Computers in Physics*](#) (now *Computing in Science and Engineering*) and a sequence of images from the movie have been published in that issue's [Internet Goldmine](#) article.

The Relative Stability of Close, Equal-Mass Binaries [1995 - 1997]

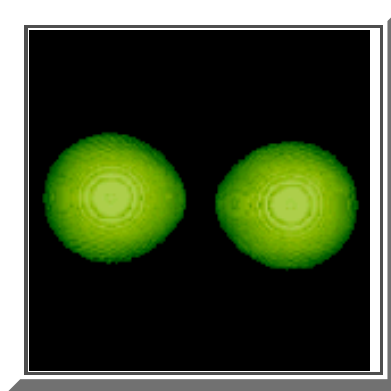
We have published the results of an extensive study of the stability of close, equal-mass binary star systems that exhibit different degrees of gas compressibility ([New and Tohline 1997](#); see also, [New 1996](#)). Two animation sequences illustrating the results of this investigation are shown here. The animation referred to as "Stable Binary" follows a binary system through approximately four full orbits in a frame of reference that is rotating with the initial orbital frequency. It illustrates clearly that even a common-envelope binary system can be dynamically stable against merger. The other animation -- labeled, "Binary Merger" -- illustrates how an evolution proceeds in a system with a stiff equation of state that is unstable toward merger. A principal conclusion of the [New and Tohline 1997](#) investigation is that close binary stars with "stiff" (almost incompressible) equations of state are unstable to merger whereas binary stars with soft equations of state -- such as would arise in binary protostellar systems -- are stable against merger.

Stable Binary



[Quicktime](#) (1,380K)

Binary Merger



[mpeg](#) (701K)

References

- New, K.C.B. (1996), *Instabilities in and Gravitational Radiation from Compact Stars and Compact*

- New, K.B.C. and Tohline, J.E. (1997), *The Astrophysical Journal*, **490**, p. 311

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